# Observations of OH, $HO_2$ , $H_2O$ , and $O_3$ in the upper stratosphere: implications for $HO_x$ photochemistry

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Abstract. Balloon-borne observations of concentrations of OH, HO<sub>2</sub>, H<sub>2</sub>O, and O<sub>3</sub> in the middle and upper stratosphere are used to test our understanding of HO<sub>x</sub> photochemistry. Assuming our photochemical model is complete, the measured [OH] and [HO<sub>2</sub>] above 38 km (where HO<sub>x</sub> partitioning is no longer dependent on [NO]) are modeled best by calculations that use a 25% reduction in the ratio of the reaction rate constants for O+HO<sub>2</sub> and O+OH as well as either a 25% reduction of the rate constant of OH+HO<sub>2</sub> (the primary HO<sub>x</sub> sink) or a 25% increase in HO<sub>x</sub> production. All of these changes are consistent with the uncertainties in the recommended rate constants. The kinetic parameters required to explain our observations of [OH] and [HO<sub>2</sub>] do not lead to a resolution of the long-standing "ozone deficit problem" above 45 km.

# Introduction

 $\mathrm{HO}_x$  (defined here as  $\mathrm{OH} + \mathrm{HO}_2$ ) is produced in the upper stratosphere and mesosphere principally via

$$H_2O + O(^1D) \rightarrow OH + OH$$
 (1)

and by photolysis of  $H_2O$  above 60 km. The primary  $HO_x$  loss reaction is

$$HO_2 + OH \rightarrow O_2 + H_2O, \tag{2}$$

although the reaction of OH with HNO<sub>3</sub> becomes an important  $HO_x$  sink in the lower and middle stratosphere. The partitioning between OH and  $HO_2$  is driven mostly by the reactions

$$HO_2 + O \rightarrow OH + O_2,$$
 (3)

$$OH + O \rightarrow O_2 + H, \tag{4}$$

$$HO_2 + NO \rightarrow OH + NO_2,$$
 (5)

and

$$OH + O_3 \rightarrow HO_2 + O_2. \tag{6}$$

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Reactions 3 and 4 dominate the partitioning of  $HO_x$  above approximately 38 km, while reactions 5 and 6 (along with  $HO_2+O_3$ ) dominate at lower altitudes. The H atom produced in reaction 4 is quickly converted to  $HO_2$  by reaction with  $O_2$ , which is about 10 times faster than its reaction with  $O_3$  that recreates OH at these altitudes.

Recent measurements of [OH] and [HO<sub>2</sub>] individually have suggested an incomplete understanding of  $HO_x$  chemistry in the middle atmosphere. Summers et al. [1997] reported measurements of [OH] between 50 and 80 km that were  $\sim 25-30\%$  lower than expected based on standard photochemical theory. Sandor and Clancy [1998] found that ground-based microwave measurements of [HO<sub>2</sub>] were 25-30% higher than expected. Since neither of these studies obtained simultaneous measurements of [OH] and [HO<sub>2</sub>], it is not possible to discern whether the reported discrepancies were due to errors in the partitioning of  $HO_x$ , in the balance of production and loss of  $HO_x$ , or a combination of both factors.

The Far Infrared Spectrometer (FIRS-2) simultaneous measurements of concentrations of OH, HO<sub>2</sub>, H<sub>2</sub>O, and O<sub>3</sub> presented here provide a stringent test of our understanding of the photochemistry of  $HO_x$ . We show: a) measured [OH] agrees reasonably well with, and measured [HO<sub>2</sub>] is significantly higher than, values calculated using standard kinetic parameters for altitudes between 40 and 50 km; b) modifications to both the partitioning of HO<sub>x</sub> and the balance of production and loss of  $HO_x$  are necessary to explain the observed [HO<sub>2</sub>]/[OH] ratio and abundance of HO<sub>2</sub>; c) the changes to reaction rates proposed by Summers et al. [1997] and Sandor and Clancy [1998] to explain their observations are inconsistent with the simultaneous FIRS-2 observations of [OH] and [HO<sub>2</sub>]. Furthermore, the FIRS-2 observations are unlikely to help resolve the "ozone deficit problem" above 45 km. Models that describe well our observed profiles of [OH] and [HO2] either have similar, or enhanced, imbalances between production and loss of  $O_3$ compared to the standard calculation.

### Measurements and Model

The FIRS-2 is a thermal emission far-infrared Fourier transform spectrometer developed at the Smithsonian Astrophysical Observatory and operates mostly from balloons [Johnson et al., 1995]. The spectrometer has recently been upgraded with a new beamsplitter [Dobrowolski and Traub, 1996] to obtain nearly complete spectral coverage between 75 and 1000 cm<sup>-1</sup> with high interferometric efficiency. Cur-

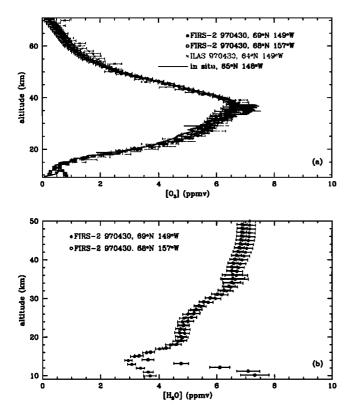


Figure 1. (a) Balloon-borne mixing ratios of O<sub>3</sub> retrieved by FIRS-2 and an in-situ photometer as well as ILAS during a nearly coincident occultation on 30 April 1997 in northern Alaska. (b) Mixing ratios of H<sub>2</sub>O retrieved from the same FIRS-2 data set. FIRS-2 data points above 38 km are derived from column measurements constrained by climatological values. Measured O<sub>3</sub> and H<sub>2</sub>O are used to constrain the photochemical model calculations.

rently, FIRS-2 spectra yield vertical concentration profiles of over 28 molecules. Measured slant columns are fit in an onion-peeling fashion and a singular-value decomposition routine is then used to retrieve mixing ratios on a 1 km vertical grid [Johnson et al., 1996]. The concentrations above the balloon altitude are determined from the observed overhead columns retrieved from the FIRS-2 spectra and constrained either by the profile shape calculated using the photochemical model described below or climatological profiles.

The measurements presented here were obtained during a balloon flight from Fairbanks, Alaska on 30 April 1997 as part of the ADvanced Earth Observing Satellite (ADEOS) Alaska validation campaign. We produced two sets of profiles: one from spectra obtained while observing to the north, at  $9:15\pm1$  hr local solar time, centered at  $69^{\circ}$ N,  $149^{\circ}$ W and a second from spectra obtained while observing to the northwest, at  $11:15\pm1.5$  hr local solar time, centered at  $68^{\circ}$ N,  $157^{\circ}$ W. The balloon float altitude was 37 km.

[OH] is retrieved from 11 rotational transitions between 83 and 230 cm<sup>-1</sup> and [HO<sub>2</sub>] is retrieved from 43 rotational transitions between 110 and 220 cm<sup>-1</sup>. The systematic error due to uncertainties in the spectroscopic parameters for OH and HO<sub>2</sub> rotational transitions is small. Our estimate of the total systematic uncertainty for the retrievals presented here is 5% for [OH] and 3% for [HO<sub>2</sub>].

Investigating  $HO_x$  photochemistry in the middle and upper stratosphere requires accurate measurements of  $[O_3]$ 

(used to derive concentrations of O and O(1D) as well as the photolytic environment) and [H<sub>2</sub>O]. The FIRS-2 profiles of [O<sub>3</sub>] are retrieved using rotational transitions between 80 and 130 cm<sup>-1</sup> and ro-vibrational transitions between 750 and 810 cm<sup>-1</sup>. The observed concentration profile of [O<sub>3</sub>] is shown in Figure 1a, along with a profile obtained by an in situ UV photometer on the same gondola and a nearby [O<sub>3</sub>] profile measured by the Improved Limb Atmospheric Sounder (ILAS) instrument (version 3.0 data) on board ADEOS. The good agreement between the FIRS-2, in-situ, and ILAS profiles (as yet unvalidated) below 38 km suggests ILAS has accurately measured [O<sub>3</sub>] in the middle stratosphere. Above 38 km, the overhead O<sub>3</sub> column retrieved by FIRS-2 is in good agreement with the column above 38 km determined from the ILAS profile. Thus, we are confident that the ILAS [O<sub>3</sub>] measurement is accurate (systematic uncertainty of about 10% depending on the inversion algorithm used) and confirms the climatological profile shape used to constrain the FIRS-2 profile above 38 km.

The FIRS-2 measurements of  $[H_2O]$  are plotted in Figure 1b.  $[H_2O]$  is retrieved using rotational transitions between 140 and 500 cm<sup>-1</sup>, selected to reduce the systematic uncertainties to 5%. The profile shape above 38 km is constrained by climatology. The FIRS-2 concentrations of  $[H_2O]$  above 35 km are ~8% higher than those of HALOE for similar latitudes and time of year. Use of HALOE  $[H_2O]$  would decrease the modeled  $HO_x$  discussed below by 4% and increase the difference between measured and calculated values.

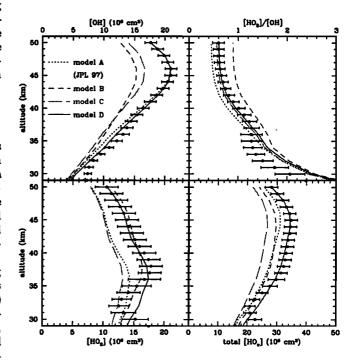


Figure 2. Measured and modeled concentrations of OH, HO<sub>2</sub>, and HO<sub>x</sub> as well as [HO<sub>2</sub>]/[OH] from the FIRS-2 balloon flight on 30 April 1997 for the first set of observations at 69°N, 149°W 9:15 local solar time. All model curves are constrained by simultaneous FIRS-2 measurements of temperature, O<sub>3</sub> and H<sub>2</sub>O. Model A uses standard photochemical reaction rates and cross sections, model B has the rate constant for O+HO<sub>2</sub> decreased by 50%, model C has O+OH decreased by 20% and OH+HO<sub>2</sub> increased by 30%, and model D has the rate constants for O+HO<sub>2</sub> and OH+HO<sub>2</sub> both reduced by 25%.

The photochemical model employed here has been used previously to interpret data from aircraft, satellite, and balloon platforms [Jucks et al., 1996 and references therein]. In the present analysis, the model is constrained by FIRS-2 measurements of  $[O_3]$ ,  $[H_2O]$ ,  $[Cl_y]$ ,  $[NO_y]$ ,  $[N_2O]$  (to infer  $[CH_4]$ ), and temperature in the middle and upper stratosphere.

## **Discussion**

The measured concentration profiles of OH,  $\rm HO_2$ , and total  $\rm HO_x$  as well as  $\rm [HO_2]/[OH]$  between 30 and 50 km are shown in Figure 2. Only the first of the two sets of data is displayed here for simplicity, but both sets show quantitatively consistent comparisons with the model calculations discussed below. Above 35 km, measured concentrations of OH are in reasonable agreement with the values calculated using recommended reaction rates and absorption cross sections from DeMore et al. [1997] (model A) while measured concentrations of  $\rm HO_2$  are  $\sim 25\%$  higher than the calculated values from model A. The discrepancy for  $\rm [HO_2]$  is similar to that reported by Sandor and Clancy, [1998] for altitudes above 50 km. The good agreement for [OH] is similar to that reported near 40 km by Pickett and Peterson [1996].

Summers et al. [1997] reported lower concentrations of OH between 50 and 80 km than calculated using recommended reaction rates and cross sections. They showed agreement with the MAHRSI observations of [OH] would be improved if either the rate constant of reaction 3 (k<sub>3</sub>) were reduced by 50% (model B in Figure 2) or k<sub>2</sub> was increased by 20% and k<sub>3</sub> was reduced by 30% (model C in Figure 2). The latter changes are within the uncertainties given by DeMore et al. [1997]. Our models B and C use the same kinetic parameters as the similarly named models of Summers et al. [1997]. Model B is also similar to the kinetic change favored by Sandor and Clancy [1998] (they used a 40% reduction in k<sub>3</sub>) to explain their measurements of [HO<sub>2</sub>].

Neither model B or C lead to calculated profiles for both [OH] and [HO<sub>2</sub>] that are consistent with the FIRS-2 observations. Both result in [OH] significantly lower than measured. The model C calculation also results in significantly lower [HO<sub>2</sub>] than observed. The difference between the observed and modeled columns of OH and HO<sub>2</sub> between 38 and 50 km (using both sets of [HO<sub>x</sub>] retrievals described above) has a reduced  $\chi^2$  of 5.3, 7.8, and 7.9 for models A, B, and C, respectively.

The contrasting conclusions regarding our understanding of  $\mathrm{HO}_x$  photochemistry drawn from the FIRS-2 and MAHRSI observations appear to result from a fundamental measurement difference, since our model reproduces the Summers et al. [1997] calculations. The two measurements of [OH] are difficult to compare directly since 50 km is both the lowest altitude point reported by MAHRSI and the highest altitude point by FIRS-2. The differences between the concentrations of [OH] measured by MAHRSI and the model A calculations are close to the systematic uncertainties of 23% quoted in Summers et al. [1997].

The FIRS-2 observations of the  $[HO_2]/[OH]$  ratio imply the partitioning of  $HO_x$  is not calculated correctly using the kinetic parameters of *DeMore et al.* [1997]. This suggests either a modification to the ratio of  $k_3/k_4$  is necessary (the observations cannot be used to discriminate between the two rates) or else there exists an undocumented reaction that converts OH to  $HO_2$  that would need to proceed at an approximate 24 hour average rate of  $2.8 \times 10^5$  molec/cm<sup>3</sup>/sec at 48 km. Such a reaction is unlikely given the extensive labo-

ratory studies of OH reactions over the past two decades. The measured  $[HO_2]/[OH]$  ratio is consistent with  $k_3/k_4$  equal to 1.34 at 250 K. Keyser [1983] reported a laboratory measurement at room temperature for  $k_3/k_4$  of 1.7  $\pm$  0.2, with an overall experimental error of  $\pm 25\%$ . While the ratio of the rate constants for reactions (3) and (4) necessary to explain the FIRS-2 observations is systematically lower than the existing laboratory data, we note that the experiment of Keyser [1983] was conducted only at room temperature and that the ratio for  $k_3/k_4$  required to account for the FIRS-2 observations is within the overall experimental error of the laboratory measurement.

The FIRS-2 observations of the abundance of OH and  $HO_2$  imply also that the balance between production and loss of  $HO_x$  is not calculated correctly by the standard model. The reaction of OH and  $HO_2$  [reaction (2)] is the dominant sink for  $HO_x$  in the upper stratosphere. Since the measured product of  $[OH] \times [HO_2]$  is larger than the product calculated using model A, either the production rate of  $HO_x$  must be increased or the rate constant of the loss reaction decreased to lead to better agreement with the measurements.

There are considerable uncertainties in the production rate of  $\mathrm{HO}_x$ , which depends on the photolysis rate of  $\mathrm{O_3}$ , the quantum yield of  $\mathrm{O(^1D)}$ , and  $\mathrm{k_1}$ . The existing laboratory measurements of  $\mathrm{k_2}$  show wide variation, reflected in the large uncertainty (52% at 240 K) associated with the recommended rate [DeMore et al., 1997]. Equivalent values for calculated [OH] and [HO\_2] result if, instead of decreasing  $\mathrm{k_2}$ , a comparable increase were used for  $\mathrm{k_1}$ . Finally, we note that there is little coupling between the partitioning reactions and the production/loss reactions at these altitudes because the partitioning reactions occur at rates about 50 times the rate of the production/loss reactions. Consequently, changes alone in the rates of the production/loss reactions have little effect on the calculated [HO\_2]/[OH] ratio.

Good agreement is obtained with the FIRS-2 observations of [OH] and [HO<sub>2</sub>] above 35 km when the rate constants of reactions 2 and 3 are both reduced by 25% (model D in Figure 2). We emphasize that comparable results are obtained for ~25% increases in the rates of reactions 1 and 4. The difference between the observations and model D profiles for [OH] and [HO<sub>2</sub>] has a reduced  $\chi^2$  of 1.2; a 5% change in either rate alteration doubles the reduced  $\chi^2$ . As with model C, the values of  $k_2$  and  $k_3$  adopted in model D are consistent with the uncertainties given by *DeMore et al.* [1997].

Below 35 km, reaction 5 starts to dominate  $HO_x$  partitioning. This is evident at the lower altitudes of the data presented in Figure 2, where the  $[HO_2]/[OH]$  ratio is small compared to the model, implying higher [NO] than calculated. Our simultaneous measurements of  $NO_2$  and  $HNO_3$  suggest that the model underestimates  $NO_x$  in a manner consistent with the imbalance in observed  $[HO_2]/[OH]$  (Jucks et al., manuscript in preparation).

#### Implications for Ozone

The required changes in rate constants needed to account for our observations of  $HO_x$  will have either a negligible or worsening effect on the "ozone deficit" problem above 45 km [Osterman et al., 1997 and references therein]. Good agreement with the FIRS-2 measurements is obtained if the rate constant for  $O+HO_2$  (k<sub>3</sub>), the dominant rate limiting step for loss of odd oxygen above 45 km, is reduced by 25% together with either an increase in k<sub>1</sub> or a decrease in k<sub>2</sub>

(model D). Since observed [HO<sub>2</sub>] is roughly 25% higher than values calculated using standard kinetics (model A), the calculated imbalance between production and loss of odd oxygen found using model D (loss is 30% greater than production at 48 km) is nearly identical to that found using model A. If the rate of O+OH (k<sub>4</sub>) is instead increased by 25% a larger imbalance between production and loss of odd oxygen results (loss ~50% greater than production at 48 km) because the higher than expected concentrations of [HO<sub>2</sub>] are no longer offset by a change in the rate of the limiting step. Sandor and Clancy [1998] proposed that reducing k3 by 40% would resolve the ozone deficit above 45 km. However, as noted above, this change alone results in discrepancies between theory and FIRS-2 observations of [OH], and consequently we do not consider it to be a viable solution to the ozone deficit problem.

It appears that the solution to the "ozone deficit" problem at these altitudes lies either with the calculated production rate of odd oxygen or with calculated concentrations of O. There are no observational constraints on the calculated concentrations of O at the required quantitative level. Recent studies have proposed scenarios involving vibrationally excited  $O_2$  that could lead to increased production of odd oxygen, but critical aspects of this process involving the quenching of  $O_2$  (v>26) await laboratory confirmation [Toumi et al., 1996]. This scenario would have the largest effect on calculated production rates near 48 km, and have a significantly lesser effect near 35 km, where calculated production and loss are found to balance.

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