# Seamless Astronomy & the WorldWide Telescope Alyssa A. Goodman

Harvard-Smithsonian Center for Astrophysics Initiative in Innovative Computing @ Harvard

#### Collaborators

Harvard-Smithsonian Center for Astrophysics & SEAS: Alberto Accomazzi, Eli Bressert, Douglas Burke, Rahul Davé, Pepi Fabbiano, Michael Kurtz, Gus Muench, Pavlos Protopapas Massachusetts General Hospital: Tim Clark & Sudeshna Das Microsoft Research: Jonathan Fay, Curtis Wong RPI: Jim Hendler & Deborah McGuinness STScl:Alberto Conti & Carol Christian UCLA: Christine Borgman

#### Seamless Astronomy

www.cfa.harvard.edu/~agoodman and worldwidetelescope.org



# What (the) "Virtual Observatory" meant/means/should mean to...

# Jim Gray & Alex Szalay Typical Astronomers Today

#### Me

## Astronomers who travel & use facebook...

## Astronomers & the V.O. c. 2006









#### ...requires: "Selection"; "Registration"; "Readable Labels"; "Highlighting"; & "Measurement"

in order to yield: "Inference": ... Wow, that's about 600 feet, hope we can change the room!





"Inference": ... Oh, that building with the funky paths outside is the Hyatt... what if I...





#### "GIS/Layering"



# What's needed?



# What's possible now?

**"Progressive Resolve"** "Zoom" "Search" "Selection" "GIS/Layering" "Registration"

"Side-by-Side **Comparison**" "Readable Labels" "Highlighting" "Custom Site" "Inference"

#### "Off-the-Desktop"

"Ontology"

"Measurement"

#### WorldWide Telescope



## WWT Demos...

Settings

IGC 281 Star-F

# \* "Dust & Us" (Sample Interactive Tour) \* Explore, Search, etc. (Basic Functionality) \* ADS/Wikipedia/SIMBAD (Deep Links)

Context Search



Digitized Sky Survey (Optical...

## Galileo's Moons of Jupiter

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SIDE ... EUSNUNCIUS

On the third, at the seventh hour, the stars were arranged in this quence. The eastern one was 1 minute, 30 seconds from Jupiter 2 closest western one 2 minutes; and the other western one wa

\* Wes

o minutes removed from this one. They were absolutely on the ame straight line and of equal magnitude.

On the fourth, at the second hour, there were four stars aroun upiter, two to the east and two to the west, and arranged precise

\* \* • • \* \* Wes

on a straight line, as in the adjoining figure. The easternmost wa listant 3 minutes from the next one, while this one was 40 second from Jupiter; Jupiter was 4 minutes from the nearest western one. d this one 6 minutes from the westernmost one. Their magnitude, ere nearly equal; the one closest to Jupiter appeared a little smaller ian the rest. But at the seventh hour the eastern stars were only o seconds apart. Jupiter was 2 minutes from the nearer eastern

\*\* O \* \* West

me, while he was 4 minutes from the next western one, and this one was 3 minutes from the westernmost one. They were all equal and extended on the same straight line along the ecliptic.

On the fifth, the sky was cloudy.

On the sixth, only two stars appeared flanking Jupiter, as is seen

jegelsee verd staats

fast

East

West

in the adjoining figure. The eastern one was 2 minutes and the vestern one 3 minutes from Jupiter. They were on the same straight fine with Jupiter and equal in magnitude.

On the seventh, two stars stood near Jupiter, both to the east.

Notes for & re-productions of Siderius Nuncius (1610)



The Spitzer image was obtained with the infrared array camera. Emission at 3.6 microns is shown as blue, emission from 4.5 and 5.8 microns has been combined as green, and 8.0 micron emission is depicted as red.

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# Seamless Astronomy



Mockup based on work of Eli Bressert, excerpted from NASA AISRP proposal by Goodman, Muench, Christian, Conti, Kurtz, Burke, Accomazzi, McGuinness, Hendler & Wong, 2008

# astrobitz?



# And to go fully "seamless"?

- "Progressive Resolve"
- "Zoom"
- "Search"
- "Selection"
- "GIS/Layering"
- "Registration"

"Side-by-Side Comparison" "Readable Labels" "Highlighting" "Custom Site" "Inference"

#### "Off-the-Desktop"

"Ontology" "Measurement"

## Going **"Off-the-Desktop"**



More information: See the SEAS/IIC "Scientists Discovery Room" project pages

#### Slideshow: Tabletop Computers Continued By Meredith Ringel Morris

#### First Published December 2008

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**UBITABLE**: Users can interact with surface computers through auxiliary devices, such as laptops, phones, and PDAs. The display on the auxiliary device can convey private or sensitive content to a single user, while group-appropriate content can appear on the tabletop display. Chia Shen and her colleagues at Mitsubishi Electric Research Laboratories, in Cambridge, Mass., have explored auxiliary interactions with surface computers in their UbiTable project, in which two people with laptops collaborate over a tabletop display. Recently, Shen expanded the UbiTable into an interactive room called the WeSpace. People can share data on their laptops with other people in the room, using both a table and a large display wall. Here, three Harvard University astrophysicists discuss radio and IR spectrum images using the WeSpace.

"Ontology"

"Measurement"

# Seamless Astronomy



Mockup based on work of Eli Bressert, excerpted from NASA AISRP proposal by Goodman, Muench, Christian, Conti, Kurtz, Burke, Accomazzi, McGuinness, Hendler & Wong, 2008



# Seamless "Measurement" via SAMP?



Mockup based on work of Eli Bressert, excerpted from NASA AISRP proposal by Goodman, Muench, Christian, Conti, Kurtz, Burke, Accomazzi, McGuinness, Hendler & Wong, 2008

## Our "Real" Work... COMPLETE, Astronomical Medicine & Taste-Testing

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Monday, November 16, 2009

Done

## Star (and Planet, and Moon) Formation 101



#### MARCE STREETE = COordinated Molecular Line Exinction Thermal Emission

#### mm peak (Enoch et al. 2006)

sub-mm peak (Hatchell et al. 2005, Kirk et al. 2006)

<sup>13</sup>CO (Ridge et al. 2006)

mid-IR IRAC composite from c2d data (Foster, Laakso, Ridge, et al. in prep.)

Optical image (Barnard 1927)

oom: 227% Angle: 0

# Astronomical Medicine

Alyssa Goodman (IIC/CfA/FAS) Michael Halle (IIC/SPL/HMS) Ron Kikinis (SPL/HMS) Douglas Alan (IIC) Michelle Borkin (FAS/IIC) Jens Kauffmann (CfA/IIC) Erik Rosolowsky (CfA/UBC Okanagan) Nick Holliman (U. Durham)





3D Viz made with VolView

# AstronomicalMedicine@



#### Some of What We're Learning...

WU 73 WHE 132



#### Cores nest in coccoons (Kauffmann et al. 2009)





#### Tripled Outflows (Borkin et al. 2008,9)



(Arce et al. 2009)

## Tasting Star Formation



Star Formation Taste Tests > Overview									
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Back to Dashboard   Switch to a different project   Log-out   Hat2									
Star Formation Taste Tests CFA									
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Project overview & activity       New message   New to-do list   New milestone   New file       III         Welcome to the Tasting Room       IIIIIIIIIIIIIIIIIIIIIIIIIIIIIIIIIIII									
			This project's RSS feed Subscribe to your project RSS feed and be notified when someone posts a message, comment or file, or adds or completes a to-do item or milestone in this project. What's RSS? People on this project HCO						

This is the collaborative space for those who do simulations of star forming regions, and those who observe them. It was inspired, in the Fall of 2006, by the NSF proposal entitled "Star Formation Taste Tests," by A. Goodman & E. Rosolowsky. Today, it is used to host conversations about and short descriptions of simulatons, along with links to longer descriptions (e.g. Journal articles & web sites). In the future, we are planning to connect more enhanced descriptions of those simulations directly to online code bases and sample outputs, via the new <u>CADAC</u> site. So, stay tuned.

#### MONDAY, 13 APRIL 2009

Message	Relevant References relating to Bayesian Methods	Posted by	Rahul S.					
JESDAY, 7 APRIL 2009								
File	dustfit_slides.pdf	Uploaded by	Rahul S.					
EDNESDAY, 1	8 FEBRUARY 2009							
Writeboard	Taste Tests we Plan (COMPLETE Group)	Updated by	Alyssa G.					
To-do	Compare PPP and PPV dendrograms to determine the correct "paradigm" for mapping between the two. (Dendrograms and Simulations)	Completed by	Alyssa G.					
To-do	Taste Testing delivery to CADAC prior to Ringberg Meeting (Dendrograms and Simulations)	Completed by	Alyssa G.					
To-do	link to http://www1.astrophysik.uni-kiel.de/asd/ (Dendrograms and Simulations)	Assigned to	Sarah B.					
Writeboard	Re: Heitsch et al: Colliding Flows	Comment by	Alyssa G.					
EDNESDAY, 21 JANUARY 2009								
Message	Decadal Survey	Posted by	Alyssa G.					
URSDAY, 20 NOVEMBER 2008								
Comment	Re: "Toward a Prescriptive Understanding of Kennicutt- Schmidt Relations"	Posted by	Alex L.					
Comment	Re: "Toward a Prescriptive Understanding of Kennicutt- Schmidt Relations"	Posted by	Alex L.					

Alyssa Goodman You are logged in right now

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Joao Alves Hasn't logged in recently

Caltech

Scott Schnee Hasn't logged in recently

## Magnetohydrodynamic Simulations



Simulations of Bate 2009



# The Taste-Testing Process



## (Tasting) The Role Self-Gravity in Star Formation

#### LETTERS

#### NATURE Vol 457 1 January 2009



Figure 2 | Comparison of the 'dendrogram' and 'CLUMPFIND' featureidentification algorithms as applied to <sup>13</sup>CO emission from the L1448 region of Perseus. a, 3D visualization of the surfaces indicated by colours in the dendrogram shown in c. Purple illustrates the smallest scale selfgravitating structures in the region corresponding to the leaves of the dendrogram; pink shows the smallest surfaces that contain distinct selfgravitating leaves within them; and green corresponds to the surface in the data cube containing all the significant emission. Dendrogram branches corresponding to self-gravitating objects have been highlighted in yellow over the range of  $T_{mb}$  (main-beam temperature) test-level values for which the virial parameter is less than 2. The x-y locations of the four 'selfgravitating' leaves labelled with billiard balls are the same as those shown in Fig. 1. The 3D visualizations show position–position–velocity (p–p–y) space. RA, right ascension; dec., declination. For comparison with the ability of dendrograms (c) to track hierarchical structure, d shows a pseudo-

dendrogram of the CLUMPFIND segmentation (**b**), with the same four labels used in Fig. 1 and in **a**. As 'clumps' are not allowed to belong to larger structures, each pseudo-branch in **d** is simply a series of lines connecting the maximum emission value in each clump to the threshold value. A very large number of clumps appears in **b** because of the sensitivity of CLUMPFIND to noise and small-scale structure in the data. In the online PDF version, the 3D cubes (**a** and **b**) can be rotated to any orientation, and surfaces can be turned on and off (interaction requires Adobe Acrobat version 7.0.8 or higher). In the printed version, the front face of each 3D cube (the 'home' view in the interactive online version) corresponds exactly to the patch of sky shown in Fig. 1, and velocity with respect to the Local Standard of Rest increases from front ( $-0.5 \text{ km s}^{-1}$ ) to back (8 km s<sup>-1</sup>).

data, CLUMPFIND typically finds features on a limited range of scales, above but close to the physical resolution of the data, and its results can be overly dependent on input parameters. By tuning CLUMPFIND's two free parameters, the same molecular-line data set<sup>8</sup> can be used to show either that the frequency distribution of clump mass is the same as the initial mass function of stars or that it follows the much shallower mass function associated with large-scale molecular clouds (Supplementary Fig. 1).

Four years before the advent of CLUMPFIND, 'structure trees'<sup>9</sup> were proposed as a way to characterize clouds' hierarchical structure

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using 2D maps of column density. With this early 2D work as inspiration, we have developed a structure-identification algorithm that abstracts the hierarchical structure of a 3D (p-p-v) data cube into an easily visualized representation called a 'dendrogram'<sup>10</sup>. Although well developed in other data-intensive fields<sup>11,12</sup>, it is curious that the application of tree methodologies so far in astrophysics has been rare, and almost exclusively within the area of galaxy evolution, where 'merger trees' are being used with increasing frequency<sup>13</sup>.

Figure 3 and its legend explain the construction of dendrograms schematically. The dendrogram quantifies how and where local maxima of emission merge with each other, and its implementation is explained in Supplementary Methods. Critically, the dendrogram is determined almost entirely by the data itself, and it has negligible sensitivity to algorithm parameters. To make graphical presentation possible on paper and 2D screens, we 'flatten' the dendrograms of 3D data (see Fig. 3 and its legend), by sorting their 'branches' to not cross, which eliminates dimensional information on the x axis while preserving all information about connectivity and hierarchy. Numbered 'billiard ball' labels in the figures let the reader match features between a 2D map (Fig. 1), an interactive 3D map (Fig. 2a online) and a sorted dendrogram (Fig. 2c).

A dendrogram of a spectral-line data cube allows for the estimation of key physical properties associated with volumes bounded by isosurfaces, such as radius (R), velocity dispersion ( $\sigma_{\nu}$ ) and luminosity (L). The volumes can have any shape, and in other work<sup>14</sup> we focus on the significance of the especially elongated features seen in L1448 (Fig. 2a). The luminosity is an approximate proxy for mass, such that  $M_{\text{lum}} = X_{13\text{CO}}L_{13\text{CO}}$ , where  $X_{13\text{CO}} = 8.0 \times 10^{20} \text{ cm}^2 \text{ K}^{-1} \text{ km}^{-1} \text{ s}$ (ref. 15; see Supplementary Methods and Supplementary Fig. 2). The derived values for size, mass and velocity dispersion can then be used to estimate the role of self-gravity at each point in the hierarchy, via calculation of an 'observed' virial parameter,  $\alpha_{obs} = 5\sigma_{o}^{2}R/GM_{back}$ In principle, extended portions of the tree (Fig. 2, yellow highlighting) where  $\alpha_{obs} < 2$  (where gravitational energy is comparable to or larger than kinetic energy) correspond to regions of p-p-v space where selfgravity is significant. As  $\alpha_{obs}$  only represents the ratio of kinetic energy to gravitational energy at one point in time, and does not explicitly capture external over-pressure and/or magnetic fields<sup>16</sup>, its measured value should only be used as a guide to the longevity (boundedness) of any particular feature.



Figure 3 | Schematic illustration of the dendrogram process. Shown is the construction of a dendrogram from a hypothetical one-dimensional emission profile (black). The dendrogram (blue) can be constructed by 'dropping' a test constant emission level (purple) from above in tiny steps (exaggerated in size here, light lines) until all the local maxima and mergers are found, and connected as shown. The intersection of a test level with the emission is a set of points (for example the light purple dots) in one dimensions, a planar curve in two dimensions, and an isosurface in three dimensions. The dendrogram of 3D data shown in Fig. 2c is the direct analogue of the tree shown here, only constructed from 'isosurface' rather than 'point' intersections. It has been sorted and flattened for representation on a flat page, as fully representing dendrograms for 3D data cubes would require four dimensions.

PDF

Goodman et al. Nature, 2009



# Seamless Astronomy



Mockup based on work of Eli Bressert, excerpted from NASA AISRP proposal by Goodman, Muench, Christian, Conti, Kurtz, Burke, Accomazzi, McGuinness, Hendler & Wong, 2008

#### BONUS DEMO

#### Ongoing(!) "Real" Work... COMPLETE, Astronomical Medicine & Taste-Testing

The Role of B-Stars in Molecular Cloud Evolution

(special thanks to Sanjana Sharma & Gus Muench!)