

Rediscovering the Milky Way



Alyssa A. Goodman (Harvard-Smithsonian Center for Astrophysics)

with collaborators at (alphabetically by institution):

Boston University: James Jackson

Caltech: Jens Kauffmann

Harvard - Smithsonian: Christopher Beaumont, Michelle A. Borkin, Thomas M. Dame

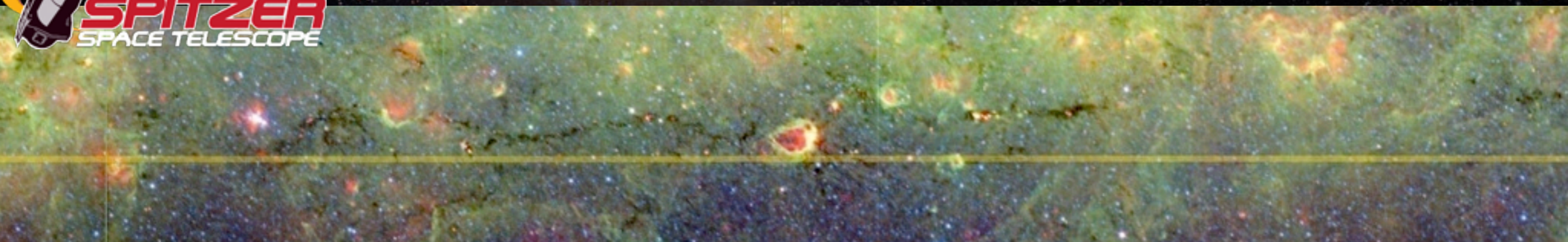
Max Planck Institute for Astronomy: Thomas Robitaille

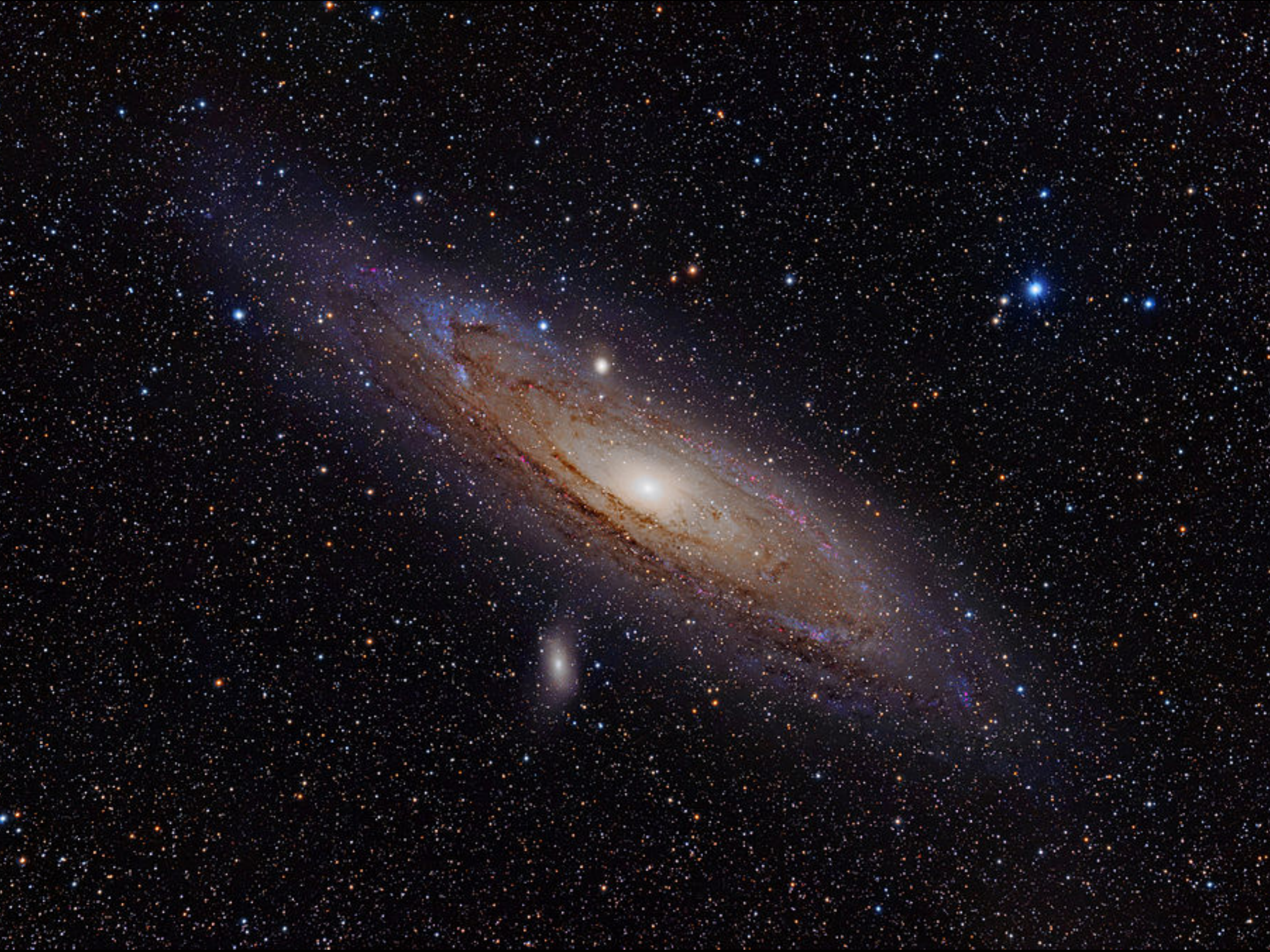
U. Munich: Andreas Burkert

U. Vienna: Joao F. Alves

U. Wisconsin: Robert A. Benjamin

Music: Davis Jerome, Richard Woodhams & The Mozart Orchestra - Oboe Concerto in C Major: II. Adagio, by Sir William Herschel





The Spitzer Infrared Nearby Galaxies Survey (SINGS)

The Spitzer Space Telescope observed 75 galaxies as part of its SINGS (Spitzer Infrared Nearby Galaxies Survey) Legacy Program. The galaxies are presented here in a Hubble Tuning-Fork diagram, which groups galaxies according to the morphology of their nuclei and spiral arms. The designation of these galaxies and their placement in the diagram is based on their visible-light appearance. The main goal of the SINGS program is to characterize the infrared properties of a wide range of galaxy types. The images of the galaxies are composites created from data taken by IRAC (the Infrared Array Camera) at 3.6 and 8.0 μm , and MIPS (the Multiband Imaging Photometer for Spitzer) at 24 μm .

The infrared range probed by these and other observations taken for the SINGS project allows for the detailed study of star formation, dust emission, and the distribution of stars in each galaxy. Light from old stars appears as blue in the images, while the lumpy knots of green and red light are produced by dust clouds surrounding newly born stars. The elliptical galaxies on the left are almost entirely made of old stars, while spiral galaxies like our own Milky Way are rich in young stars and the raw materials for future star formation.

More information can be found at:
<http://sings.stsci.edu/>



Ellipticals

Irregulars

Strong Bulge

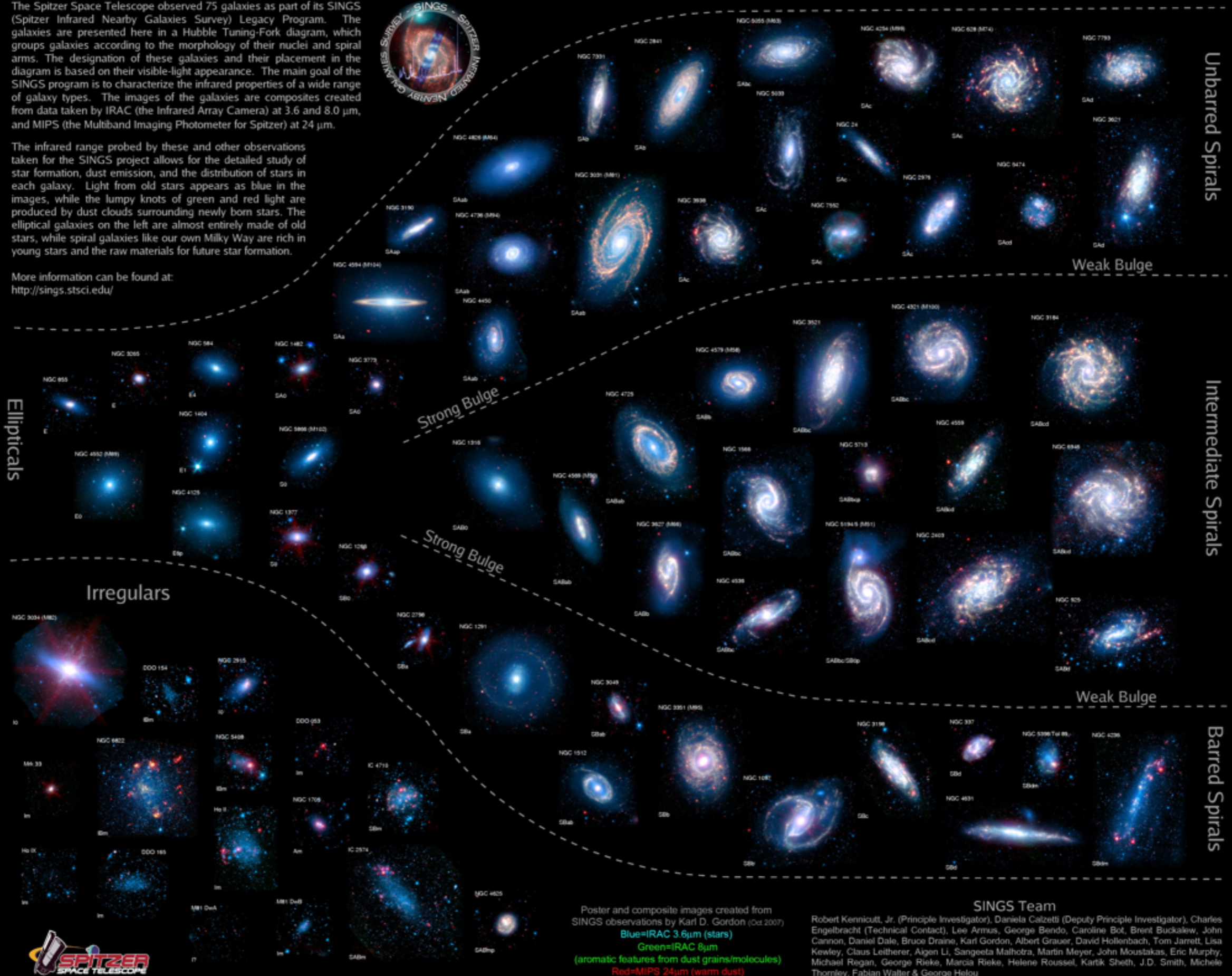
Strong Bulge

The Spitzer Infrared Nearby Galaxies Survey (SINGS) Hubble Tuning-Fork

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Poster and composite images created from SINGS observations by Karl D. Gordon (Oct. 2007)
 Blue=IRAC 3.6 μm (stars)
 Green=IRAC 8 μm (aromatic features from dust grains/molecules)
 Red=MIPS 24 μm (warm dust)

SINGS Team
 Robert Kennicutt, Jr. (Principle Investigator), Daniela Calzetti (Deputy Principle Investigator), Charles Engelbracht (Technical Contact), Lee Armus, George Bendo, Caroline Bot, Brent Buckalew, John Cannon, Daniel Dale, Bruce Draine, Karl Gordon, Albert Grauer, David Hollenbach, Tom Jarrett, Lisa Kewley, Claus Leitherer, Aigen Li, Sangeeta Malhotra, Martin Meyer, John Moustakas, Eric Murphy, Michael Regan, George Rieke, Marcia Rieke, Helene Roussel, Kartik Sheth, J.D. Smith, Michele Thornley, Fabian Walter & George Helou



1936: "The Realm of the Nebulae" by Edwin Hubble

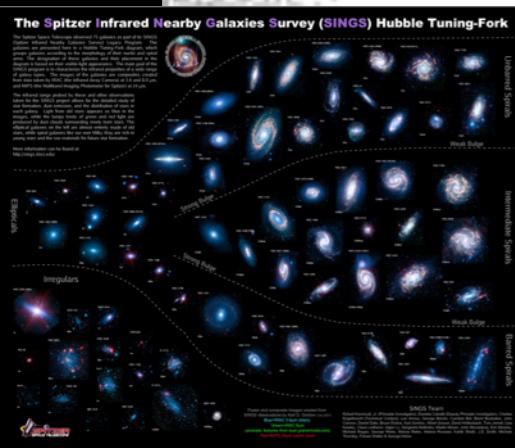
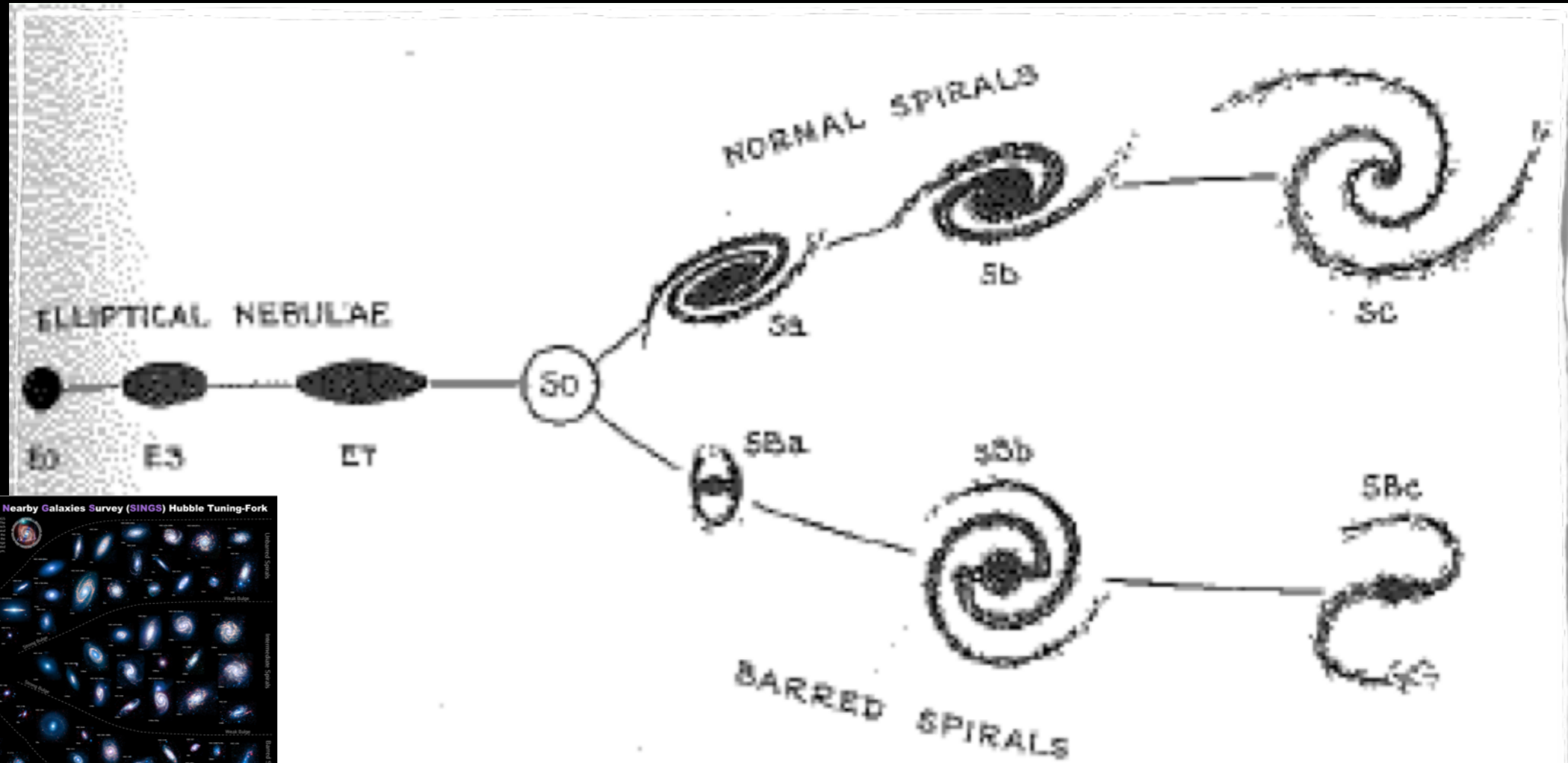


FIG. 1. *The Sequence of Nebular Types.*

The diagram is a schematic representation of the sequences of classification. A few nebulae of mixed types are found between the two sequences of spirals. The transition stage, S0, is more or less hypothetical. The transition between E7 and SB_a is smooth and continuous. Between E7 and S_a, no nebulae are definitely recognized.

“Hubble’s Tuning Fork Diagram”



The Shapley-Curtis Debate at the Smithsonian Natural History Museum, 1920

From National Academy of Sciences,
Smithsonian Institution, Washington, D. C.
(Carl H. Butman, Representative).

For Release to Afternoon Papers,
Monday, April 26

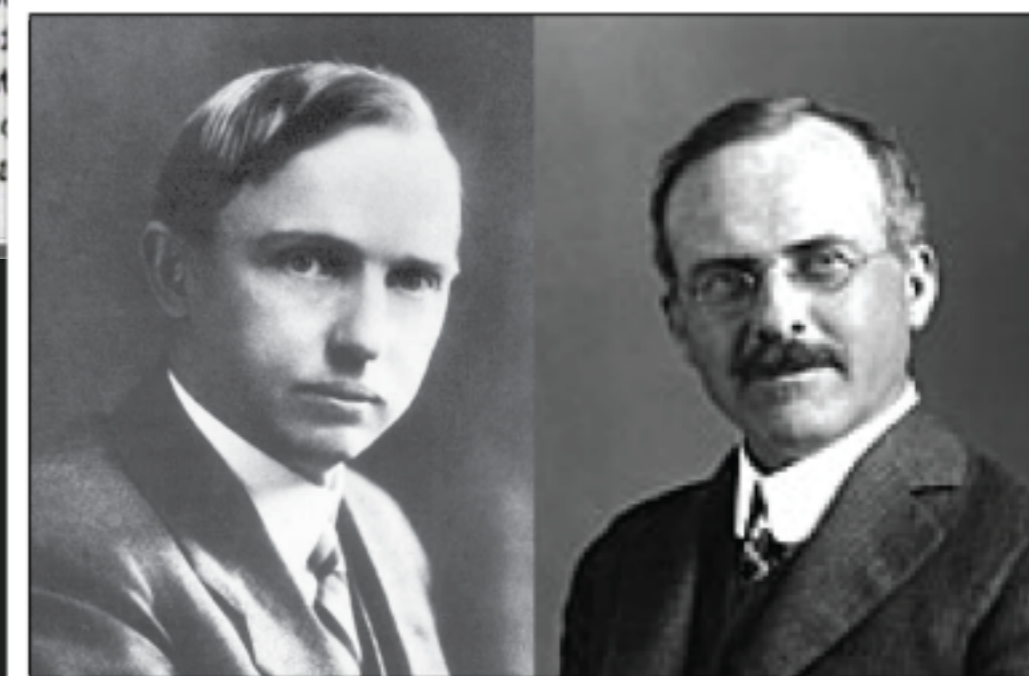
HOW MANY UNIVERSES ARE THERE?

This evening two California astronomers will discuss the Size of the Universe, and present their views as to whether or not there is only one or several universes, before the National Academy of Sciences, which is now in session in Washington.

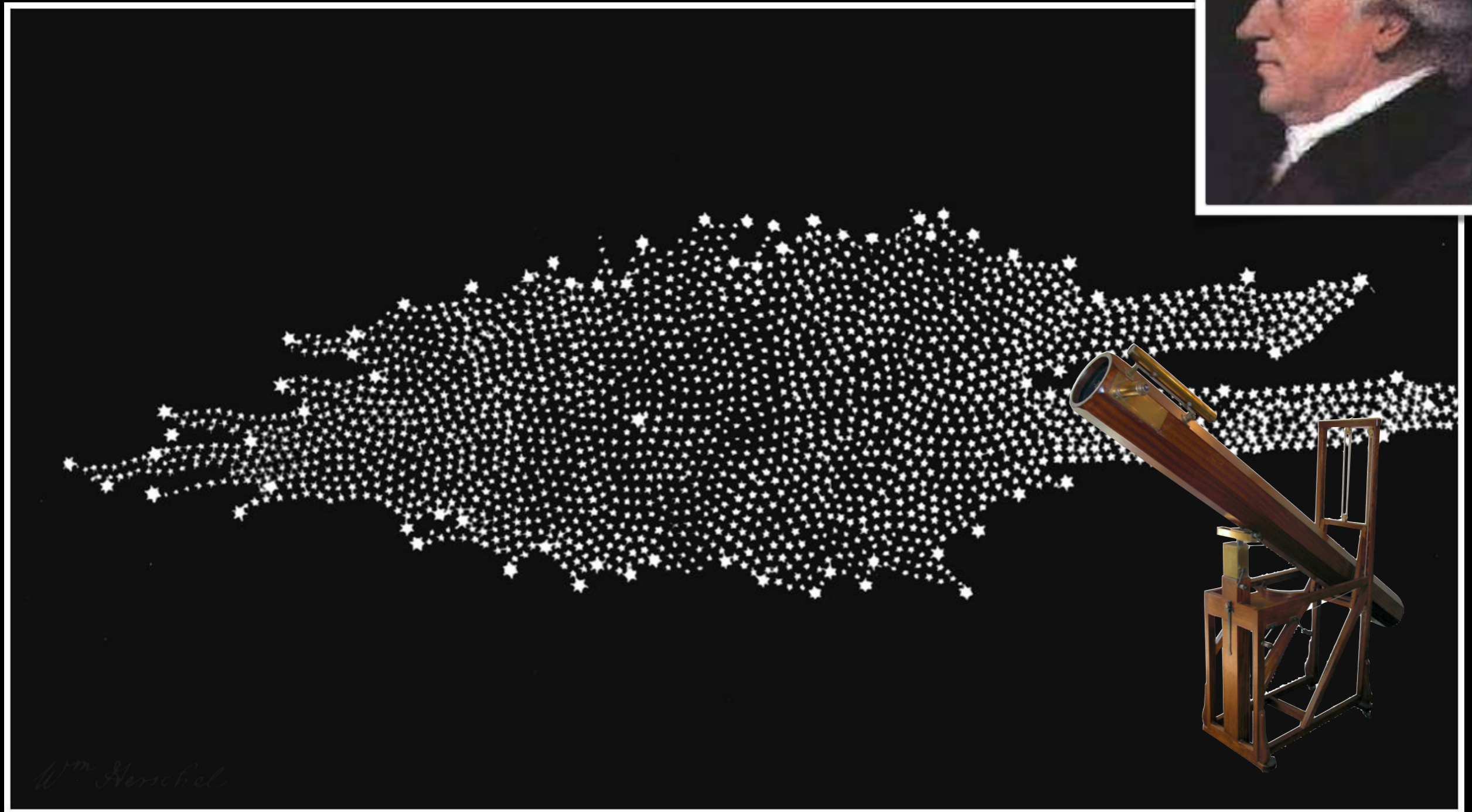
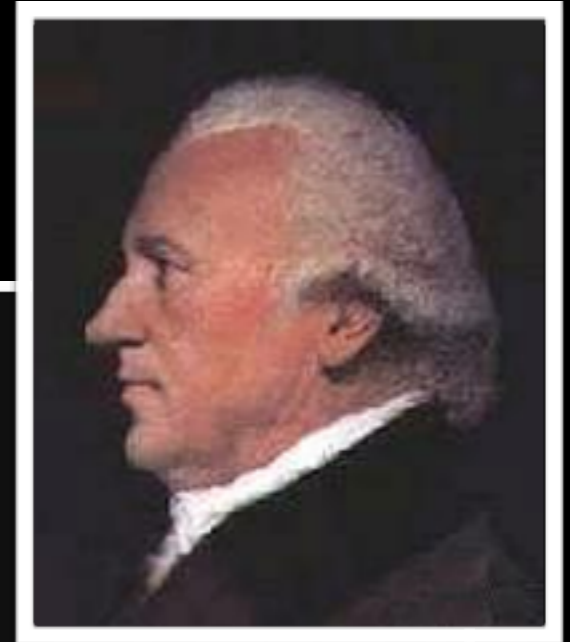
In this public meeting, Dr. Harlow Shapley of the Mt. Wilson Solar Observatory, will discuss recently secured evidence pointing to the dimensions of our galaxy of stars, known popularly as the Milky Way, which he believes to be ten times greater than is held in the older theories concerning the dimensions and composition of the Milky Way. In other words, he claims that it takes light about three hundred thousands of years to cross from one side to the other of the space occupied by the 3,000,000,000 stars of which our sun is the nearest one. He holds the spiral nebulae, those clam-shell-like cloudy luminous objects seen by great telescopes, to be inside our system.

Doctor Shapley's views will be followed by the discussion of Doctor Heber D. Curtis of the Lick Observatory, who will defend the older view that our Milky Way is approximately of the dimensions suggested by Newcomb, about 30,000 light-years in diameter, with the spiral nebulae regarded as very probably individual galaxies of "island universes", like ours. Thus there may be a million other universes each having 3,000,000,000 stars. Inhabitants of numerous universes would see our Milky Way as a spiral nebula. The lectures of these two learned astronomers will be followed by a general discussion of the auditors present who are interested in the development of this new work in scientific research.

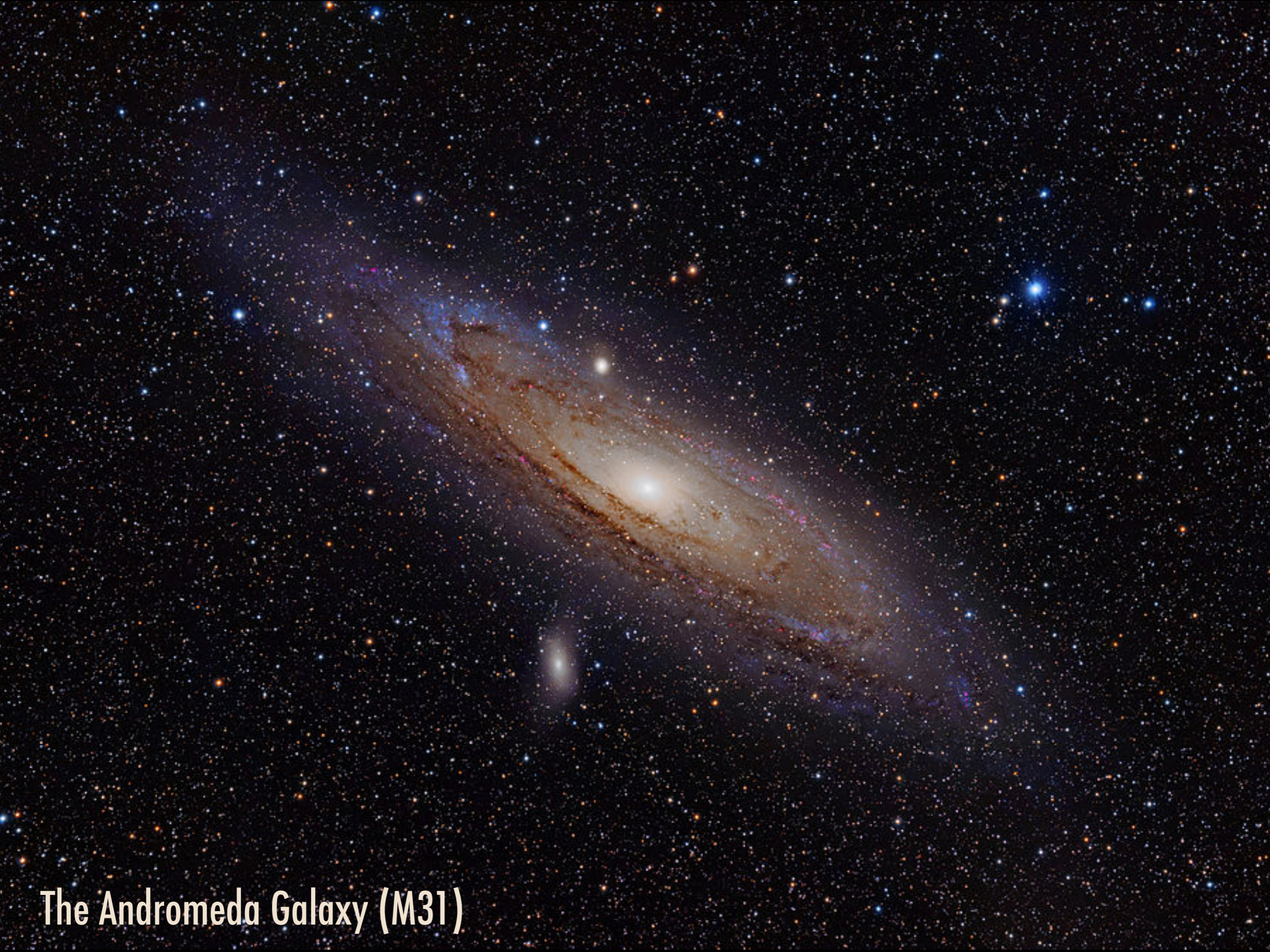
The Shapley-Curtis Debate at the Smithsonian Natural History Museum, 1920



William Herschel's Milky Way Galaxy in 1781



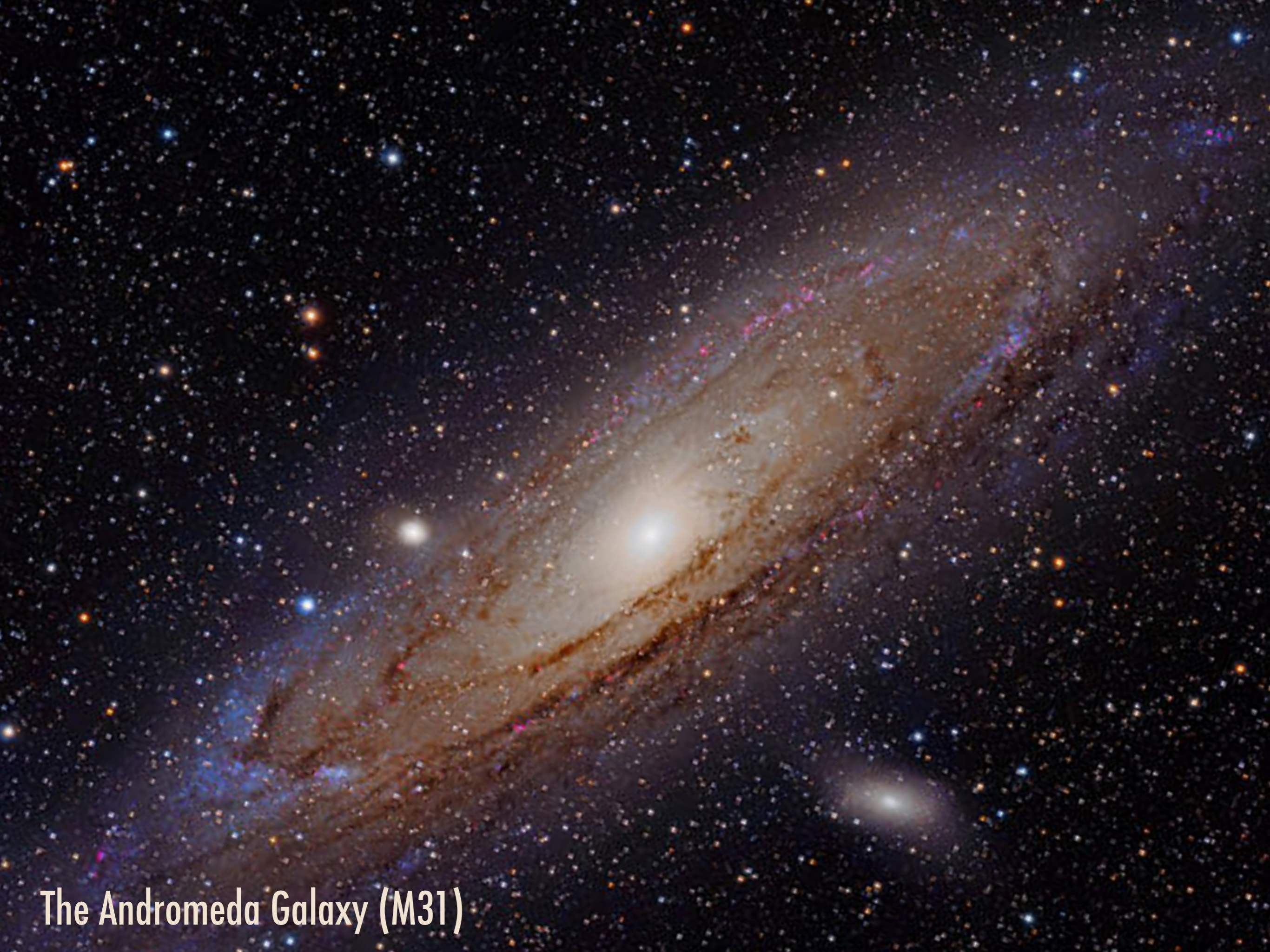
More info at <http://cosmology.carnegiescience.edu/timeline/1781> (Herschel); <http://cosmology.carnegiescience.edu/timeline/1920> (Shapley-Curtis)



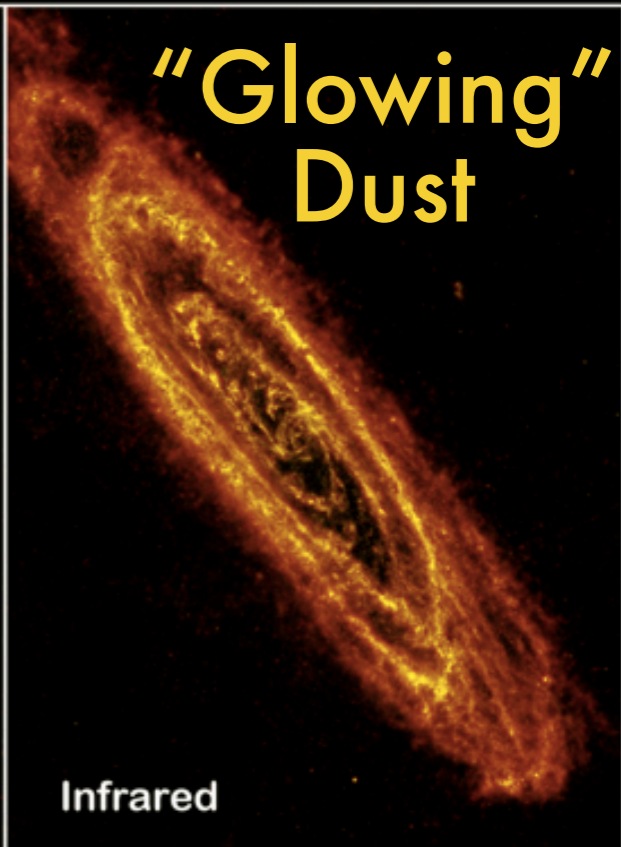
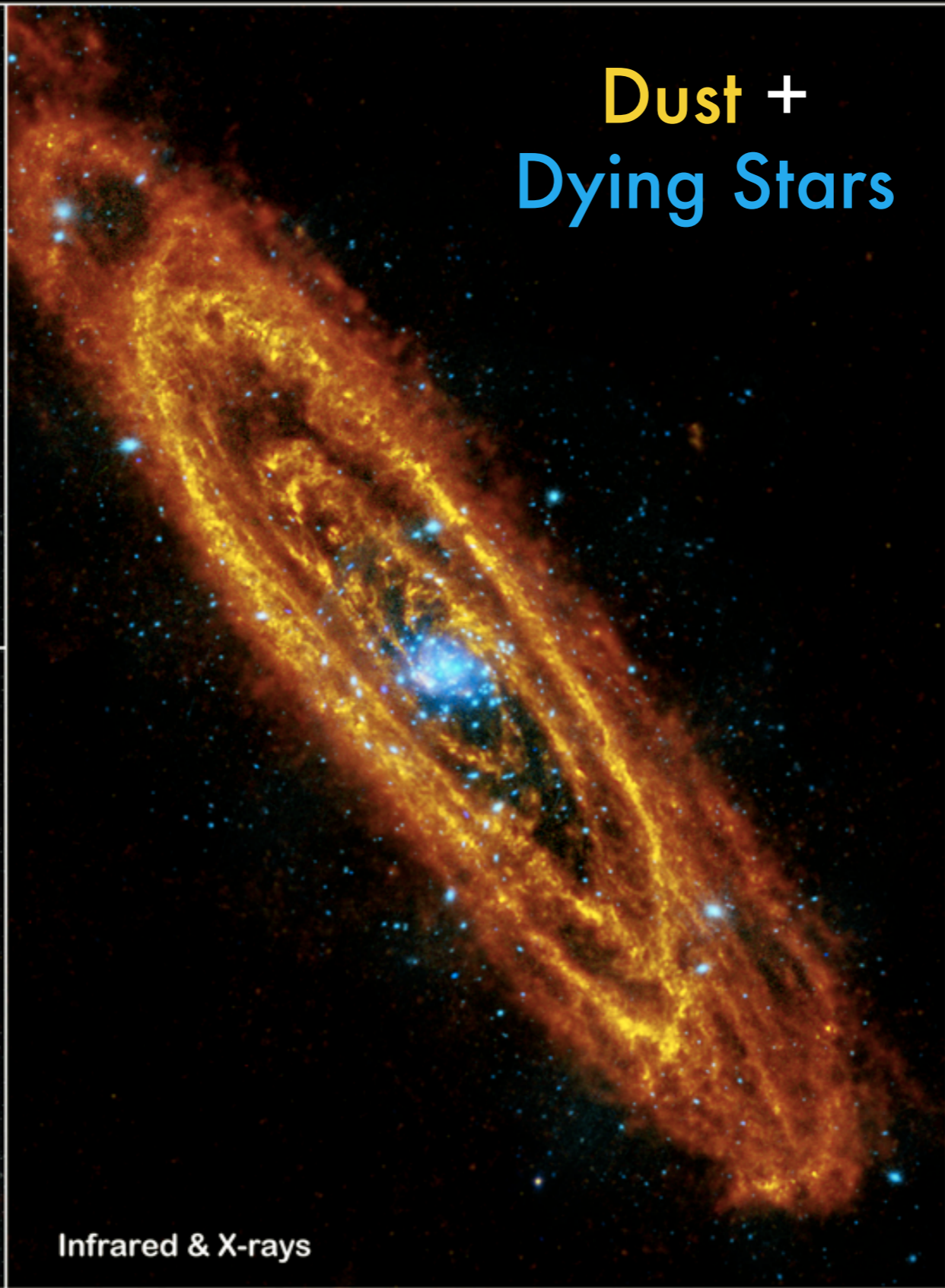
The Andromeda Galaxy (M31)



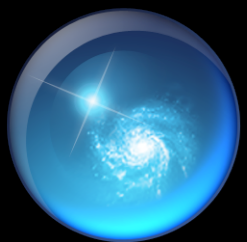
The Andromeda Galaxy (M31)



The Andromeda Galaxy (M31)



The Andromeda Galaxy (M31)



Microsoft® Research WorldWide Telescope

[demo]

worldwidetelescope.org

The screenshot shows the WorldWide Telescope interface with a top navigation bar containing 'Explore', 'Guided Tours', 'Search', 'View', and 'Settings'. Below this is a 'Collections > All-Sky Surveys >' section with a row of eight image thumbnails: 'Digitized Sky Survey', 'VLSS: VLA Low-frequency Sky Survey', 'WMAP ILC 5-Year Cosmic Microwave Background', 'SFD Dust Map (Infrared)', 'IRIS: Improved Resolution', '2MASS: Two Micron All Sky Survey', and 'Hydrogen Alpha Filter'. The main view area shows a large, detailed image of a galaxy with a central crosshair. A 'Finder Scope' window is open, displaying a smaller image of a galaxy and the following text: 'Classification: Spiral Galaxy In Andromeda', 'NGC224', 'RA: 00h42m42s Magnitude: 3.4', 'Dec: 41 : 16 : 00 Distance: 2.5 million light years', 'Alt: 70 : 06 : 26 Rise: 00:35', 'Az: 275 : 42 : 17 Transit: 00:35', 'Set: 00:35', and 'Image Credits: Data provided by two NASA satellites, the Infrared Astronomy Satellite (IRAS) and the Cosmic Background Explorer (COBE). Processing http://astro.berkeley.edu/~marc/dust/'. Below the Finder Scope is a 'Look At' dropdown menu set to 'Sky', with 'Andromeda' selected. To the right is a 'Context bar' showing 'NGC221' and 'M31' with their respective thumbnails. Further right is a 'Context globe' showing the current field of view on a celestial sphere. At the bottom right, a 'Context globe' shows the current location on a celestial sphere with coordinates 'RA : 00h42m40s' and 'Dec : 41:13:35'. A '3D' button is visible in the bottom right corner.

Seamlessly explore imagery from the best ground and space-based telescopes in the world

Expert led tours of the Universe

Control time to study how the night sky changes

View and compare images from across the electromagnetic spectrum

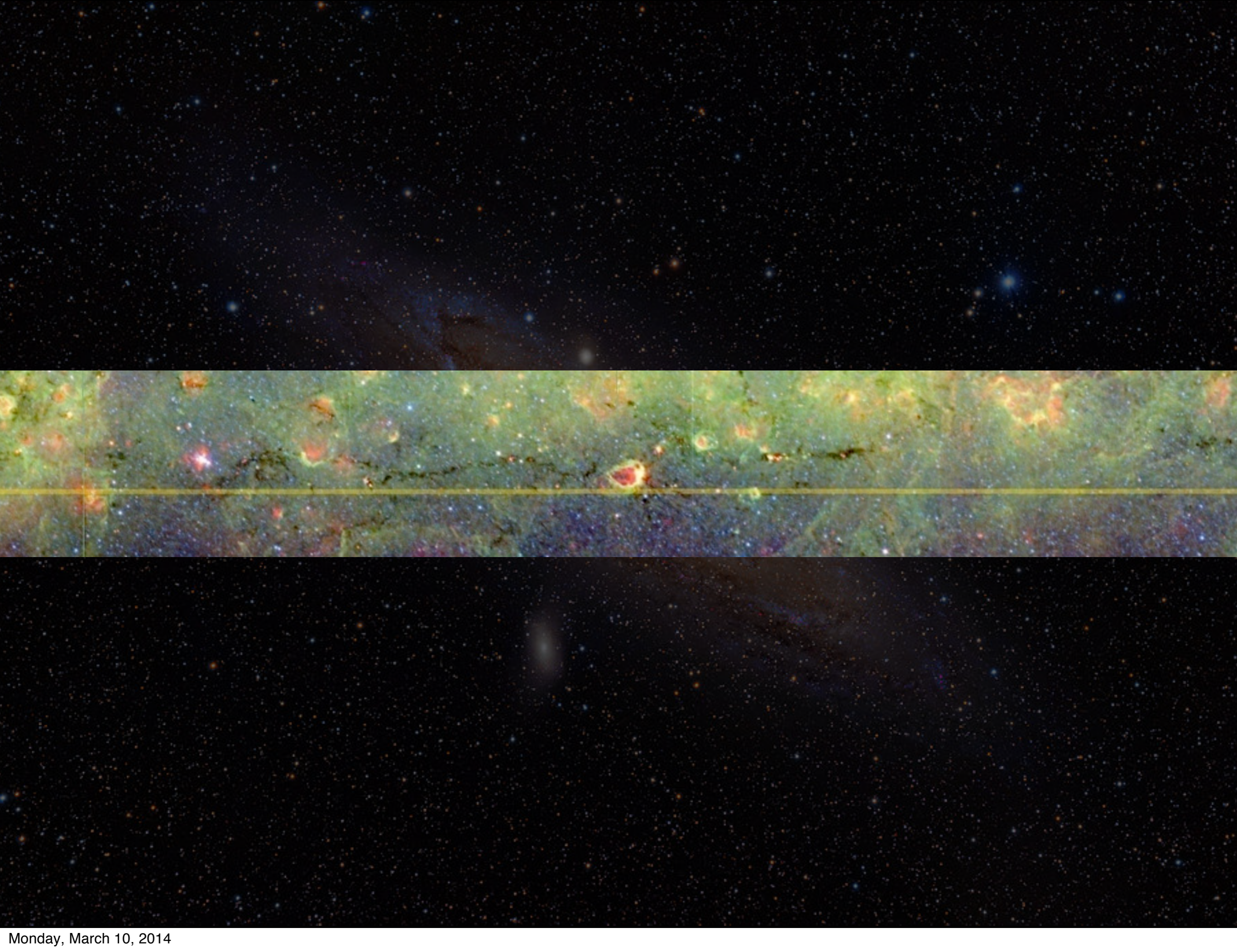
Much more than "just" the sky at night! 3D features can take you to other planets, stars & galaxies.

Finder Scope links to Wikipedia, publications, and data, so you can learn more

Context bar shows items of interest in current field of view

Context globe shows where you're looking.







"The Bones of the Milky Way"

Open Access, Open Data (Authorea.com)

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The Bones of the Milky Way

[Alyssa Goodman](#), [João Alves](#), [Chris Beaumont](#), [Tom Dame](#), [James Jackson](#), [Jens Kauffmann](#),
[Thomas Robitaille](#), [Alberto Pepe](#), [Michelle Borkin](#), [Andreas Burkert](#), [Robert A Benjamin](#)

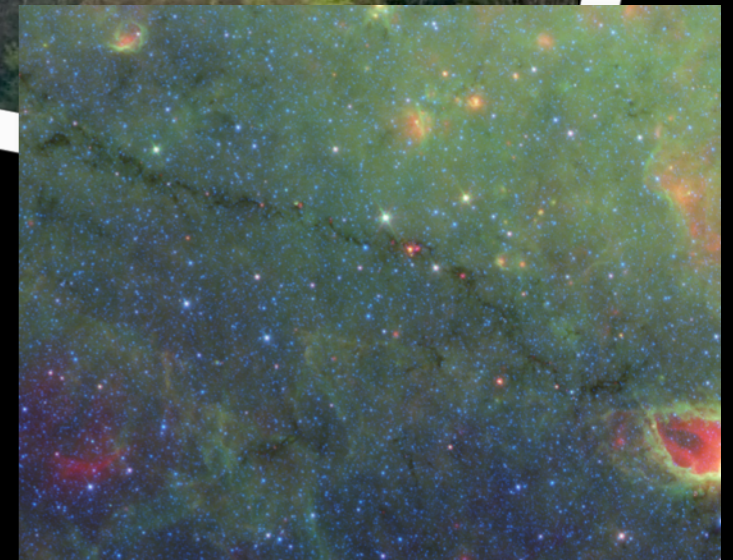
NOTES TO ONLINE READERS

This article was submitted to the *Astrophysical Journal* in December 2013.

The current online version is citable as an online "Authorea" preprint, and you can use the article's URL to do

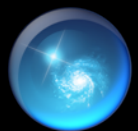
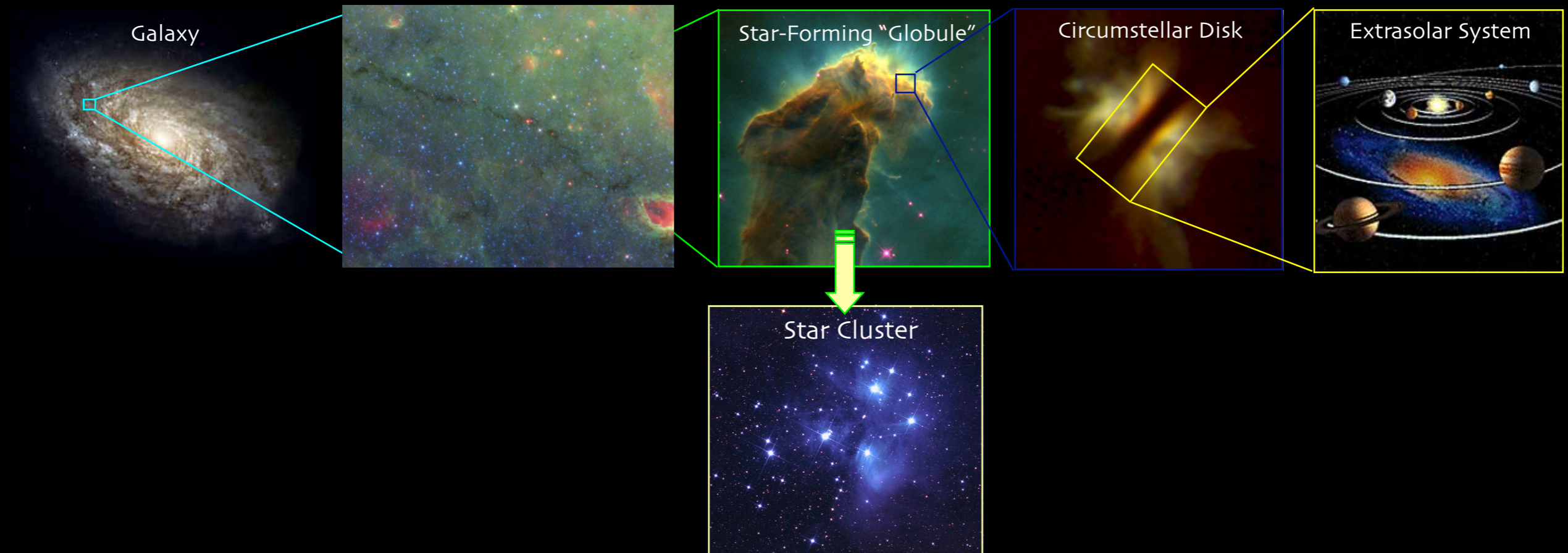
**Once upon a time (2012), in an
enchanted castle (in Bavaria)**

...at a conference about star formation





Star and Planet Formation

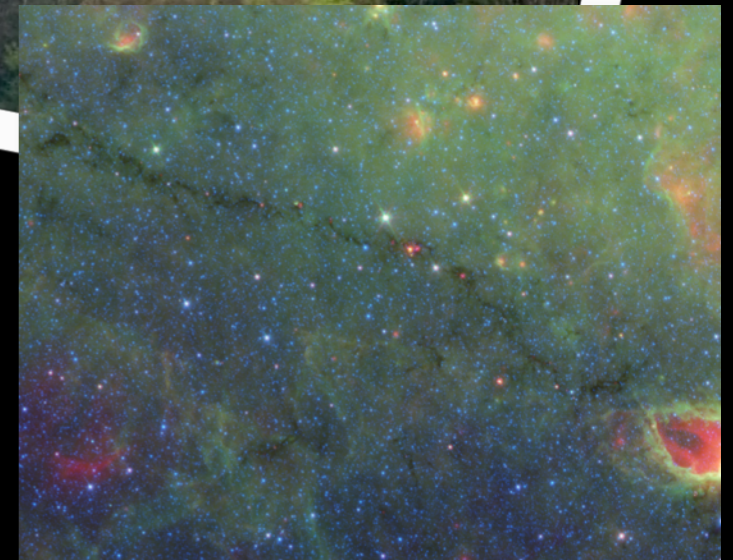


**Once upon a time (2012), in an
enchanted castle (in Bavaria)**

...at a conference about star formation

**Question *Andi Burkert*: Is Nessie
“parallel to the Galactic Plane”?**

**Answer *no one* immediately knew the
answer!**



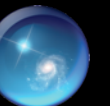
The Milky Way



"Galactic Plane"



The Milky Way
(Artist's Conception)



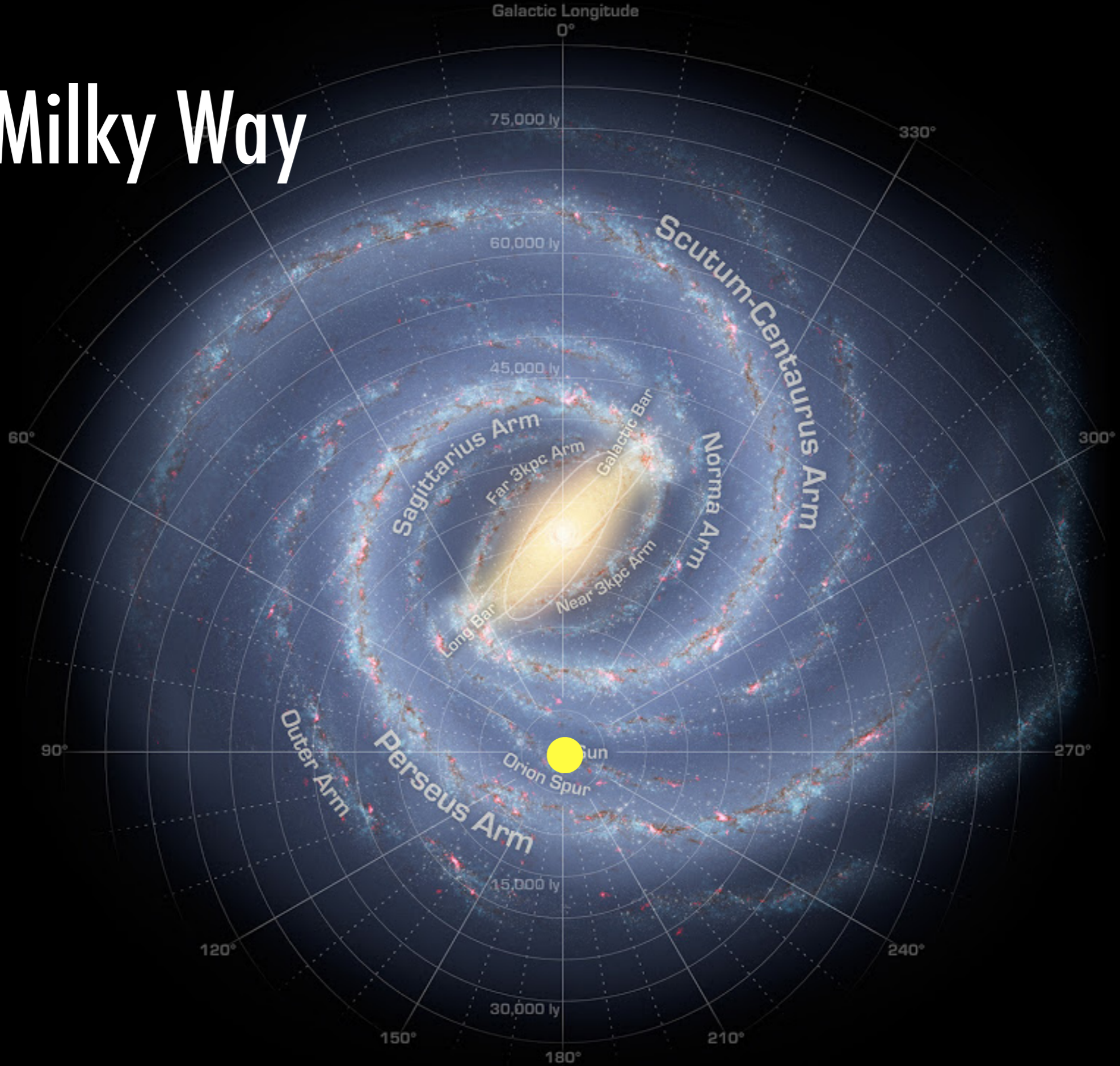
"Is Nessie Parallel to the Galactic Plane?"

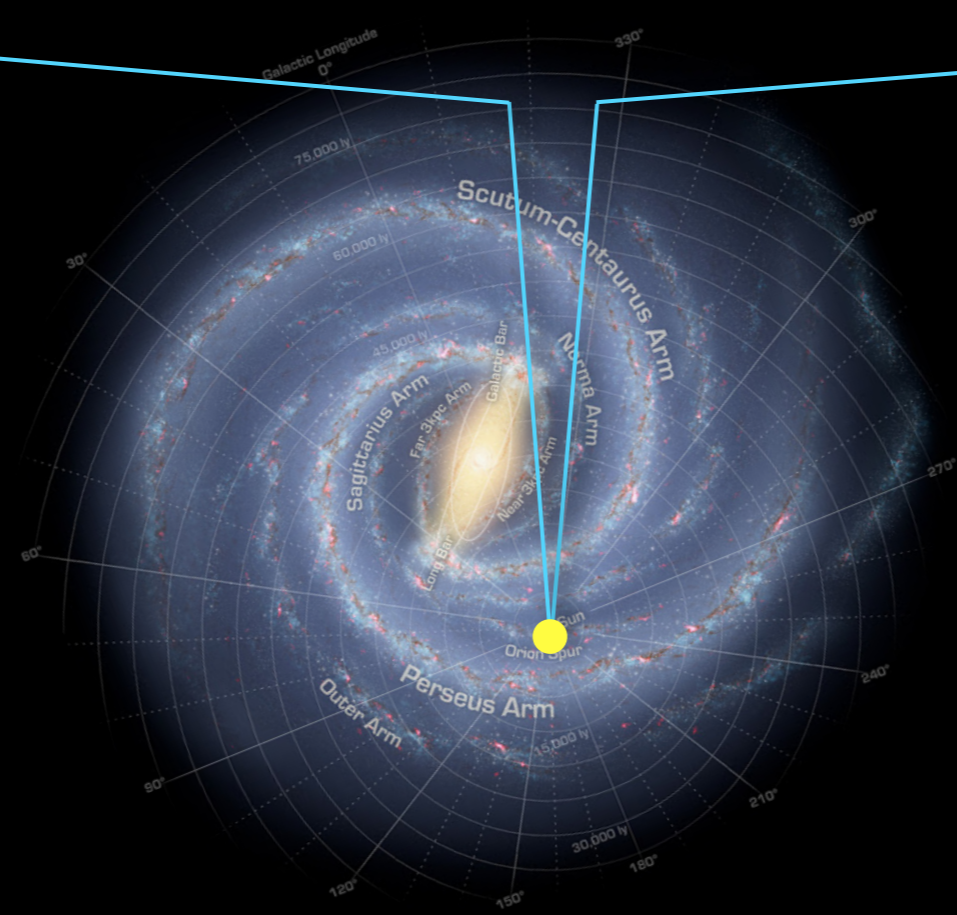
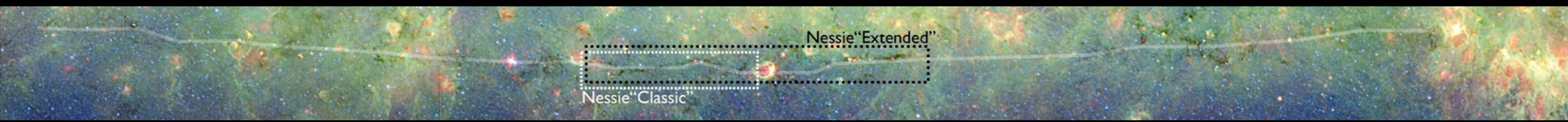
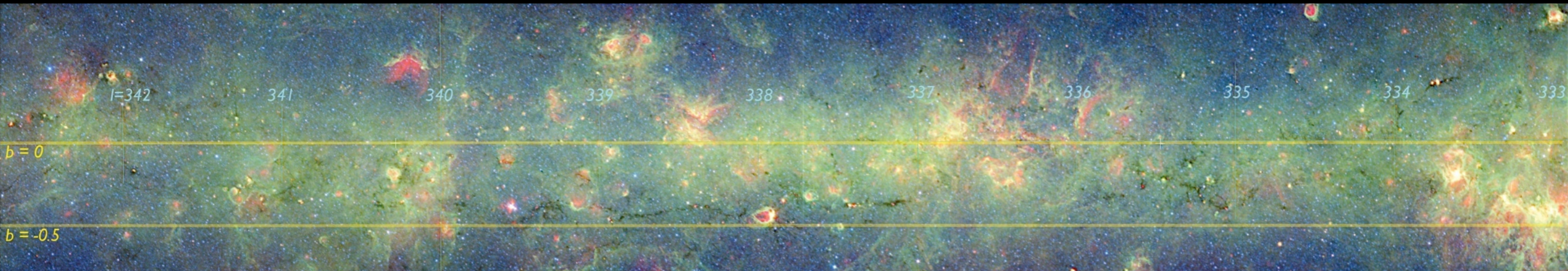


Yes, but why not at Zero of Latitude ($b=0$)?

The image shows a screenshot of the "GLIMPSE / MIPS GAL VIEWER" interface. At the top, there is a header with the text "GLIMPSE / MIPS GAL VIEWER" and three buttons: "LINK TO CURRENT VIEW", "TOGGLE PINS", and "QUESTIONS?". Below the header is a dark blue band representing the galactic plane, with several satellite icons. A green horizontal line is drawn across the image, labeled "b=0". Below this line is a large field of red and orange stars, with a yellow horizontal line labeled "b=-0.5 deg". A blue dinosaur is overlaid on the star field. At the bottom, there is a control panel with a "IRAC" filter selector, a "IRAC/MIPS" filter selector, and navigation buttons including zoom in (+), zoom out (-), and directional arrows. The copyright notice "©2008 Space Science Institute" is at the bottom left, and the URL "back to: alienearths.org/glimpse" is at the bottom right.

The Milky Way







"Nessie Extended"

~500 light years long & 1.5 light years thick

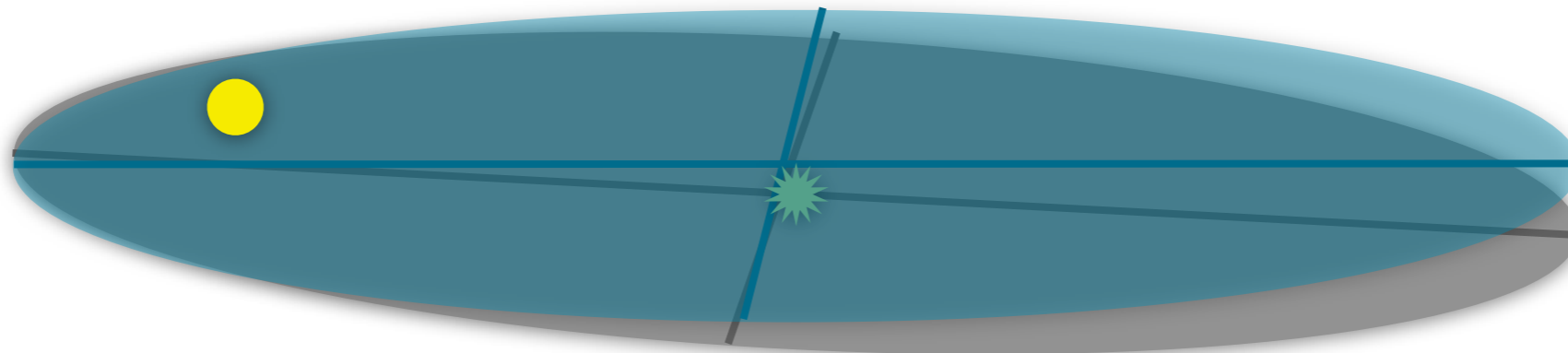
300:1 axial ratio

200,000 solar masses

BUT, why is it near $b=-0.5$, and not $b=0$?

Where are we, really?

“IAU Milky Way”, est. 1959



True Milky Way, modern

The equatorial plane of the new co-ordinate system must of necessity pass through the sun. It is a fortunate circumstance that, within the observational uncertainty, both the sun and Sagittarius A lie in the mean plane of the Galaxy as determined from the hydrogen observations. If the sun had not been so placed, points in the mean plane would not lie on the galactic equator. *[Blaauw et al. 1959]*

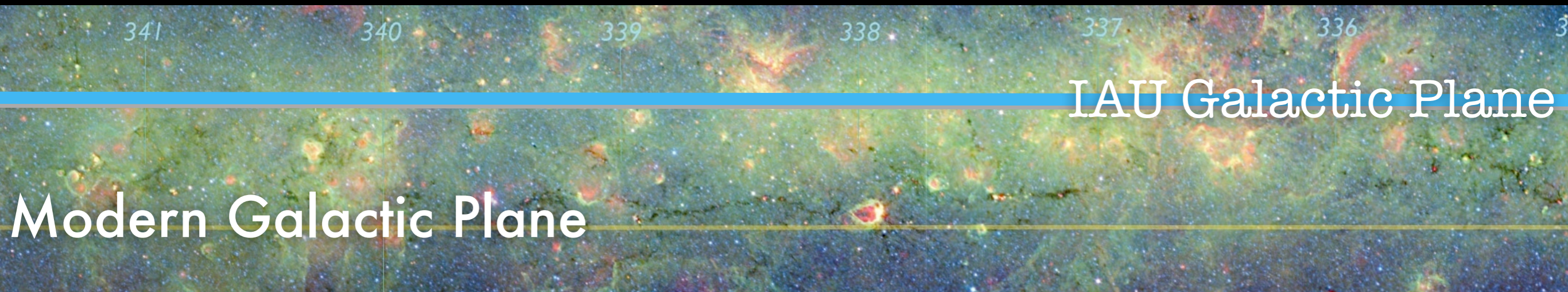
Sun is
~75 light years
“above” the
IAU Milky Way
Plane

+

Galactic
Center is
~20 light years
offset from the
IAU Milky Way
Center

=

The **Galactic Plane is not quite
where you’d think it is**
when you look at the sky



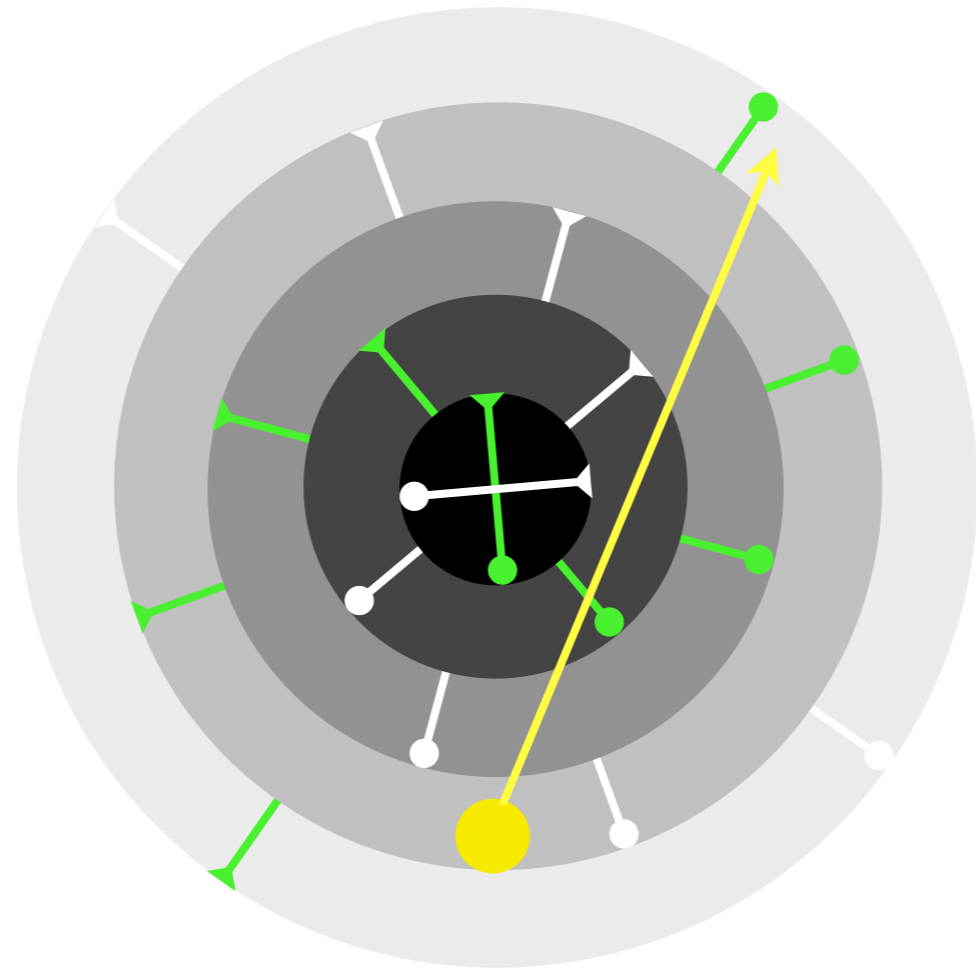
Yes, Nessie is EXACTLY in the Galactic Plane!

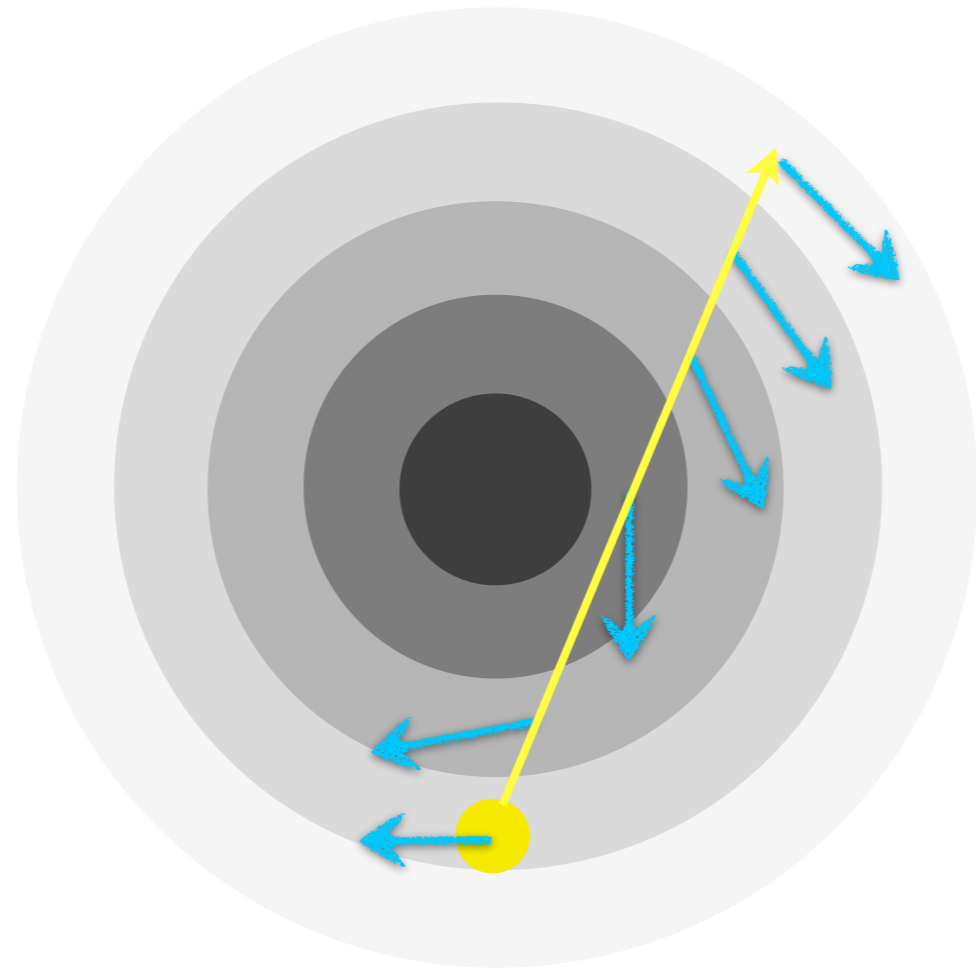
What about its distance?

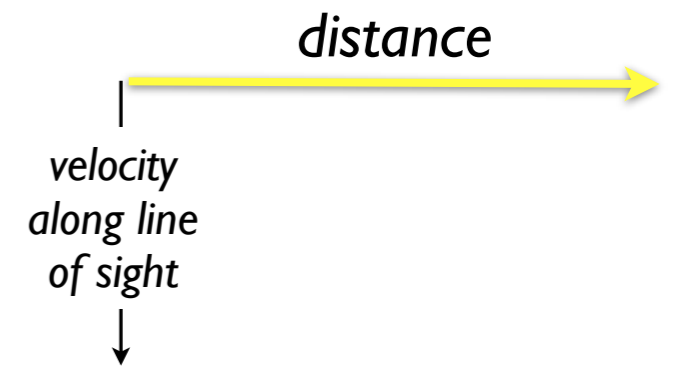
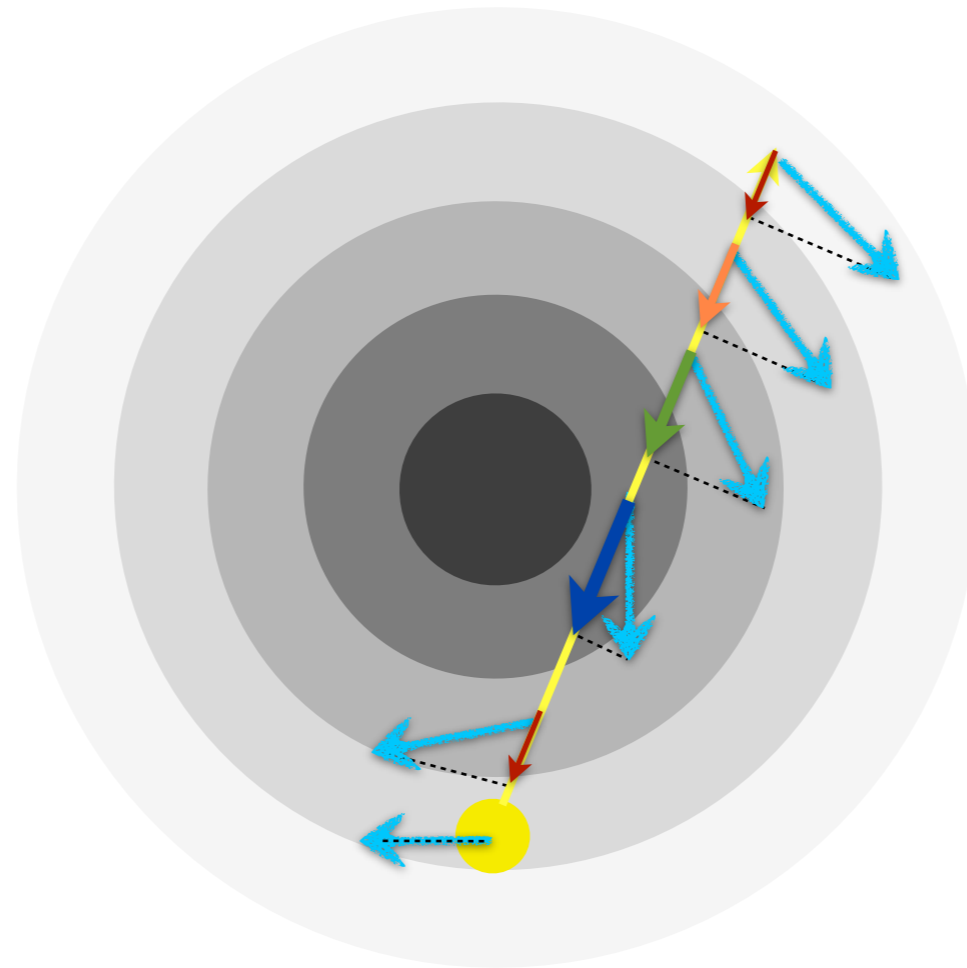
we can use “radial velocities” to estimate distance in a rotating galaxy..

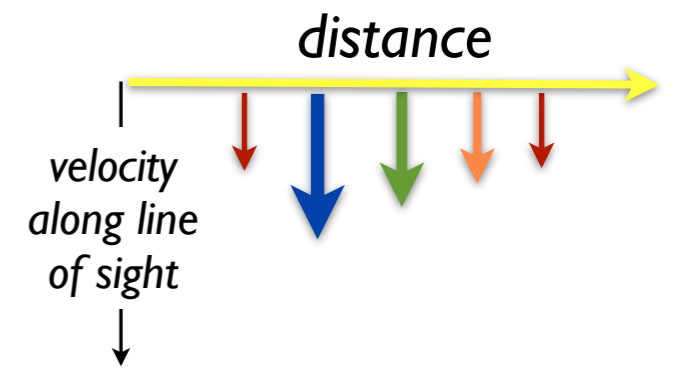
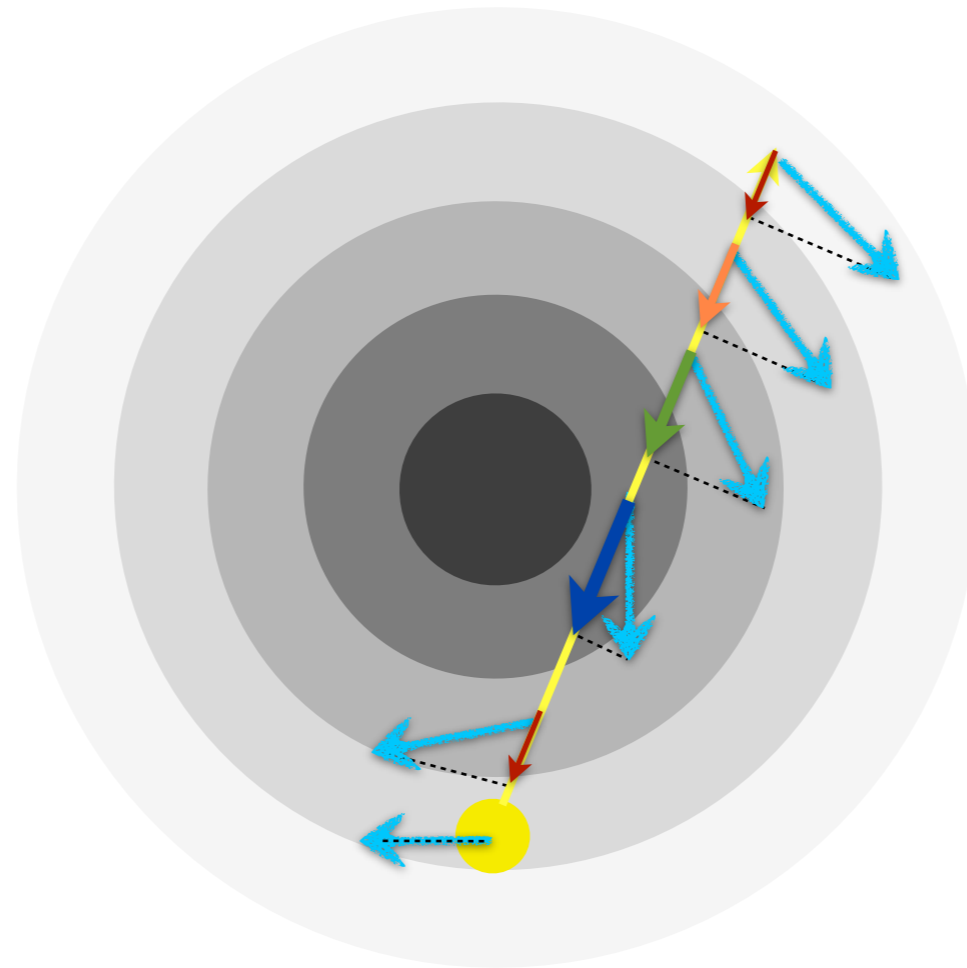
A Rotating (Spiral) Galaxy Observed from its Outskirts...



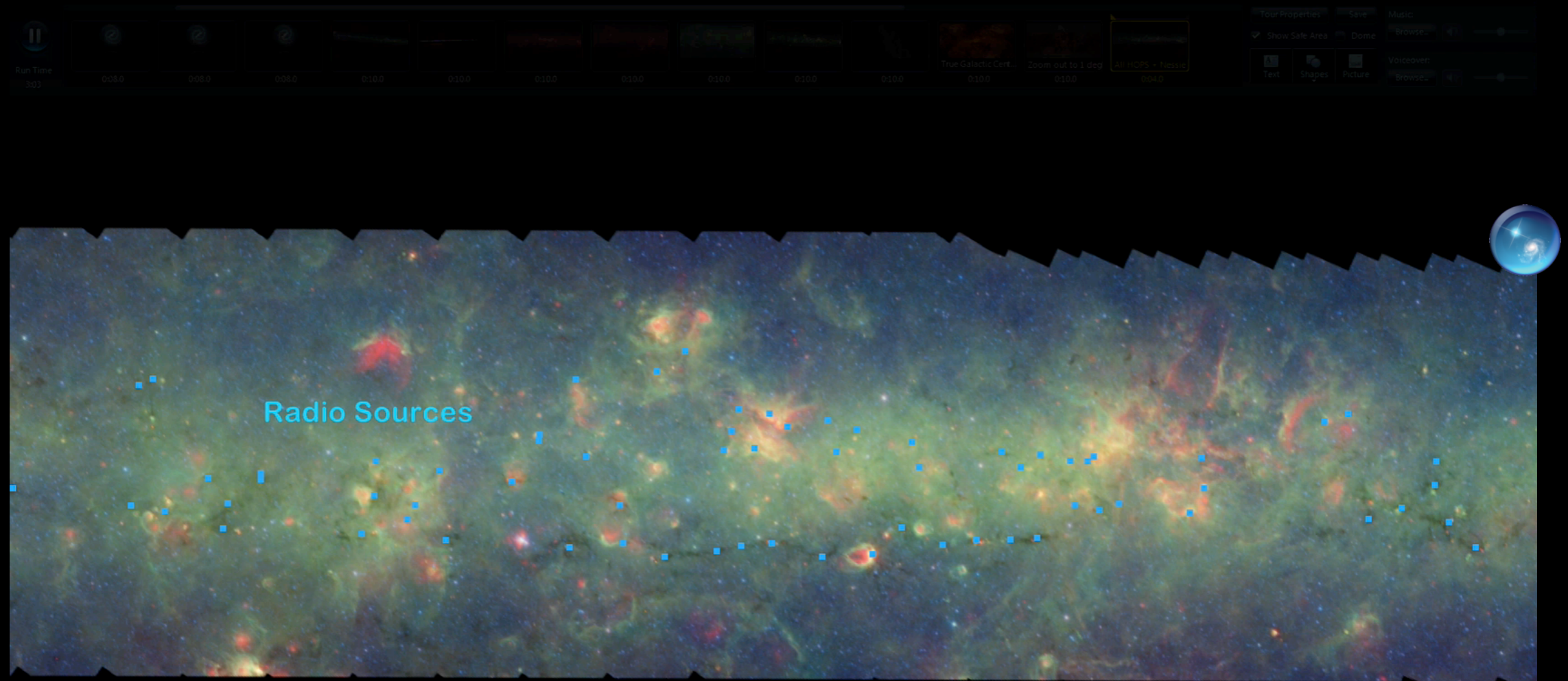




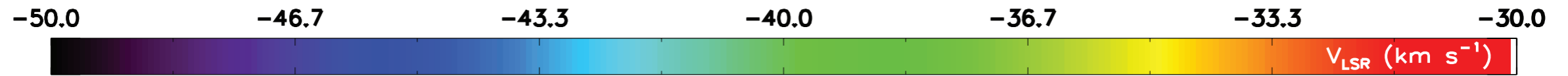




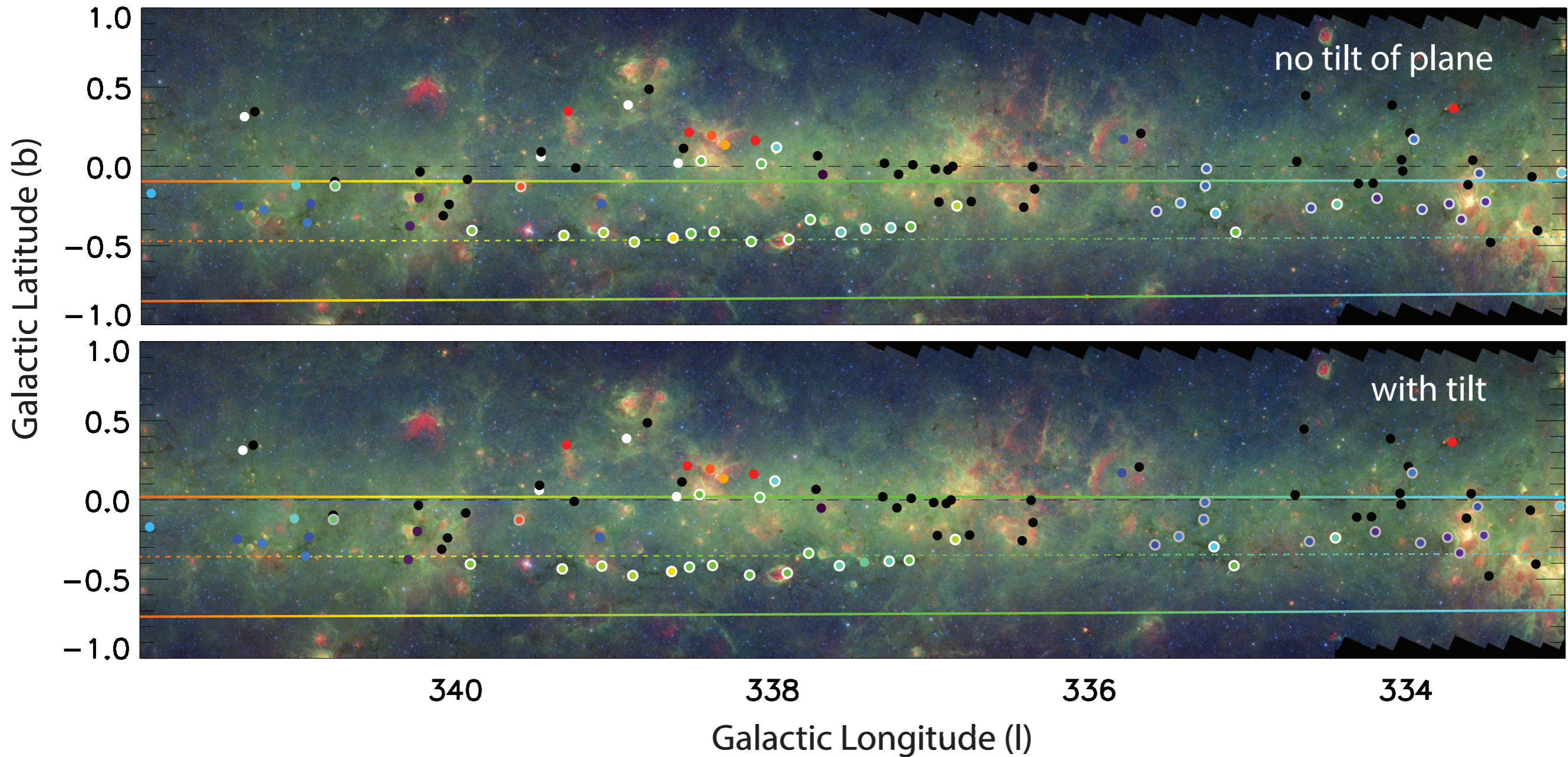
Velocity to Distance

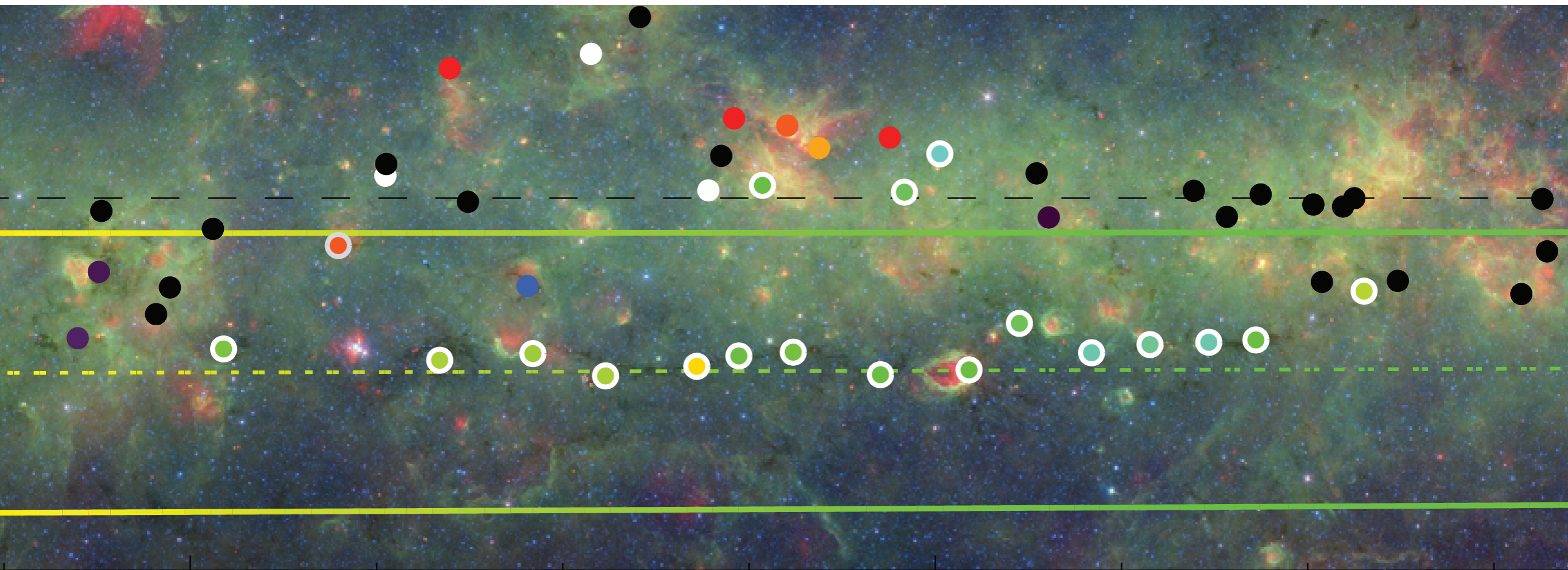


In the plane and at the distance of spiral arm!



$[Z_0=25.0 \text{ pc}, R_0=8.5 \text{ kpc}, \Theta_0=220 \text{ km/s}]$





...eerily precisely...

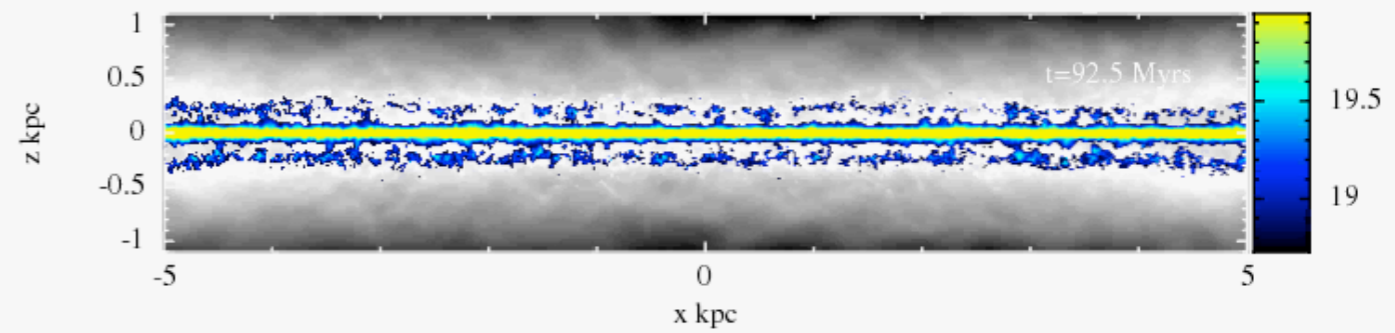
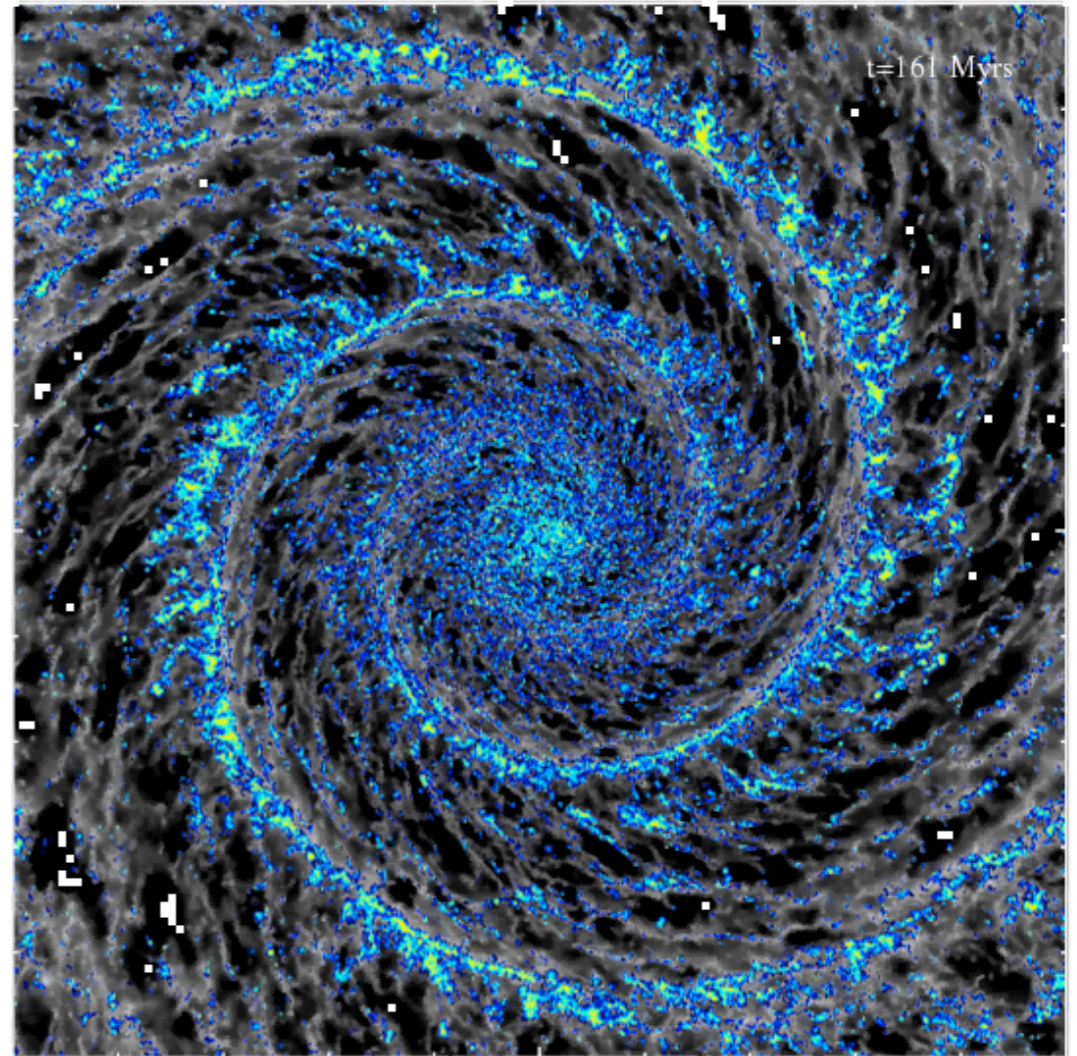
Monster to Bone

There could be 1000s more of these to find...a full skeleton perhaps?

A full 3D skeleton?



(flipped) image of IC342 from Jarrett et al. 2012; WISE Enhanced Resolution Galaxy Atlas



simulations courtesy Clare Dobbs

$z=0.00$

Formation of a Milky-Way-like Galaxy (Stars)



http://www.tapir.caltech.edu/~phopkins/Site/Movies_cosmo.html

Future of the Milky Way (Collision with Andromeda)



Sea Monster to Skeletal Shadow

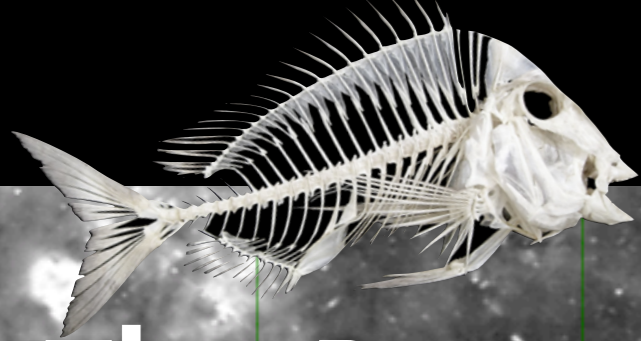


Spitzer GLIMPSE Image

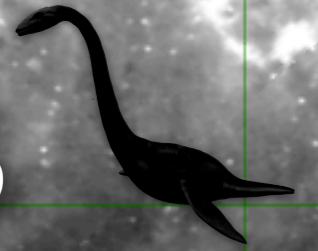
Peculiar dust cloud named "Nessie" much larger than thought.

Nessie more important as "bone" than sea monster.

Sun's height above Plane may make full Milky Way skeleton mappable.



The Bones of the Milky Way: Credits



Seamless Astronomy-style tools used in this project



authorea.com (open publishing)

theastrodata.org (open data)

glueviz.org (open source tools)

universe3d.org (collaborative data)

worldwidetelescope.org (universe information system)

[virtual observatory standards](#) (international online information-sharing systems)

Supported by



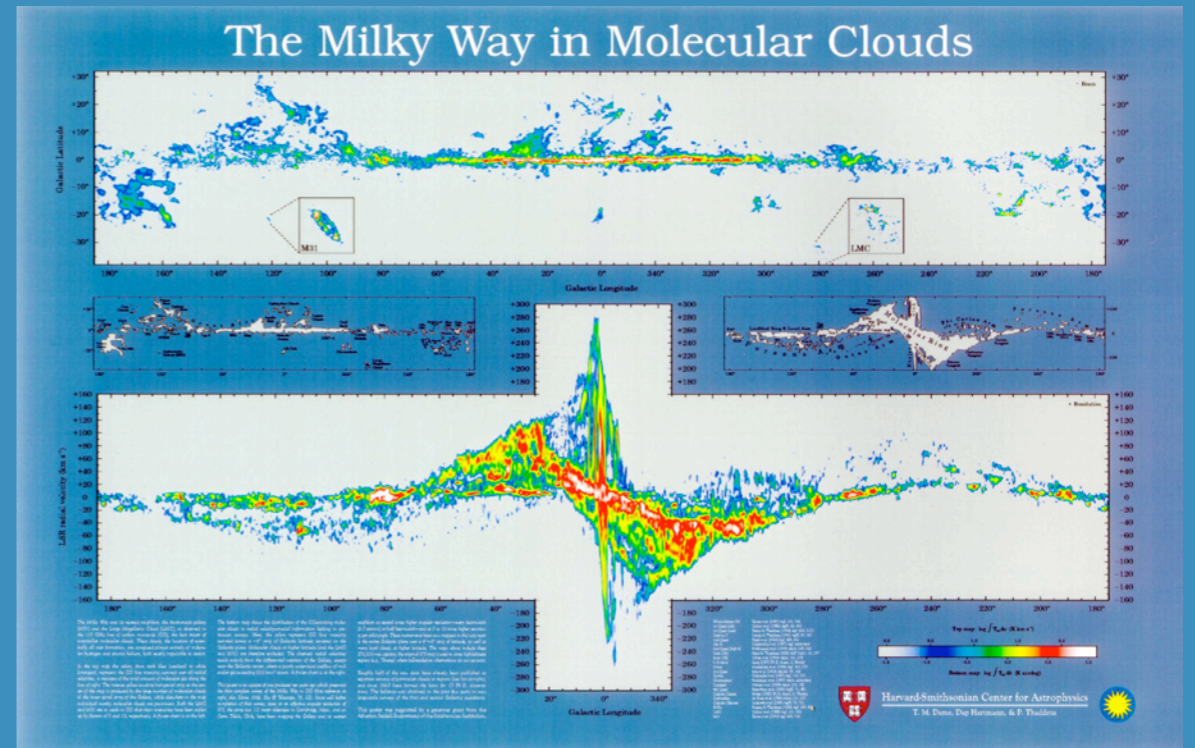
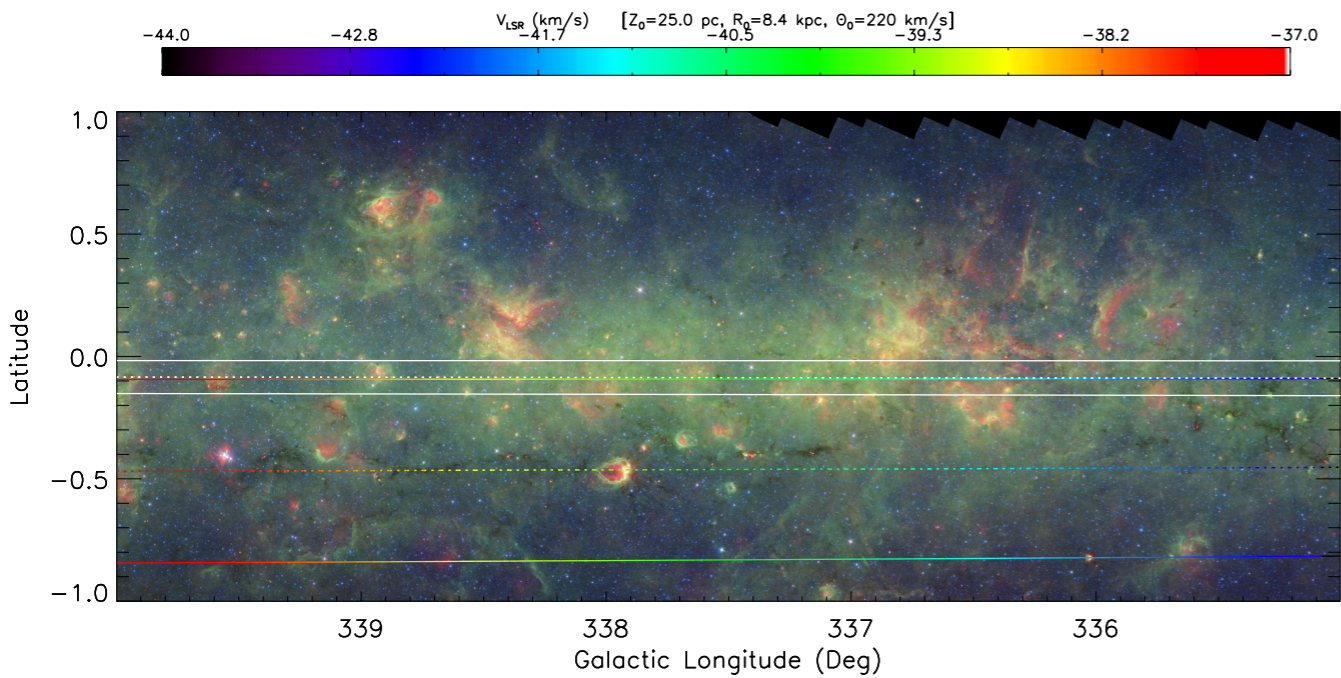
Alyssa Goodman milkywaybones.org

milkywaybones.org
worldwidetelescope.org

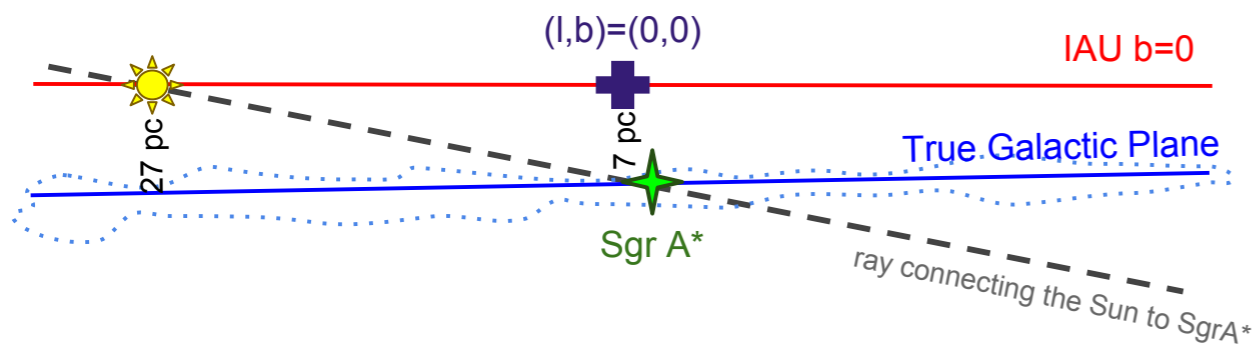
Where is "Nessie," in 3D?

How close to "in" the plane?

At what distance & inclination to l.o.s?



Drawing is schematic--NOT to scale



Notes:
 IAU b=0 set from HI, which is uncertain by ~0.1 degrees
 tilt of red w.r.t. blue would be $(20/8400) * 180/\pi = 0.13$ degrees

