

Part I Seamless Astronomy

Alyssa A. Goodman Harvard-Smithsonian Center for Astrophysics

with Alberto Accomazzi, Rahul Davé, Gus Muench & Michael Kurtz (Harvard-Smithsonian CfA); Tim Clark (Massachusetts General Hospital/Harvard Medical School); Jonathan Fay & Curtis Wong (Microsoft Research)

+extended & upcoming collaboration with Chris Borgman & Alberto Pepe* (UCLA);
 Doug Burke; Sarah Block, Pepi Fabbiano, et al. (CfA); E. Bressert (U. Exeter);
 J. Hendler & D. McGuinness (RPI); A. Conti & C. Christian (STScI); A. Connolly et al. (U. Washington)

Realm of Seamless Astronomy



3500 years of Observing

Stonehenge, 1500 BC



Ptolemy in Alexandria, 100 AD



Observatory Tower, Lincolnshire, UK, c. 1300



Galileo, 1600



The "Scientific Revolution"

Reber's Radio Telescope, 1937





NASA/Explorer 7 (Space-based Observing) 1959

"The Internet"



NVS

Long-distance remote-control/ "robotic" telescopes I 990s



"Virtual" observing

000	0										COM	PLETE Data C	overage Tool								
4 1-	36	\$	2	A	A	+	O http://	/www.worldwid	fetelescope.o	g/COMPLE	TE/WWTCoverageToolJ	Norw#					6 Qr (Google) 🕐
Ⅲ		ie w K	odak	EXP	LO I	ling	WHITSL	Alyssa Good.	Home Page	Toodleda	Harvard BC: Projects	Wikis Etc.v	Google Calendar	\$55	Image Search +	fbk share	Directories *	ADS Best	RSS (3387) +	BeyondAD5 +	39

0

•

4

1

1

×

8

8

3

ø

0

1

Data

Data

Data

Data

Data

Data

Data

Data

D th

Data

Data

Data

Data

0.15



What can today's Astronomer's "Research" look like?

Research

In my Astronomy research, I am primarily interested in how the gas in galaxies constantly re-arranges itself over huge time spans to constantly form new stars. I have also had a long-standing interest in data visualization, and in improving the use of computers in all aspects of scientific research. I teach a course at Harvard called "The Art of Numbers," and I am very involved in the WorldWide Telescope Project, which brings astronomical data to everyone through an interface that demonstrates data delivery for the 21st Century of "e-Science."



http://www.cfa.harvard.edu/~agoodman/

Publishing Simulation Data WorldWide Telescope Ambassadors Program http://www.cfa.harvard.edu/WWTAmbassadors. Harvard University, WGBH & Microsoft Research 3D WorldWide Telescope Alyssa Goodman, Patricia Udomprasert, Annie Valva & Curtis Wong In WorldWide Telescope and Its Ambassadors Pro-Wold Telescope (WWT) is a finitatic "Universe I I primarily by Carlos Wang and Ionathum Riy at Net is as a Virtual Astronomical Observatory Inineigi Iss, atore of online data and Information about our Univer in a kay related hood within the online autonomy ; presently as the "Wol" (see A. Geotman's "Semice circle), but I also offens apprecienting new op-curity). each. WorldWide Telescope Ambassadors Program promotes WWT as ire-leaning way to teach and learn STEM concepts by recruit/ onomically-literate volunteers who are trained to be experts in using WWT Our current Collaboration brings together preference activity and science educations at herein, comparational vituations at MS. Research, and Stere education and outraceh apecalistica at WCBH. The sub-place has a table below) will include participants from selected areas within the US, including Wissimption, Torida, Arcana, Alaska, and Appatchia. e-Science Tools Viz Who are the WWT Ambassadors, and what do they do? WW Ambassions are carefully exclude for barring from amongst: 1) retried STEM professionals and amsteur astronomers with deep knowledge of astronomy and physics; 2) undergoduste and grobuse students and postocicral fellows in Astronomy and i source trachers. In their training, Amtessades team how to use WW's tools in general, and also how to create and pusies astrophysical concepts. These Tours allow users to display beaufild astronomical images in their proper contacts in the demonstrating the physical principies at work in those images. Ambasadero care create and use materials within WYT. WWT Tour WCEH What have we done so far? Our program begun in the fail of 2009. Initial Ambassadors are currently working with 80 middle statout students and their bascher, Michelle Bacilloy, at the Clarke Meddle School in provider have the students and their bascher, Michelle Bacilloy, at the Clarke Meddle School in provider have the students and their bascher backware generated the topological the fram the students, and have inspired outly learning through exploration and discovery fram the students, and have inspired outly learning through exploration and discovery site coles to all students, and an analysis of the Mick experience will serve to inform the NSF provided beng such the original in the Spring 4708. What's the whole plan, and what are the program's goals? times to be write previous into write the up regress system. We are presently reparking a proposite to the National Schnice Foundation, hased in large part on our "Nict" experiences, to implement "Phase II" of the Ambassadors Project (see table), where we will be plan a limited expansion within in School Schnick (School School A relical goal of this project to be needs a **full actoromy curriculum using WWT Tours enabled by our Ambassadon**. These Tours will be wheth by the sector the declaration professionalis within tour ballboardon, and there will be integrated with **WGBH Teachers' Demain** which currently has needy 400,000 mightend users. NorldWide Telescope can help change how students learn science by demonstrating the joys of inquiry and discovery, and the WWT Program is designed to help to increase science literacy in the general public while forming intergenerational connections within their co Fall 2009-Spring 2010 Research 😈 🚧 🗰 Fall 2010-Summer 2011 US-wide Fall 2011-Summer 2012 2012+

External Research

Research

WorldWide Telescope: a UIS from Microsoft Research [UIS=Universe Information System]



Seamless Data/Literature Connections (e.g. ADS) "Modular Craftsmanship" (e.g. flickr) Collections, Communities & Guided Tours



The World Wide Telescope an Archetype for Online-Science

Jim Gray (Microsoft)

Alex Szalay (Johns Hopkins University)

Microsoft Academic Days in Silicon Valley

http://research.microsoft.com/~gray/talks



Created by Curtis Wong and Jonathan Fay at MSR; AG is "Academic Partner" on the WWT Project

The (US) Backstory

R

2.4

UNDING | AWARDS

Science New

\$10 Million N

ScienceDaily (Oc its users the world research institutic starting an ambiti universe online.



200

See Also:

(NVO), headed by astronomer Alex

NVO senior personnel:

Charles Alcock, University of Pennsylvania Kirk Borne, Astro Tim Cornwell, NSF National Radio Astronomy Observatory I Optical Astronomy Observatory Giuseppina Fabbiano, Smit Observatory Alyssa Goodman, Harvard University Jim Gray Hanisch, Space Telescope Science Institute George Helou, N Analysis Center Stephen Kent, Fermilab Carl Kesselman, Uni Miron Livny, University of Wisconsin, Madison Carol Lonsdo and Analysis Center Tom McGlynn, GSFC/HEASARC/USRA A University Reagan Moore, San Diego Supercomputer Cente Naval Observatory, Flagstaff Station Ray Plante, University Thomas Prince, California Institute of Technology Ethan Sch STScI Nicholas White, NASA Goddard Space Flight Center Ro of Technology

Find Funding

- A-Z Index of Funding **Opportunities**
- Recent Funding Opportunities
- Upcoming Due Dates
- Advanced Funding Search
- **How to Prepare Your Proposal**
- About Funding

Proposals and Awards

Proposal and Award Policies and Procedures Guide Introduction **Proposal Preparation and**



SEARCH NSF Web Site ICS | ABOUT Print,

Management and Operation of the Virtual Astronomical Observatory

CONTACTS

Name	Email
Nigel Sharp	nsharp@nsf.gov
Eileen D. Friel	<u>efriel@nsf.qov</u>

PROGRAM GUIDELINES

Solicitation 08-537

Please be advised that the NSF Proposal & Award Policies & Procedure (PAPPG) includes revised guidelines to implement the mentoring pro the America COMPETES Act (ACA) (Pub. L. No. 110-69, Aug. 9, 2007.) specified in the ACA, each proposal that requests funding to support postdoctoral researchers must include a description of the mentoring that will be provided for such individuals. Proposals that do not com this requirement will be returned without review (see the PAPP Guide Grant Proposal Guide Chapter II for further information about the implementation of this new requirement)



and meanwhile...





Seamless Astronomy



But, that was 2009...





This simple argument, first made at the 2009 WWT session at AAS, seems to be working:

"Astronomy research tools should work as seamlessly as travel research tools."

"Astronomy research tools should work as seamlessly as travel research tools."

When the concept of a "Virtual Observatory" (VO) was first discussed by future-looking astronomers in the mid-1990s, all thoughts were about **distributed data** and a **common system** to access it. But, information access on today's web primarily works in the **reverse**: distributed tools accessing **common data centers**. Capability and ease-of-use improvements to the web typically now come in the form of **nesting, aggregating or connecting tools.** Think **kayak.com**, iGoogle, or Bing Maps. In the "Seamless Astronomy" view to be discussed, today's "VO" should be thought of as the ever-improving set of data archives, tools, interconnections, and standards that strive to make astronomical research as "seamless" as travel research. The good news is that the cutting-edge of the astronomical research environment is moving rapidly in this seamless direction. The most savvy institutions are beginning to realize that the original VO model of data distributed on thousands of individual researchers' desktop hard drives is not a sustainable model, and that they need to offer data hosting, archiving, and stewardship services the way libraries offer such services for printed matter. Software tools are becoming much more interoperable thanks to protocols for messagepassing such as "SAMP." And, the improved speed of web applications is to some extent removing platform-dependence as an obstacle to programmers and users alike. The bad news is that most astronomers are largely unaware of the tools that this new nirvana offers, and instead still conduct online research in the same way they did a decade ago. In this talk, I will focus in particular on how our recent work on connecting Microsoft's WorldWide Telescope program to other commonly-used astronomical research tools--most notably literature searching tools--has made the astronomical research environment more seamless. More generally, I will emphasize and demonstrate that an ever-increasing diversity of tools allow researchers to carry out a particular research task, so that the important research for the future lies in figuring out how to make the tools, their interconnections, and their connections to data and literature resources useful and well-known to the astronomical community.

(Abstract of the Evermore Seamless Astronomy presentation by A. Goodman at the Microsoft External Research Symposium, April 2010.)

From: Abstract Service <ads@cfa.harvard.edu>

Subject: myADS Notification (Astronomy database)

Date: March 23, 2010 12:19:23 AM EDT

To: Alyssa Goodman



myADS Personal Notification Service for Alyssa Goodman Tue Mar 23 00:19:23 2010 Astronomy database

No new articles found

Favorite Authors - Recent Papers

ADS Main Queries	GOODMAN, ALYSSA -	
Astronomy	Citations: 3310 (total 4002)	
Physics	2010NewA15444K: Karatas,+:	New
arXiv e-prints	photometry	/-beta
FAQ	2010MNRAS.403.1054D: Dabring	hause
What's new	Mass loss and expansion of ultra	oomna
	dwarf galaxies through gas expu	F
	stellar evolution for top-heavy ste	Sub
Current Tables of	mass functions	1
Contents	2010ApJ713269F: Federrath,-	
Astronomical Journal	Collapse and Accretion in Turbul	Repl
Astronomy &	Clouds: Implementation and Con	
Astrophysics	Sink Particles in AMR and SPH	
<u>/////////////////////////////////////</u>	2010ApJ712.1403P: Pech,+: C	
Astronomy &	a Recent Bipolar Election in the V	14
Astrophysics	Hierarchical Multiple System IBA	
Supplements	2422	_
Astrophysical Journal	2010ApJ712.1137K: Kauffmani	1

Astronomers can see parallels...

PROPER MOTION, etc - Recent Papers 2010A&A...511A..90B: Breddels,+: Distance en.+: determination for BAVE stars using stellar From: Kayak Alert <alert@kayak.com> bject: Your KAYAK Fare Alert: Boston (BOS) > Munich (MUC) Date: March 26, 2010 3:52:30 AM EDT To: Alyssa Goodman y-To: Kayak Alert <alert@kayak.com>



Literature Handling: Diverse Apps, Common Data



What fraction of astronomy researchers know about these tools?



zotero

zotero

Leveraging the long tail of scholarship.

= Use of Zotero 2.0 with your Word and OpenOffice d

"writemypaper.org?"

ARXIVSORTER

logged in as agoodman

Arxivsorter uses the network of co-authorship (based on papers on arxiv.org since 1992) to estimate proximities between authors and consequently papers.

Please enter a list of authors whose publications are particularly relevant for you. They will define a reference region in the network of co-authorship. Experience shows that, on average, satisfactory results are obtained by entering about five names. However, depending on the range of interests, a longer list might be needed (it is usually a good idea to include yourself).

Enter a last name: alves J. Alves(ALVES_J)	Retrieve Add		Favo	Arxiv Remove Mark R. Kru Arxiv Remove	HEILI Imholz KRUMHOL	ES_C
XIVSORTER	New papers	Recent papers	Select a month	Arxiv Remove	MCK	EE_C Info
4x10 ⁻² arXiv:1003.4900 [pdf] Star-forming gas in young Philip C. Myers Comments: To appear in Astrophysical Journ Subjects: Galaxy Astrophysics (astro-ph.GA)	g clusters nal, May 2010				39	paper
Initial conditions for star formation in o to 10 solar masses. Star-forming infall feedback. For spherical infall, the typic shallower gradient, as in the clumps ar mean infall stopping time is 0.05 Myr a to 0.3 Myr, and the mass accretion rate ~5 solar luminosities, reducing the lum shallow at radius 0.04 pc, enclosing 0.9	clusters are estimated for p is assumed equally likely is al initial condensation mus round cores. These properti and the accretion efficiency e increases from 3 to 300 x hinosity problem to a factor 9 solar masses, with mean	to stop at any mom to stop at any mom t have a steep dens es match observed is 0.5. The infall d (10^(-6) solar mas ~3. The initial cor column density 2 x	asses follow the init ient, due to gas disp sity gradient, as in lo column densities in uration increases wit sses/yr. The typical idensation density gr 10^(22) cm^(-2),	ial mass function persal dominated ow-mass cores, cluster-forming th final protosta spherical accret radient changes and with effective	n (IMF) from l by stellar surrounded regions wh r mass, from ion luminos from steep ve central	by a ben the m 0.01 ity is to

esults are

"writemypaper.org?"

Selected and retrieved 200 abstracts.

SAO/NASA Astrophysics Data System (ADS)

Query Results from the ADS Database

Related Objects
NAME LMC (26)
NGC 292 (15)
SN 1987A (13)
M 31 (9)
NGC 7293 (6)
NGC 6888 (6)
NGC 6543 (6)
<u>M 33 (6)</u>
<u>HIP 54283 (6)</u>
HIP 33165 (6)
<u>VV 344a (5)</u>
V* eta Car (5)
<u>V* CW Leo (5)</u>
NGC 7027 (5)
SNR G111.7-02.1 (4)
NGC 6826 (4)
NGC 2438 (4)
NAME BUTTERFLY NEBULA (4)
MCG+12-08-033 (4)
<u>GSC 06253-02182 (4)</u>
<u>WR 147 (3)</u>
<u>V* V1302 Aql (3)</u>
<u>V* V1042 Cyg (3)</u>
SNR J052501-693842 (3)
PN G208.5+33.2 (3)
NOVA Aql 1919 (3)
<u>NGC 7009 (3)</u>
NGC 6537 (3)
<u>NGC 3132 (3)</u>
<u>NGC 2440 (3)</u>
<u>NGC 2359 (3)</u>
<u>NGC 891 (3)</u>
NAME MAGELLANIC CLOUDS (3)
NAME LOCAL GROUP (3)
NAME HOMUNCULUS NEBULA
NAME FROSTY LEONIS NEBULA



ADS Faceted Topic Search (alpha)

Enter one or more keywords on your subject of interest, sit back and relax.

1	#	Bibcode Authors	Score Title	Date	List Acc	of L ess (winds and e.g.: <u>"dark energy"</u> , <u>"ex</u>	shells xtrasola	from st ar plane	tars ets", <u>"we</u>	eak lensing" "spin hall"	Search
	1	 <u>1995RvMP67661B</u> Bisnovatyi-Kogan, G. S.; Silich, S. A. 	19.000 Shock-w	Jul 1995 ave propagatio	<u>A</u> on in	E the	Keyword Search: Most relevant Most recent Most important	nt			Subject Area Search: Most popular Most useful Most instructive	
1]]	2	<u>1999NewAR4331F</u> Frank, A.	18.000 Bipolar	May 1999 outflows and th	A ne ev	E olution	of stars	<u>ADS</u>	Home	<u>e Abs</u>	stract Search Help	
<u>02.1 (4)</u>	3	2007ARA&A45177C Crowther, Paul A.	13.000 Physical	Sep 2007 Properties of V	A Wolf-	<u>E</u> Rayet S	<u>X</u> Stars	<u>R</u> (<u>C</u>	<u>s</u>	<u>U</u>	
<u>2182 (4)</u>	4	2002ARA&A40439B Balick, Bruce; Frank, Adam	13.000 Shapes a	n/a 2002 and Shaping of	A Plane	E E etary N	lebulae	<u>R</u> (<u>C</u>	<u>s</u>	<u>U</u>	
<u>(3)</u> - <u>693842 (3)</u> 3.2 (3) 19 (3)	5	2008A&ARv16209P Puls, Joachim; Vink, Jorick S.; Najarro, Francisco	12.000 Mass los	Dec 2008 ss from hot ma	<u>A</u> ssive	<u>E</u> stars	X	<u>R</u> !	<u>C</u>		U	
LLANIC CLOUDS (3)	6	 <u>2005ApJ631435R</u> Ramirez-Ruiz, Enrico; García- Segura, Guillermo; Salmonson, Jay D.; Pérez-Rendón, Brenda 	12.000 The Stat Afterglo	Sep 2005 e of the Circun w Appearance	<u>A</u> nstell	E E ar Me	X dium Surrounding	<u>R</u> (g Gar	<u>C</u> c mma-	<u>S</u> Ray B	U Burst Sources and It	s Effect on the
GROUP (3) JNCULUS NEBULA Y LEONIS NEBULA	7	 <u>1992ARA&A30235C</u> Chiosi, Cesare; Bertelli, Gianpaolo; Bressan, Alessandro 	12.000 New dev	n/a 1992 velopments in u	<u>A</u> under	standi	ng the HR di			<u>S</u>	U	Untitle

This page generated using ADS faceted Search, as developed by Michael Kurtz, Alberto Accomazzi & Jonathan Fay









Classification: Reflection Nebula in Cepheus

NGC 7023

RA:

Dec:

Alt:

1.2

		Image	
		lack j	ĺ
Look At	Imagery	1	
Sky	Digitized Sky	y Survey (Optionally)	
-			
Sculptor	Earth	Uranus	ļ



Done

0	0 0		Object Query Results				
•	🕨 🙆 🖌 A A + 🏧 http://adsab	s.harvard.edu/cgi-bin/al	bs_connect?db_key=AST∼_query	=YES&object=NGC%20	7 🛛 🥵 🤇	Q- Google	
Β	DCDCLXV Texify newKodak EXPLO B	ing Clarke WWTSL	Alyssa Good Home Page Toodlede	Harvard IIC: Projects	Wikis Etc.	Google Calendar	\$\$\$ Image Search • >
					ag	oodman@cfa.harvar	d.edu l <u>my Account</u> l <u>Sign off</u>
<u>SA</u>	O/NASA Astrophysics Data System	(ADS)					
Q	ery Results from the Astronomy Da	tabase					Go to bottom of page
Re	rieved 200 abstracts, starting with number 1. T	'otal number selected: 3	993.			Sort	options 🛟
#	Bibcode	Score Date	List of Links				
	Authors	Title	Access Control Help				
1	□ 2009ApJ700.1609M	1.000 08/2009	AZEE LX	<u>RCS</u>	U		
	Myers, Philip C.	Filamentary Structure	of Star-forming Complexes				
2	□ 2009ApJ700.1190D	1.000 08/2009	AZEE LX	<u>RCS</u>	U		
	Desai, Vandana; Soifer, B. T.; Dey, Arjun; LeFloc'h, Emeric; Armus, Lee; Brand, Kate; Brown, Michael J. I.; Brodwin, Mark; Jannuzi, Buell T.; Houck, James R.; and 8 coauthors	Strong Polycyclic Are	omatic Hydrocarbon Emission fro	m z≈2 ULIRGs			
3	2009MNRAS.396.1851N	1.000 07/2009	AZEF LX	R S	U		
	Nutter, D.; Stamatellos, D.; Ward- Thompson, D.	The initial conditions	of isolated star formation - IX. Ak	ari mapping of an ext	ternally heat	ed pre-stellar core	
4	□ 2009A&A502175B	1.000 07/2009	AZEE L	<u>R S</u>	U		
	Boersma, C.; Peeters, E.; Martín- Hernández, N. L.; van der Wolk, G.; Verhoeff, A. P.; Tielens, A. G. G. M.; Waters, L. B. F. M.; Pel, J. W.	A spatial study of the	mid-IR emission features in four I	Herbig Ae/Be stars			
5	□ 2009MNRAS.395.1695H	1.000 05/2009	AZEE LX	RC S	U		
	Hernán-Caballero, A.; Pérez-Fournon, I.; Hatziminaoglou, E.; Afonso-Luis, A.; Rowan-Robinson, M.; Rigopoulou, D.; Farrah, D.; Lonsdale, C. J.; Babbedge, T.;	Mid-infrared spectros	copy of infrared-luminous galaxie	s at z ~ 0.5-3	-		







Done

000					SI	MBAD query	y result					
< r di 🖉	AA	+ 🛋 http:/	//simbad.u-stra	sbg.fr/simbad/	sim-id?ldent	NGC+7023	&jsessionid=B	8F7CD92574727FFC	8 R55 C	Q- Google		
	Texify new	wKodak EX	PLO Bing Cla	arke WWTSL	Alyssa Good.	Home Page	Toodledo	Harvard IIC: Projects	Wikis Etc. v	Google Calendar	SSS Image	Search 🔻 🚿
	Simbad	WizieR	<i>روی</i> Aladin	Catalogs	Dictionary	Biblio	Tutorials	Developers				ſ
				SIMBA	D query re	sult						
other query modes : <u>query</u>	<u>Coordinate</u> <u>guery</u>	Criteria Bi	guery Bas	ic <u>Script</u> ry <u>submission</u>	Output options	Help						
Object query : NGC	7023								CDS.	- SIMBAD4 rel 1.13	2 - 2009.10.23C	EST21:59:31
Available data		Basic data	Identifi	<u>ers</u>	Plot & images		Bibliography	Measure	rments	External a	rchives	Notes
Basic data : NGC 7023 O Other object types:	pen (gala e1•	ctic) Clus	ster	(OCISM) .HII	(LBN) , V* (A	AVSO) .IR	(IRAS)			(query around) with radius 2	aremin
ICRS coord. (ep=2000)	: 21	01 36.9 +68	8 09 48 (-)	[] D	-							
FK5 coord. (ep=2000 e	q=2000) : 21 104	01 36.9 +68	8 09 48 (~) 1926 (~) [[.]	[]D-	-							
Fluxes (1):	B 7	7.20 [-] D	-									
Identifiers (11) :												
NGC 7023 C 2059+679 C1 VDB 139		IRAS 20599 IRAS F2059 LBN 104.08	+6755 9+6755 +14.21			LBN 487 DCISM 50 DC1 235		[BDB2003] G104.00 AAVSO 2044+67	6+14.19			
Plots and Images	(plet	around)	(Aladin previewe) (Nadin app	det)							
	Taulus 10	artinin	-									
References (371 bet Simbad bibliographic surve	ween 1983 : y began in 1950	and 2009)) for stars (at lea	ast bright stars) and	in 1983 for all othe	7 objects (outside	e the solar system	nj.		(display) from: 1	reference summary	ŧ	
												E

.

A

Q• Google

Ç



ADS Faceted Topic Search (alpha)



ADS Home | Abstract Search | Help

"alpha" Faceted Topic Search in ADS (courtesy of Michael Kurtz & Alberto Accomazzi) ٠.

ADS Query Results

🙆 🖉 A A + 🖶 http://adsres.cfa.harvard.edu/cgi-bin/topicFacetSearch?q=PAH;qtype=RELEVANT & Q* Google

SAO/NASA Astrophysics Data System (ADS)

Query Results from the ADS Database

Related Objects	Selected and retrieved 200 abstrac	ts.	Sort options rol Help X R C C S N O U Dust. IV. The Silicate-Graphite-PAH Model in the Post- X R C C S N U X R C C S N U 3 Starlight Intensities in the SINGS Galaxy Sample X R C C S N U Based on Silicate Obscuration and PAH Equivalent Starlight Intensities Starlight Intensities Starlight Intensities
M 82 (14) NGC 7027 (12) NGC 7023 (10) NAME ORI BAR (10)	# Bibcode Authors	Score Date List of Links Title Access Control Help	
NAME RED RECTANGLE (9) QSO B1254+571 (8) NGC 2023 (8) NGC 253 (8) M 17 (8)	1 <u>2007ApJ657810D</u> Draine, B. T.; Li, Aigen	100.000 Mar 2007 <u>A E F X R C</u> Infrared Emission from Interstellar Dust. IV. The Silicate- Spitzer Era	<u>c SN QU</u> Graphite-PAH Model in the Post-
PN G093.9-00.1 (7) NGC 7714 (7) IC 4553 (7) NGC 6240 (6) NGC 292 (5) NAME RHO OPH REGION (5) NAME LMC (5) MCG+10-14-025 (5) 4C 47.36A (5) VV 65 (4)	2 2007ApJ663866D Draine, B. T.; Dale, D. A.; Bendo, G.; Gordon, K. D.; Smith, J. D. T.; Armus, L.; Engelbracht, C. W.; Helou, G.; Kennicutt, R. C., Jr.; Li, A.; and 10 coauthors	96.842 Jul 2007 <u>A E F X R C</u> Dust Masses, PAH Abundances, and Starlight Intensities i	c <u>SNU</u> in the SINGS Galaxy Sample
<u>SBSG 0335-052 (4)</u> <u>QSO B2300+086 (4)</u> <u>NGC 7331 (4)</u> <u>NGC 4151 (4)</u> <u>NGC 1808 (4)</u> <u>NGC 1097 (4)</u> <u>NAME CAMPBELL'S HYDROGEN</u> STAR (4)	3 Description 2007ApJ654L495 Spoon, H. W. W.; Marshall, J. A.; Houck, J. R.; Elitzur, M.; Hao, L.; Armus, L.; Brandl, B. R.; Charmandaris, V.	95.232 Jan 2007 <u>A E F X R C</u> Mid-Infrared Galaxy Classification Based on Silicate Obs Width	<u>c SN U</u> curation and PAH Equivalent
Mrk 273 (4) M 81 (4) M 42 (4) GSC 02342-00359 (4) [KJB2003] G29.957-0.018 (3) [KJB2003] G23.955+0.150 (3)	 2005ApJ628L29E Engelbracht, C. W.; Gordon, K. D.; Rieke, G. H.; Werner, M. W.; Dale, D. A.; Latter, W. B. 	95.090 Jul 2005 <u>A E F X R C</u> Metallicity Effects on Mid-Infrared Colors and the 8 μm	<u>C SN U</u> PAH Emission in Galaxies

list of objects with links to WWT browser (thanks to ADS team & Jonathan Fay)

Go to bottom of page

.

And now we got to NGC 7023 by using the literature as a filter.



Image ssc2003-06f

http://www.spitzer.caltech.edu/Media/releases/ssc2003-06/ssc200 - C

Seamlessness

through...

flickr

astrometry.net

WWT !?



The Spitzer image was obtained with the infrared array camera. Emission at 3.6 microns is shown as blue, emission from 4.5 and 5.8 microns has been combined as green, and 8.0 micron emission is depicted as red.

interacts with the surrounding interstellar medium. These young stars are often detected only in the infrared.

JU AC/A7 is a striking avample of a low many protostor significant and aracting a hisalar artur aided autilow. The control





Coming (Very) Soon...

Historical Image Layer Extracted from ALL ADS holdings (using astrometry.net)

Faceted Heat Map of Articles on the Sky

[e.g.ADS-CDS-WWT]

Prototype of Articles on the Sky (April 2010)



with thanks to CDS/Pierre Fernique

The future is here... data IN articles

Note: This work came from the "'AstroMed" project am.iic.harvard.edu



LETTERS



Figure 2 Comparison of the 'dendrogram' and 'CLUMPFIND' feature identification algorithms as applied to ¹³CO emission from the L1448 region of Perseus. a, 3D visualization of the surfaces indicated by colours in the dendrogram shown in c. Purple illustrates the smallest scale selfgravitating structures in the region corresponding to the leaves of the dendrogram; pink shows the smallest surfaces that contain distinct selfgravitating leaves within them; and green corresponds to the surface in the data cube containing all the significant emission. Dendrogram branches corresponding to self-gravitating objects have been highlighted in yellow over the range of T_{mb} (main-beam temperature) test-level values for which the virial parameter is less than 2. The x-y locations of the four 'selfgravitating' leaves labelled with billiard balls are the same as those shown in Fig. 1. The 3D visualizations show position-position-velocity (p-p-v) space. RA, right ascension; dec., declination. For comparison with the ability of dendrograms (c) to track hierarchical structure, d shows a pseudodendrogram of the CLUMPFIND segmentation (b), with the same four labels used in Fig. 1 and in a. As 'clumps' are not allowed to belong to larger structures, each pseudo-branch in d is simply a series of lines connecting the maximum emission value in each clump to the threshold value. A very large number of clumps appears in **b** because of the sensitivity of CLUMPFIND to noise and small-scale structure in the data. In the online PDF version, the 3D cubes (**a** and **b**) can be rotated to any orientation, and surfaces can be turned on and off (interaction requires Adobe Acrobat version 7.0.8 or higher). In the printed version, the front face of each 3D cube (the 'home' view in the nteractive online version) corresponds exactly to the patch of sky shown in Fig. 1, and velocity with respect to the Local Standard of Rest increases from front (-0.5 km s^{-1}) to back (8 km s^{-1})

data, CLUMPFIND typically finds features on a limited range of scales, above but close to the physical resolution of the data, and its results can be overly dependent on input parameters. By tuning CLUMPFIND's two free parameters, the same molecular-line data set⁸ can be used to show either that the frequency distribution of clump mass is the same as the initial mass function of stars or that it follows the much shallower mass function associated with large-scale molecular clouds (Supplementary Fig. 1).

Four years before the advent of CLUMPFIND, 'structure trees'⁹ were proposed as a way to characterize clouds' hierarchical structure

©2009 Macmillan Publishers Limited. All rights reserved

NATURE Vol 457 1 January 2009

using 2D maps of column density. With this early 2D work as inspiration, we have developed a structure-identification algorithm that abstracts the hierarchical structure of a 3D (p-p-v) data cube into an easily visualized representation called a 'dendrogram'¹⁰. Although well developed in other data-intensive fields^{11,12}, it is curious that the application of tree methodologies so far in astrophysics has been rare, and almost exclusively within the area of galaxy evolution, where 'merger trees' are being used with increasing frequency¹³.

Figure 3 and its legend explain the construction of dendrograms schematically. The dendrogram quantifies how and where local maxima of emission merge with each other, and its implementation is explained in Supplementary Methods. Critically, the dendrogram is determined almost entirely by the data itself, and it has negligible sensitivity to algorithm parameters. To make graphical presentation possible on paper and 2D screens, we 'flatten' the dendrograms of 3D data (see Fig. 3 and its legend), by sorting their 'branches' to not cross, which eliminates dimensional information on the x axis while preserving all information about connectivity and hierarchy. Numbered 'billiard ball' labels in the figures let the reader match features between a 2D map (Fig. 1), an interactive 3D map (Fig. 2a online) and a sorted dendrogram (Fig. 2c).

A dendrogram of a spectral-line data cube allows for the estimation of key physical properties associated with volumes bounded by isosurfaces, such as radius (R), velocity dispersion (σ_v) and luminosity (L). The volumes can have any shape, and in other work14 we focus on the significance of the especially elongated features seen in L1448 (Fig. 2a). The luminosity is an approximate proxy for mass, such that $M_{\text{lum}} = X_{13\text{CO}}L_{13\text{CO}}$, where $X_{13\text{CO}} = 8.0 \times 10^{20} \text{ cm}^2 \text{ K}^{-1} \text{ km}^{-1} \text{ s}$ (ref. 15; see Supplementary Methods and Supplementary Fig. 2). The derived values for size, mass and velocity dispersion can then be used to estimate the role of self-gravity at each point in the hierarchy, via calculation of an 'observed' virial parameter, $\alpha_{obs} = 5\sigma_v^2 R/GM_{hum}$ In principle, extended portions of the tree (Fig. 2, yellow highlighting) where $\alpha_{obs} < 2$ (where gravitational energy is comparable to or larger than kinetic energy) correspond to regions of p-p-v space where selfgravity is significant. As α_{obs} only represents the ratio of kinetic energy to gravitational energy at one point in time, and does not explicitly capture external over-pressure and/or magnetic fields16, its measured value should only be used as a guide to the longevity (boundedness) of any particular feature.



Figure 3 | Schematic illustration of the dendrogram process. Shown is the construction of a dendrogram from a hypothetical one-dimensional

emission profile (black). The dendrogram (blue) can be constructed by 'dropping' a test constant emission level (purple) from above in tiny steps (exaggerated in size here, light lines) until all the local maxima and mergers are found, and connected as shown. The intersection of a test level with the emission is a set of points (for example the light purple dots) in one dimension, a planar curve in two dimensions, and an isosurface in three dimensions. The dendrogram of 3D data shown in Fig. 2c is the direct analogue of the tree shown here, only constructed from 'isosurface' rather than 'point' intersections. It has been sorted and flattened for representation on a flat page, as fully representing dendrograms for 3D data cubes would require four dimensions.



Goodman et al. Nature, 2009



Jim Gray (& Alex Szalay) had it right (in 2004)

All Scientific Data Online

- Many disciplines overlap and use data from other sciences
- Internet can unify all literature and data
- Go from literature to computation to data back to literature
- Information at your fingertips for everyone-everywhere
- Increase Scientific Information Velocity
- Huge increase in Science Productivity



Jim Gray (& Alex Szalay) had it right (in 2004)

The World Wide Telescope an Archetype for Online-Science

Jim Gray (Microsoft)

Alex Szalay (Johns Hopkins University)

Microsoft Academic Days in Silicon Valley

http://research.microsoft.com/~gray/talks

World Wide Telescope Virtual Observatory

http://www.ivoa.net/

 Premise: Most data is (or could be online).

- The Internet is the world's best telescope:
 - It has data on every part of the sky
 - In every measured spectral band: optical, x-ray rad
 - As deep as the best instruments (2 years ago)
 - It is up when you are up.
 The "seeing" is always great (no working at night, no clouds no moons no..).

 It's a smart telescope: links objects and data to literature on them.



All Scientific Data Online

- Many disciplines overlap and use data from other sciences
- Internet can unify all literature and data
- Go from literature to computation to data back to literature
- Information at your fingertips
 for everyone-everywhere
- Increase Scientific Information Velocity
- Huge increase in Science Productivity



The Big Picture



- Data ingest
- Managing a petabyte
- Common schema
- How to organize it?
- How to reorganize it
- · How to coexist with others

- Query and Vis tools
- Support/training
- Performance
 - Execute queries in a minute
 - Batch query scheduling

How do we increase the fraction of astronomy researchers who know about these tools?



+Suggestions?!





CfA Virtual Observatory Users Group

Want to hear more? Please see Gus Muench!

	Blog
	Calendar
	Contributors
1	Description
	Files
	Glossary
	Planning
	Presentations
	Resources
1	Surveys
	Workflows
	Sitemap

Home

Discussion (Google Group)



Join the Discussion

Friends

Astrobetter Astropython IVOA VAO

My recent activity

There are no recent activities.







VIRTUAL ASTRONOMICAL OBSERVATORY

This website provides a platform for sharing resources, workflows, and basic organizational information about networked astronomy databases and tools. Its intended audience includes anyone performing astronomical research online. It originated from the activities of scientists at the Harvard Smithsonian Center for Astrophysics in Cambridge, MA.

By Virtual Observatory (VO), we mean all forms of network tools, databases and websites that are utilized for astronomical research.

By Users Group, we mean a group of individuals who meet approximately monthly to discuss their solutions and problems with doing their research online.

Messages

More on NSF data management... Since the ScienceInsider article, NSF has since issued a press release about requiring data management plans as part of all NSF funding proposals starting in October 2010: "This is the ... Posted May 12, 2010 12:23 PM by August Muench

May 2010 Meeting reminder. This is a reminder that the next VO users group meeting is: Tomorrow, May 7th 10-11am Pratt Conference Room. Data "publishing" is the subject for our meeting and I ...

Posted May 6, 2010 1:41 PM by August Muench

NSF Guidelines on Data Access ScienceInsider reports that NSF is moving towards requiring that a data management plan to be submitted as part of future NSF grant applications. To quote:NSF's current policy requires ...

Posted May 6, 2010 1:23 PM by August Muench

May 2010 Meeting date/time Our next meeting will be 10-11am Friday May 7th in the Pratt Conference room (60 Garden Street). Our inaugural meeting touched on many topics from software to data archiving ...

Posted Apr 28, 2010 1:14 PM by August Muench

How do we increase the number of people who create and interlink new tools?



Now being implemented through VAO "Associates" and WWT Partners.

How do we organize such diverse tools, so as to make them interoperably useful?....

"SAMP" is a great technical start, but offers a very significant user interface challenge.

SAMP



Think about the "modules" needed to make this work...but do the details matter, to your research, if the system works seamlessly?

C	Commo	on "Clo	oud" C	Data	base	es			
	Services (ADS, SIMBA						D, etc.	.)	
Γ/Silve			SAMP						
Browsers				Java VVV			/WT/E	Direc	tΧ
Mac OS X				Windows 7					
Intel Mac									

Seamless Astronomy



Mockup based on work of Eli Bressert, excerpted from NASA AISRP proposal by Goodman, Muench, Christian, Conti, Kurtz, Burke, Accomazzi, McGuinness, Hendler & Wong, 2008

s,,	Top Stories	
loc	Obama Promotes New Health Care Law Voice of America - all 26785 related »	
fТ	Waste issue hurting US nuclear revival-p Reuters - all 92 related »	<u>banel</u>
λο	Dems, GOP Trade Accusations of Politi Exploiting Threats FOXNews - all 900 related »	<u>cally</u>
rsit	Pope accountable for hiding priest abuse U.S. victim Reuters - all 1832 related »	95:
Dive	NYPD: Powder sent to Congressman no hazardous The Associated Press - all 158 related »	<u>n-</u>
₹,	Gmail	
de: ,	Movies: 02421	
800	Astronomy Picture Of the Day (APOD)	
([–])	NGC 2442: Galaxy in Volans	
)!		

Distorted galaxy NGC 2442 can be found in the southern constellation of the flying fish, (Piscis) Volans. Read More

Facebook	
Welcome, Alyssa Goodman loge What's on your mind?	nng
Toodledo - Your to-do list	
1000 CO	3
✤ Hotlist	>
★ Starred	>
Folders	>
Due-Dates	>
A Priorities	>
Recently Completed	>
Mall Tasks	>
Settings	>
	4

Currency Converter

Weather Cambridge, MA 58°F Current: Mostly Cloudy Wind: S at 11 mph Humidity: 41% Sat Fri Sun Thu 65° | 39° 39° | 22° 41° | 31° 50° | 44° College Park, MD Current: Sunny 72°F Wind: S at 11 mph Humidity: 29% Sat Thu Fri Sun 52° | 40° 74° | 49° 49° | 32° 59° | 50° **Google Translate Google Latitude** marginal St 143 reiginta Saratoga + Eagle Hill Idian \$2 Bremen st Boston Street Pa AP 3 90 Jeffries Point General Edward / Airport Lawrence Logan International Airport

010 Google - Map data @2010 Google - Terms of Use



🕈 🕙 🔿	1.harvard.edu/cgi-bin/topicFacetSearch	ADS Query Results 1?q=%22weat%20lensing%22;qtype=INSTRU C Qr Goo	gle	Why?
Apple Yahoo! Google	Maps YouTube Wikipedia News (284) v Popular v		s/why.html C Q- Google
Query Results from the A	ADS Database		Go to bottom of page	
Related Objects	Selected and retrieved 200 abstract	S.	Sort options	Cott.
NAME HDF (12) NAME Chandra Deep Field- South (8) ACO 1689 (8)	# Bibcode Authors	Score Date List of Links Title Access Control Help		ne en c:
ACO 901 (8) ZwCl 0024+1652 (7) ACO 902 (7) ACO 2390 (6)	1	81.000 n/a 2006 A E X Weak Gravitational Lensing	R C c U	1111111
ZwCI 1455+2232 (5) NAME SGP (4) CIG 2137.3-2353 (4) ACO 2219 (4) ACO 2219 (4) ACO 2218 (4)	Facete	ed Browsii	1g"	
ACO 1835 (4) ACO 1763 (4) ACO 370 (4) ACO 267 (4)	3 2003ARA&A41645R Refregier, Alexandre	61.000 n/a 2003 <u>A E F X</u> Weak Gravitational Lensing by Large-Scale Structure	RCCUH re	n easy to use, web based application using the IVOA which allows a user to perform:
ACO 209 (4) ACO 68 (4) 1E 0302.5+1717 (4) ZwCL 1358+6245 (3) NAME HDF-S (3)	4 2008ARNPS5899H Hoekstra, Henk; Jain, Bhuvnesh	51.000 Nov 2008 A X Weak Gravitational Lensing and Its Cosmological A	R C c U pplications	ations
NAME COSMOS FIELD (3) NAME CBI (3) NAME ACO 901A (3) NAME ACO 901-902	5 <u>2003astro.ph6465S</u> Schneider, Peter	44.000 Jun 2003 A X Gravitational lensing as a probe of structure	<u>RCcUH</u>	wsing
SUPERCLUSTER (3) CIG 1224.7+2007 (3) CIG 1054.4-0321 (3) ACO 1942 (3)	6 2006MNRAS.368.1323H Heymans, Catherine; Van Waerbeke, Ludovic; Bacon,	41.000 May 2006 A E G X The Shear Testing Programme - I. Weak lensing and observations	R C c U lysis of simulated ground-based	s, observations to participate in an infrastructure which allows one to:

- create intelligent applications which can reason and inference with
- publish resources as Linked Data, externally indexed.
- easily aggregate metrics of interest to publishers, funding agencies
- let others build applications on this substrate using SPARQL querie

We will be carrying out these efforts as part of ADS Labs.

Bootstrapping in ADS Labs

ADS Labs is an effort to put out more forward thinking, somewhat unstat applications will be incubated in ADS Labs before being pushed out to AD

- The results of queries on a bibliographic database will be made available. We will build a user interface on the above (see first image b)
- 2. We will switch to a semantic backend with a SPARQL interface
- 3. Development on Ontologies (which this site details) continues and
- Finally we'll combine the databases so as to have one large semant below.

Examples of Applications

Here are examples of what such applications might look like:



Ongoing "ADS Labs" Work: Rahul Davé, Alberto Accomazzi, Michael Kurtz, AG Thanks to ADS (NASA)/VAO(NASA+NSF)/MSFT funding.

Article Markup via ADS will be similar to working Life Sciences Tool (thanks IIC!)

	SAO/NASA Astrophysics Data System (ADS)		
	Query Results from the A	ADS Database	Go to bottom of page
	Related Objects	Selected and retrieved 200 abstracts.	Sort options
SCE SCE SCF Annotat	NAME HDF (12) NAME Chandra Deep Field- South (8) ACO 1689 (8)	# Bibcode Score Date List of Links Authors Title Access Control Help	
	ACO 901 (8) ZwCl 0024+1652 (7) ACO 902 (7) ACO 2390 (6)	1 □ 2006glsw.book269S 81.000 n/a 2006 A E X Schneider, P. Weak Gravitational Lensing	R C c U
The neural stem cell microenvironme	ZwC1 1455+2232 (5) NAME SGP (4) CIG 2137.3-2353 (4) ACO 2261 (4) ACO 2219 (4) ACO 2218 (4)	Faceted Browsin	ng"
Save Annotations Original Listings	ACO 1835 (4) ACO 1763 (4) ACO 370 (4) ACO 267 (4)	3 ⊟ 2003ARA&A.41.645R Refregier, Alexandre 61.000 n/a 2003 A E F X Weak Gravitational Lensing by Large-Scale Structure	RCc UH re
Enter term to search for:	ACO 2019 (4) ACO 68 (4) 1E 0302.5+1717 (4) ZwCl 1358+6245 (3) NAME HDF-\$ (3)	4 ⊡ 2008ARNPS5899H Hoekstra, Henk; Jain, Bhuvnesh 51.000 Nov 2008 A X Weak Gravitational Lensing and Its Cosmological A	<u>RCc</u> U pplications
Terms: neurogenesis :59 (<u>x d a s o</u>) growth :30 (<u>x d a h o</u>) cell-cell signaling :28 (<u>x d a s o</u>	NAME COSMOS FIELD (3) NAME CBL(3) NAME ACO 901A (3) NAME ACO 901-902	5 ⊇ 2003astro.ph6465S 44.000 Jun 2003 ▲ X Schneider, Peter Gravitational lensing as a probe of structure	R <u>C</u> c <u>U</u> H
	SUPERCLUSTER (3) CIG 1224.7+2007 (3) CIG 1054.4-0321 (3) ACO 1942 (3)	6 <u>⊇ 2006MNRAS.368.1323H</u> 41.000 May 2006 <u>A</u> <u>E</u> <u>G</u> X Heymans, Catherine; Van Waerbeke, Ludovic; Bacon, observations	<u>R</u> <u>C</u> <u>c</u> <u>U</u> alysis of simulated ground-based
Annotation id 6. Updated at 2009-07-20 19:47:51.020766 by Mining Robot.			

- Ilias Kazanis¹⁴,
- · Justin Lathia1.2,
- . Lara Moss1.3,
- Charles ffrench-Constant^{1.3}

Department of Pathology University of Combridge Teanin Court Read, CP2 SOR Combridge UK	
Delete All [Add Mapping	
² Current address: Dept. of Surgery, Division of Neurosurgery, Duke University Medical Center, Durham, NC 27710, USA go : nervous system development	
³ Current address: MRC Centre for Regenerative Medicine, The Queen's Medical Research Institute, 47 Little France Crescent Delete Alter	
In mammals, neural stem cells appear early in development and remain active within the central nervous system for the whole go : cell growth Delete Alter	s they
and reside within changing microenvironments whilst retaining the basic properties of a stem cell: multipotentiality and the at go : growth pattern Delete Alter Inclogical characteristics	teristi
along with the fundamental structural components and signalling molecules of their microenvironments. In early neural devel. go : cell-cell signaling Delete Alter em is established	d, neu
cells; they are situated among other neuroepithelial cells and they are exposed to various signals such as retinoic acid, sonic go : system development Delete veurogenesis con	nmen
glial cells and the complexity of their microenvironment increases due to the emergence of various types of neuronal progenit Alter	ally, d
astrogial morphology and reside in specific microenvironments that are cared neurogenic nones; small neurogenic islands w	101 01
go . growth boliete frater	
go : stem cell development Delete	
1. The empryonic neural stem cell (NSC) microenvironment	
Central nervous system (CNS) development is an intricate process relying on a series of mechanisms precisely regulated in ti go : cell development Delete Alter is present in the	adult

Annotate

spective destination within an approximately one-week period during embryogenesis. The embryonic CNS is a dynamic structure, constantly increasing in size due to histogenesis, while the ster

Collaborative Astronomy at University of Washington

Research in a Browser

– "iGoogle" for Astronomy

- Collections of simple atomic applications (gadgets)
- Users choose the view they want
- All gadgets can communicate with each other

Customizable and sharable

- Users can build and share "mashups"
- Widgets are simple to create
- Widgets call virtual observatory resources

Efficient

- Communication is within the browser (fast)
- Built from javascript (standard)













Collaborative Astronomy (Connolly, Gibson, Krughoff, Sayers, Smith 2010)

Create, store and share multiple views of gadgets



Interaction allows selections to be shown on the viewport



Collaborative Astronomy (Connolly, Gibson, Krughoff, Sayers, Smith 2010)



Part 2 WorldWide Telescope Ambassadors Program

Alyssa Goodman

Harvard University Professor of Astronomy, WGBH Scholar-in-Residence, Microsoft Academic Partner

Annie Valva WGBH Interactive, Director of Research & Development

Pat Udomprasert

WWT Program Coordinator



WWT Ambassadors

Who?

Harvard/CfA,WGBH and Microsoft Research staff in collaboration with Volunteer Ambassadors

What?

Future-leaning way to teach and learn STEM concepts

How?

Use new WWT platform to give experts and learners access to the Universe

Where?

Public spaces and schools in a variety of regions

Where? ... and When? Public spaces and schools in a variety of regions

Pilot 🔵 Boston Area

Phase I candidates Tucson, AZ; Seattle, WA; Appalachia; Gainesville, FL; Fairbanks, AK



Phase II: US-wide; Phase III: International



Harvard University, WGBH & Microsoft Research Alyssa Goodman, Patricia Udomprasert, Annie Valva & Curtis Wong



hat is WorldWide Telescope and its Ambassadors Program? InddWide, Telescope (WWT) is a fastatic Tilebarra Information St

WorldWide Telescope (WWT) is a fantastic "Universe Information System" created primarily by Curtis Wong and Jonathan Ray at Microsoft Research. It functions as a Virtual Astronomical Observatory linking its users to much of the world's store of online data and information about our Universe. WWT is evolving to become a key research tool within the online autonomy ecosystem known in the US presently as the "VAO" (see A. Goodman's "Seamless Astronomy" talk at this meeting), but it also offers unprecedented new opportunities for STEM outreach.

The WorldWide Telescope Ambassadors Program promotes WWT as a future-leaning way to teach and learn STEM concepts by recruiting astronomically-liberate volunteers who are trained to be experts in using WWT as a teaching tool.

Who are we?

Cur current collaboration brings together professional astronomers and science educators at Harvard, computational virtuosos at MS Research, and STEM education and outreach specialistis at WGBM. The next phase of the project (see table below) will include participants from selected areas within the US, including Washington, Florida, Arizona, Alaska, and Appalachia.

Who are the WWT Ambassadors, and what do they do?

WWT Ambassadors are carefully recruited for training from amongst: 1) retired STEM professionals and amateur astronomers with a demonstratile deep knowledge of astronomy and physics; 2) undergraduate and graduate students and postdoctoral fellows in Astronomy and Physics; and 3) science teachers. In their training, Ambassadors learn how to use WWTS tools in general, and also how to create and publish guided "Burs" of astrophysical concepts. These Tours allow users to display beautiful astronomical images in their proper context in the night sky, while demonstrating the physical principles at work in those images. Ambassadors can create and use materials within WWT; give volunteer presentations at variety of public venues; help out in classroom settings; or choose to do more than one of the above!



What have we done so far?

Our program began in the fail of 2009. Initial Ambassadors are currently working with 80 middle achool students and their backet, Michelle Bactley, at the Clarke Middle School in Leongton, MA, helping the students to prepare tours within WWT based on a six-week-long research experience. WWT and its Ambassadors have generated tremendous enthusksem from the students, and have inspired quality learning through exploration and discovery. Results from the Riot at Clarke are being collected online through a dedicated commenting site open to all students, and an analysis of the Riot experience will serve to inform the NSF proposal being submitted to expand the program in the Spiring of 2010.

and discovery. Versions, examine, not, incredible (report at time)? of commenting inform the NSF "The set option the weilest ground and gas bent to regar and to maginess from the tables." Theory is every spores of the table and and the bills. It's but resulting.

The same we also as several map of the universe."

What's the whole plan, and what are the program's goals?

We are presently preparing a proposal to the National Science Foundation, based in large part on our "Plich" experience, to implement "Phase I" of the Ambassadors Project (see table), where we will begin a limited expansion within the US, carefully selecting cities and partners where we will be able to maximise success with the available misources, while increasing the socioeconomic diversity of our sites. We plan to expend nationally in Phase II, and internationally in Phase III, while minimal advertising, we have already received inquiries from dozens of interested and qualified potential volunteers in multiple states and countries.

A critical goal of this project is to create a **full astronomy curriculum using WWT Tours created by our Ambassadors**. These Tours will be vetted by the astronomy and science education professionals within our collaboration, and they will be freely available, centrally managed, and searchable, through web services at WGBH. The entire WWT Ambassadors "Tour Curriculum" will be integrated with WGBH Teachers' Domain, which currently has nearly 400,000 registered usars.

WorldWide Telescope can help change how students learn science by demonstrating the joys of inquiry and discovery, and the WWT Ambassadors Program is designed to help to increase science literacy in the general public while forming intergenerational connections within their communities.



To find out more...

Please see Pat Udomprasert, Sarah Block, Jeremy Cushman, Sana Sharma, me and the web and the wall outside this room...

A A + C http://www.cfa.harvard.edu/WWTAmbassadors/

WorldWide Telescope Ambassadors Program

About Galleo Tour Project Team How to get involved Tour-making Tutorials Documents Events Protected WorldWide Telescope

Galileo Tour

See a video of ou

teractive Tour in WW7

ing Galileo's his

ations of Jupiter'

environment that functions as a virtual telescope, ving anyone to make use of professional astronomical data to explore and understand the universe. As of early 2010, the new WWT Ambassadors Program is recruiting astronomically literate volunteers, including retired scientists engineers-all of whom will be trained to be experts in using WWT as a teaching tool. Ambassadors will give volunteer presentations at public libraries, community centers, museums, and schools, demonstrating WWT's power to help laypeople visualize and understand our universe. Ambassadors will learn how to create and publish guided "tours" of astrophysical concepts, which allow users to display beautiful astronomical images in their proper context in the night sky, while demonstrating the physical principles at work in those images.

WorldWide Telescope (WWT) is a rich visualization

Tour creators will be able to draw upon and link tours to highly vetted multimedia content from NOVA, the renowned PBS multi-platform series produced by WGBH. Virtual tours will be freely available and centrally managed in order to form a comprehensive astronomy curriculum for both formal and informal educational use. The tours will be searchable and distributed online from popular websites such as NOVA Online and WGBH Teachers' Domain, touting almost 400,000 registered users. [www.teachersdomain.org]

WWT Ambassadors will help to increase science literacy in the general public while forming intergenerational connections within their communities.



C Q- Google

6th grade students at Clarke Middle School, Lexington, MA learn about the universe using the WorldWide Telescope



WWT allows users to explore our universe in rich detail, from our solar system out to the largest observed structures in the cosmos.

www.cfa.harvard.edu/WWTAmbassadors/

What? Future-leaning way to teach and learn STEM concepts



WWT Tours,

including creation by Ambassadors & learners + hosting WorldWide Telescope Scavenger Hunt Mac Web Client version Grade 6 Science Name

This Scavenger Hunt is designed to help you learn how to navigate around and research objects in the night sky using the WorldWide Telescope Web Client.

You and your partner(s) will search within WorldWide Telescope for various items, and answer questions about the things that you find.

Some Tips:

"Home" settings.

If, on your explorations, you find yourself in a state where what you see doesn't match up with what's described in this sheet, please verify that the settings are as follows:

- 1. Along the top row, click "View." (note that you should click the top part of the button, not the lower part with a little downward pointing triangle in it.)
 - a. In the lefthand box: uncheck everything except "Figures" and "Ecliptic."
 - b. In the 2nd box from the left, check everything.

Guided WWT

Exploration

activities created by program staff & Ambassadors/teachers

How?

Using new WWT platform to give experts and learners access to the Universe

WWT Ambassadors Program Recruiting, Vetting, Coordination

ours



data, literature, media

WWT

Ambassadors









WorldWide Telescope Scavenger Hunt Mac Web Client version Grade 6 Science

Name

This Scavenger Hunt is designed to help you learn how to navigate around and research objects in the night sky using the WorldWide Telescope Web Client.

You and your partner(s) will search within WorldWide Telescope for various items, and answer questions about the things that you find.

Some Tips:

"Home" settings.

If, on your explorations, you find yourself in a state where what you see doesn't match up with what's described in this sheet, please verify that the settings are as follows:

- 1. Along the top row, click "View." (note that you should click the top part of the button, not the lower part with a little downward pointing triangle in it.)
 - a. In the lefthand box: uncheck everything except "Figures" and "Ecliptic."
 - b. In the 2nd box from the left, check everything.



Clarke Middle School, Lexington, MA (WWT Ambassadors Pilot School)

"Why is one polar ice cap on Mars bigger than the other?" - Clarke Middle School 6th Grader





Michelle Bartley interviews her 6th-grade science class about WWT December 19, 2009



"I never knew programs like this could even exist. It's just amazing." –Clarke Middle School 6th grade student

More quotes from Clarke 6th Graders

"Learning about our Universe by actually seeing and exploring it makes it easier to contemplate and more fun."

"You can explore the Universe yourself and you don't always have to only learn from the teacher."

"It gave me a better mental map of the universe."

(And of the 72 surveys we've collected, 71 are positive toward WWT Ambassadors.)

Spring 2010 Pilot Results

15 WWT Tours Created by groups of 4 to 5 students in 4 sixth-grade Science Classes (of M. Bartley, Clarke MS)

Facilitated by 4 WWT **Ambassadors**: I Ph.D. Astrophysicist (Udomprasert); I Harvard Undergrad Physicist (Cushman); I retired Ph.D. Physicist (Post); I planetarium-show writer (Becker)

Students, parents & teachers report creating WWT Tours is a better vehicle for learning than PPT presentations, papers

Unexpected finding: Tour creation is a non-threatening way to reveal students' understanding/mis-understanding of concepts

Why? Increase STEM literacy in US now. Demonstrate cyberlearning's value to the "Cone of Experience"



Edgar Dale, "Audio Visual Methods in Teaching", 1946-69

	PROPOSED INDEX & SEARCH SCHEME FOR WWT AMBASSADORS	
	WWT TIPS AND TRICKS	
	POPULAR INTERESTS	
	 Beauty in the Night Sky Space Exploration 	
Tour	 Exotic Objects and Events 	
Tour.	Collisions and Explosions	Tamparatura
Duct & Lla	♦ For and By Kids	Temperature
Dustaus	♦ NIGHT SKY OBSERVING	
	Beauty in the Night Sky GSTCODVSICISTS 3S	
	Constellations	
Dlanata	Finding Objects in the Sky	
Flanets	Brightness and Magnitudes V V V I AIIIDASSAGOIS	
	Telescopes and Technology	
	Solar System - Ours and Beyond to create SIEM-	
	 Stars and Their Evolution 	
Stars	Between the Stars - Gas and Dust	
Stars	♦ Galaxies OTETLEU TOULS	
	Cosmology	
	Exotic Objects (hint. hint)	
	ASTROPHYSICAL CONCEPTS	
Galaxies	Light Scales in the Universe and we'll help you!	
JaiaAiCS	Scales in the Universe and we in the Universe And we in the Universe	
	 Changes with Time 	
	♦ Big Theories in Physics	
	SCIENTIFIC PROCESSES AND SKILLS	
Gas	How We Know	
Gas	Limitations to What We Can Know	
	SCIENCE AND SOCIETY	
	♦ Current Research	
Exotica	History of Science	
(Black Holes	Cultural Influences and Perspectives	
	 Space Exploration 	
etc.)		







http://www.skyalert.org/static/wwt.html

How best to work with NSF OCI, MPS/AST, EHR & CISE simultaneously?

Proposals/Phases

>RAPID/EAGER?
>Unsolicited?
>Cyberlearning?
>Other?

We are ready-to-go in 2010 on Phase 1. Phase II possible in 2011-12.

