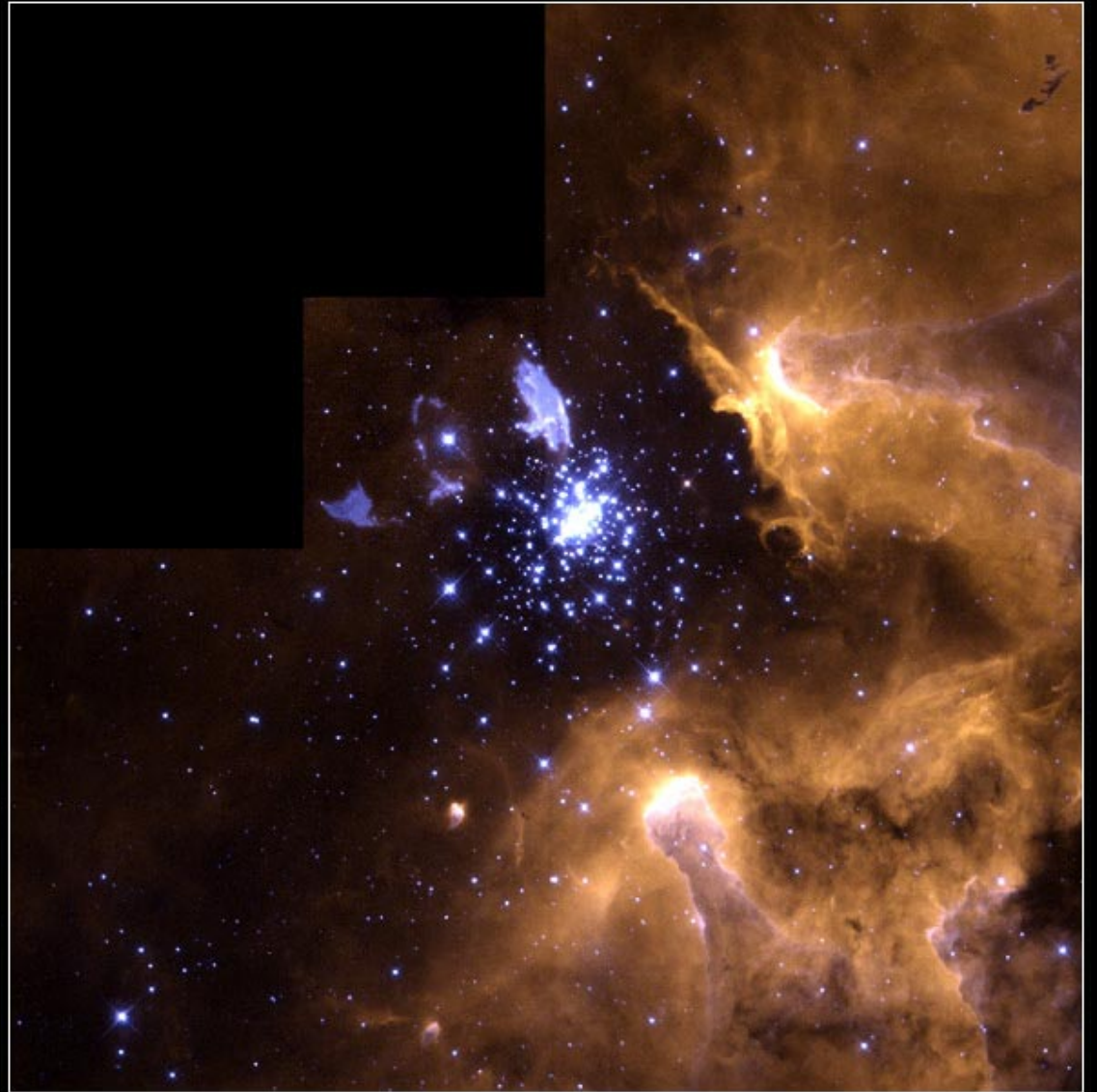


Recycling in the Universe

Alyssa A. Goodman

Department of Astronomy
Harvard University



NGC 3603

HST • WFPC2

PRC99-20 • STScI OPO • June 1, 1999

Wolfgang Brandner (JPL/IPAC), Eva K. Grebel (Univ. Washington),
You-Hua Chu (Univ. Illinois, Urbana-Champaign) and NASA

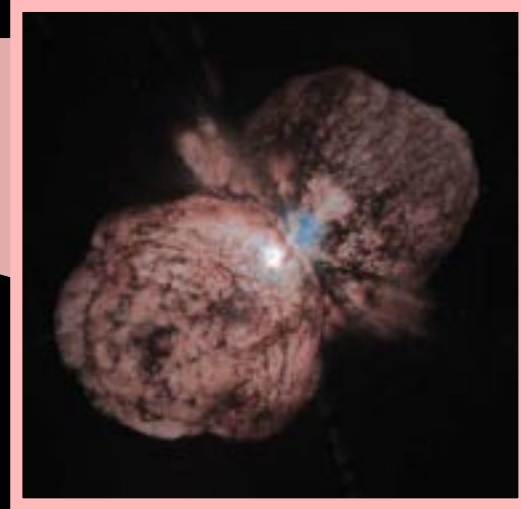
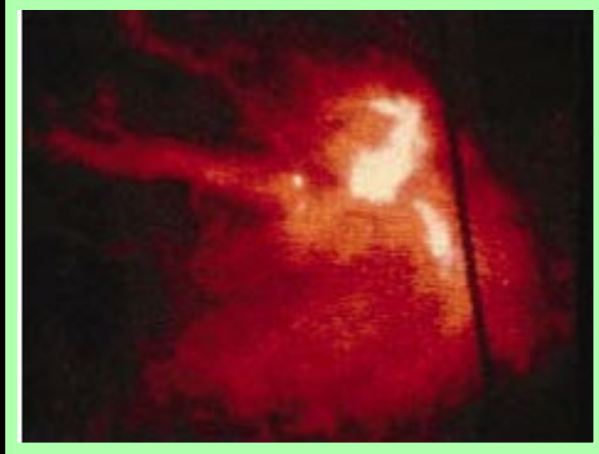


Recycling on Earth & In Galaxies

	On Earth	In Galaxies
Storage	Neatly, in a "Recycling" Bin	Not as neatly, in the Interstellar Medium
Collection	Big Trucks	Gravity & Supernova "Snowplows"
Processing	Recycling Plant	Molecular Clouds
Production	Factories	Star-forming Cores in Molecular Clouds
Consumption	Humans	Stars
Discarding	Human Tosses	Stellar Winds
Efficiency of One Cycle	Pretty Low, Maybe 10%	Pretty High, Maybe 90%
Timescale for One Cycle	Weeks to Years	Millions to Billions of Years



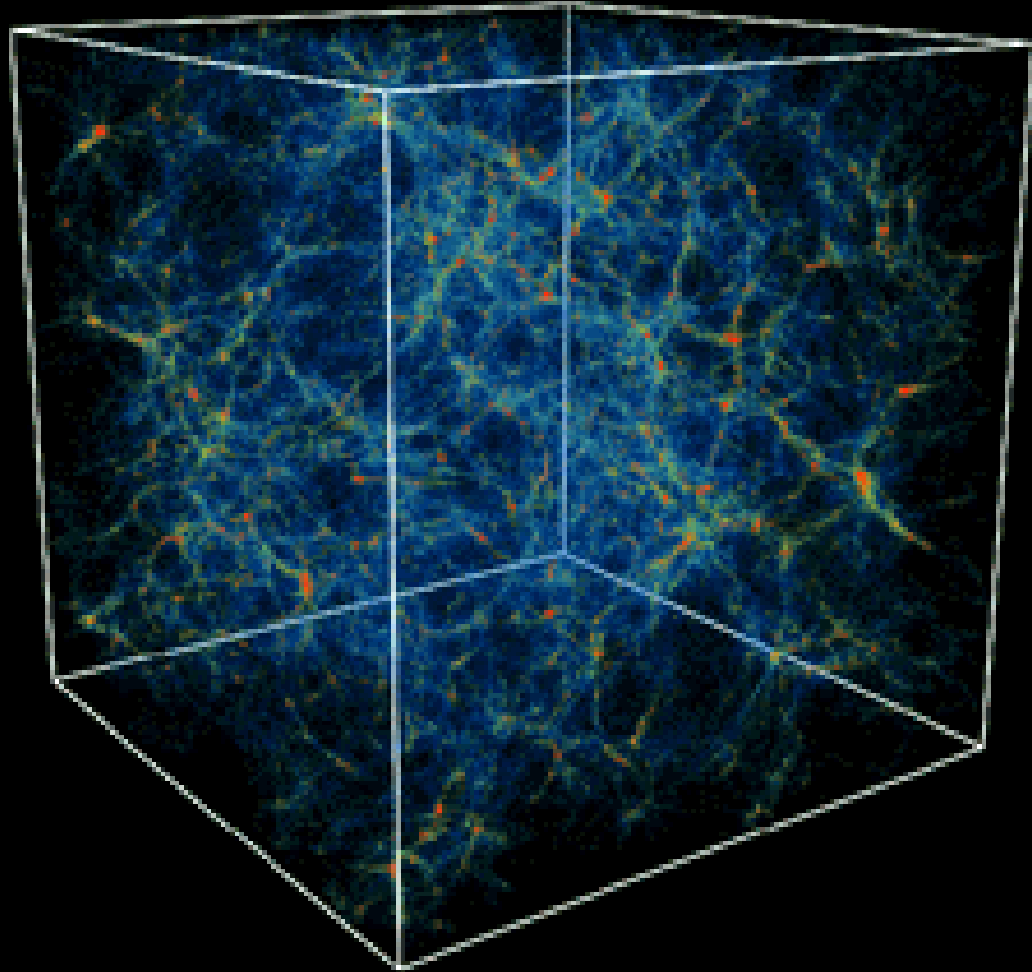
Recycling in the Universe



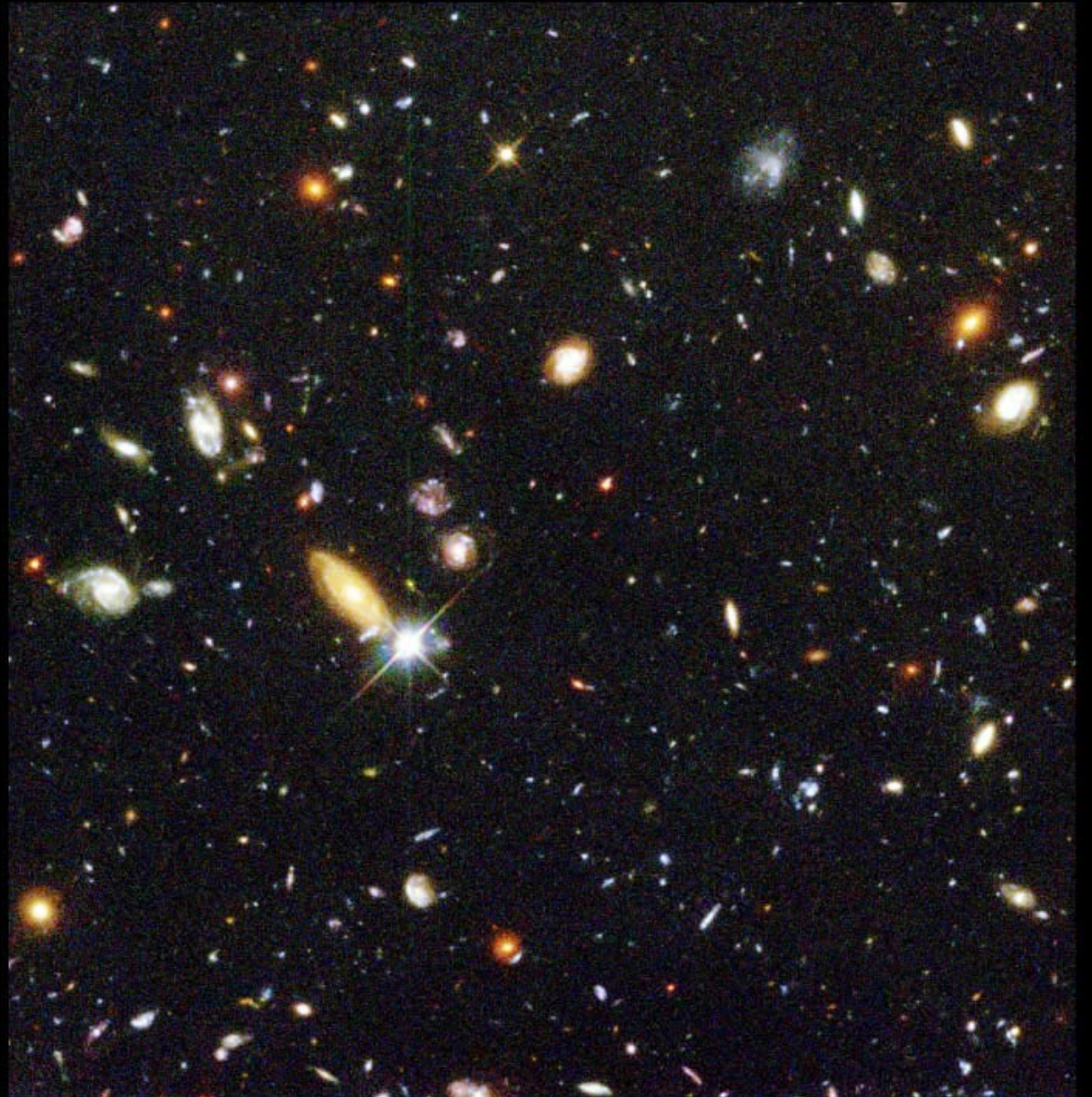
Making the First Recyclables...

Fluctuations about 300,000 years after the Big Bang lead to "**Structure Formation.**"

Gravitational collapse of some of these "structures" produces the **first stars and galaxies.**



Pretty
young
galaxies

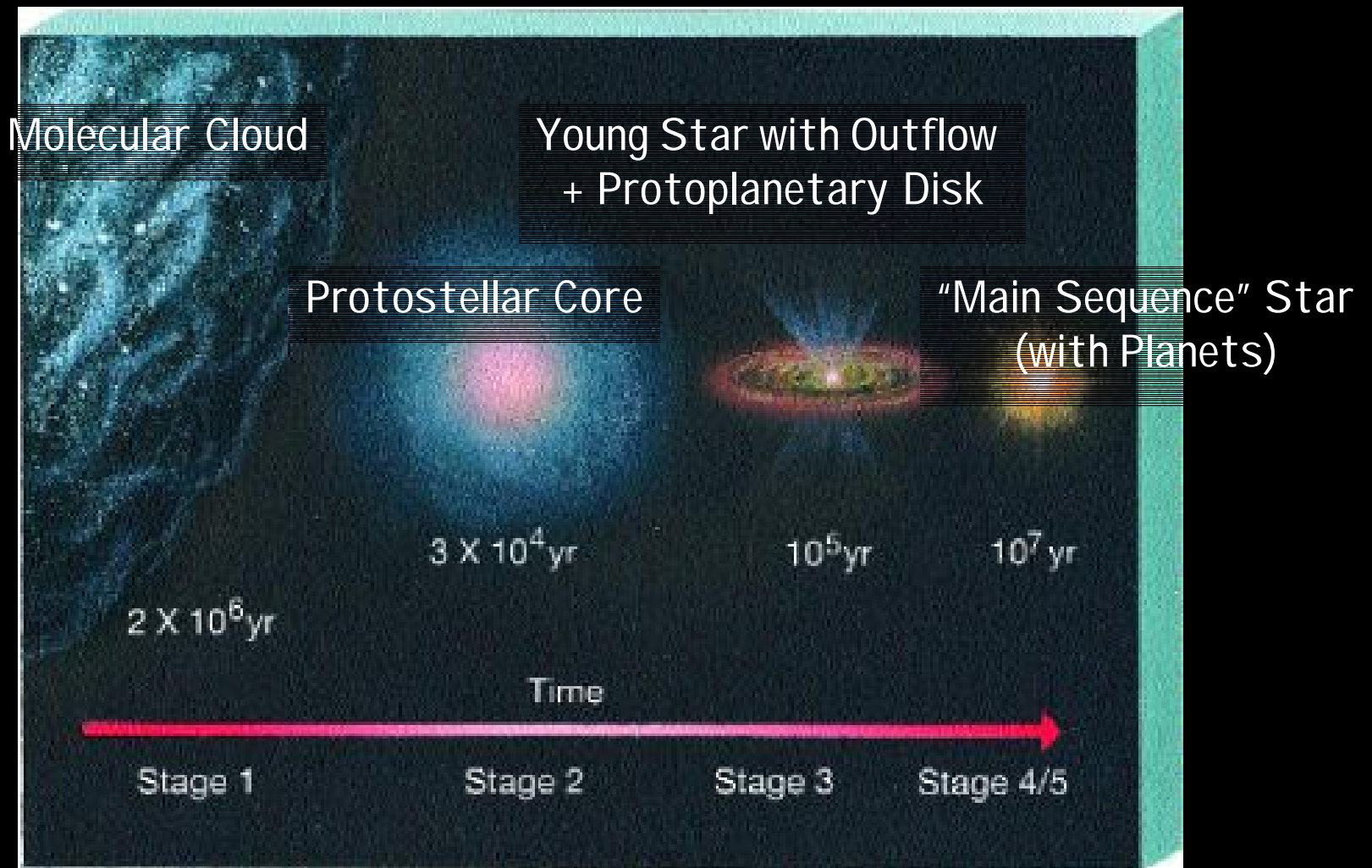


Hubble Deep Field

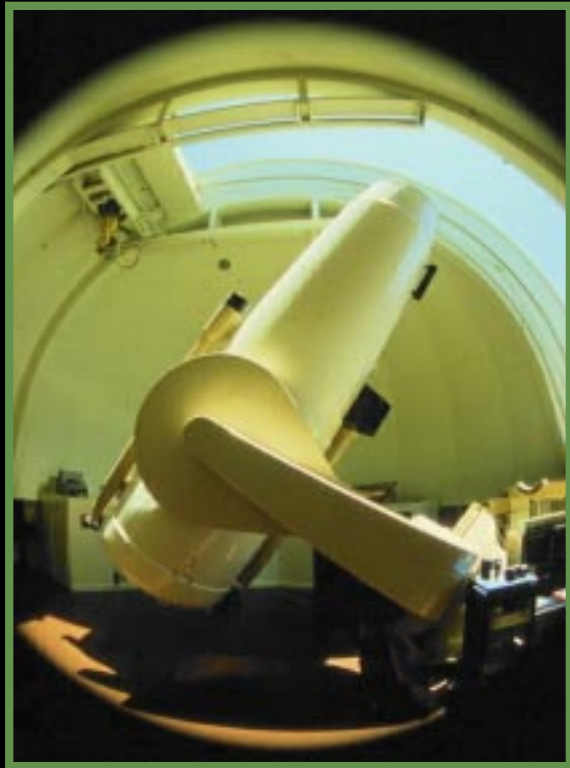
HST · WFPC2

PRC96-01a · ST ScI OPO · January 15, 1996 · R. Williams (ST ScI), NASA

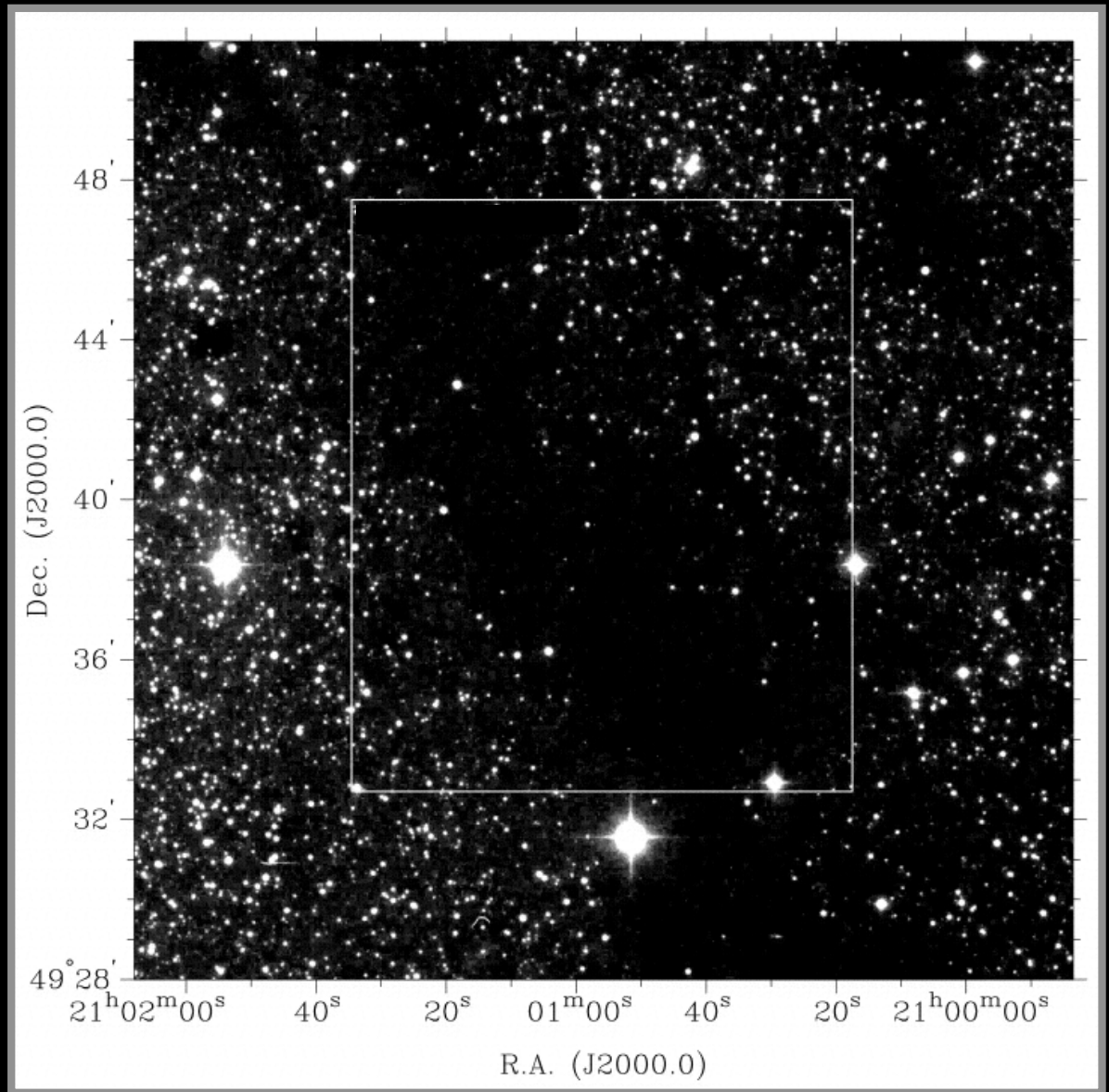
"Star Formation 101"



Molecular Clouds: The Stuff of New Stars



The Oschin telescope,
48-inch aperture wide-field
Schmidt camera at Palomar

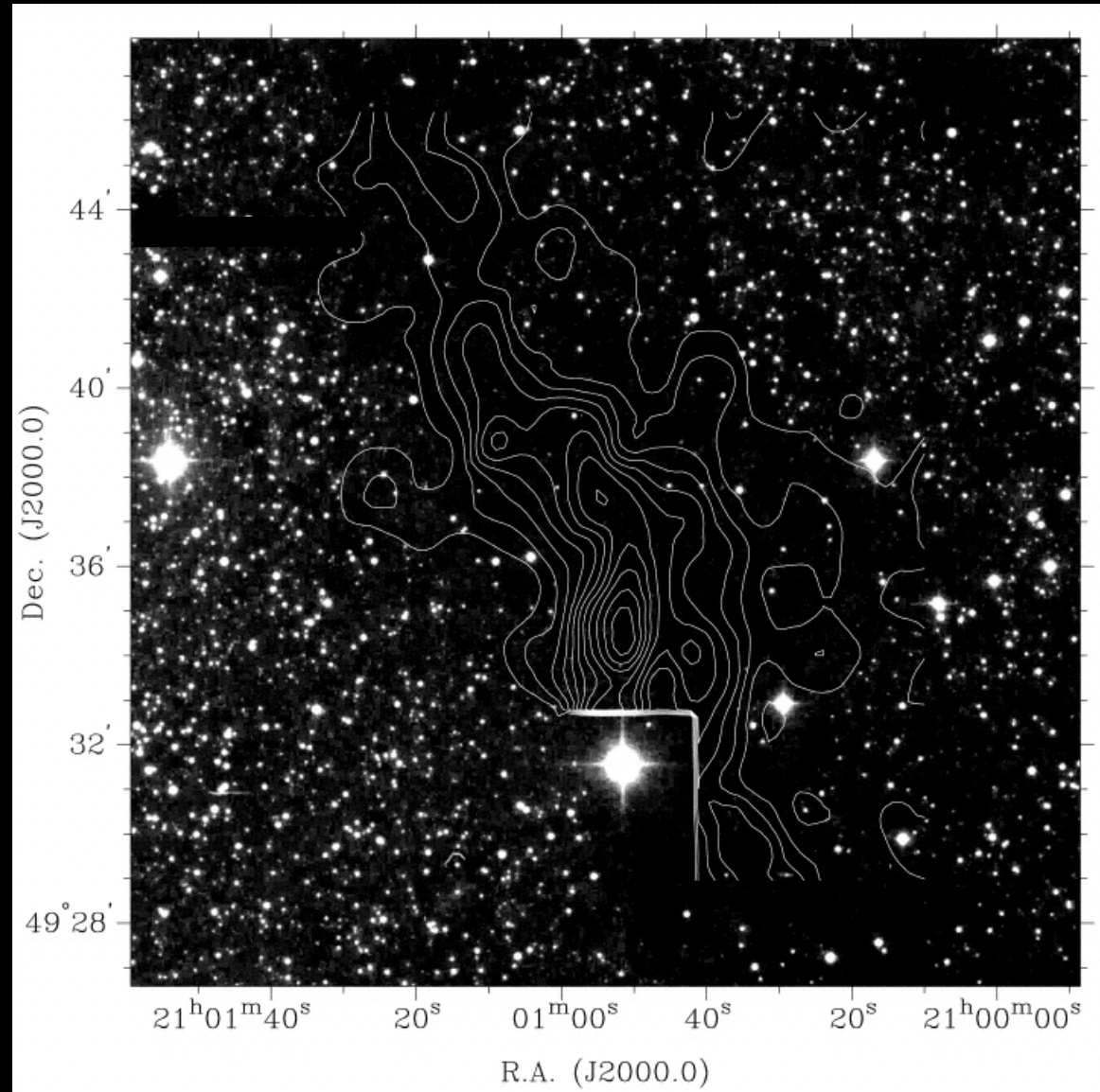


Red Plate, Digitized Palomar Observatory Sky Survey

How much stuff is there?

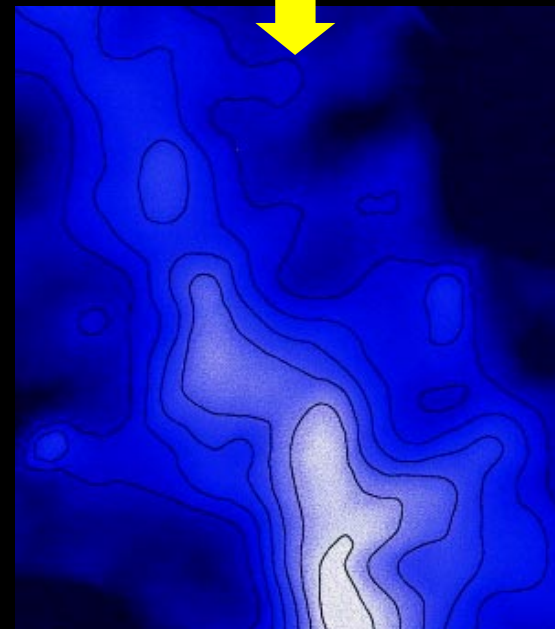
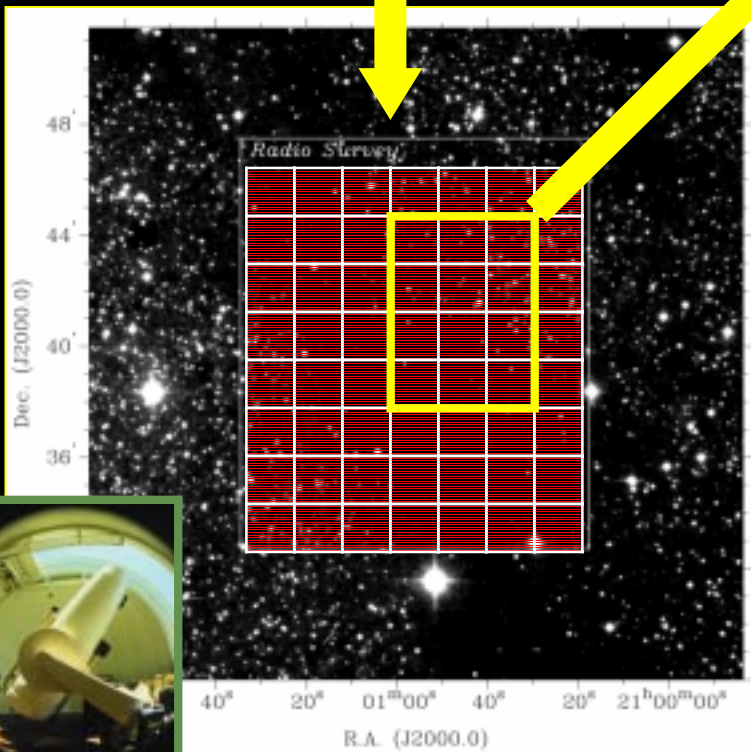
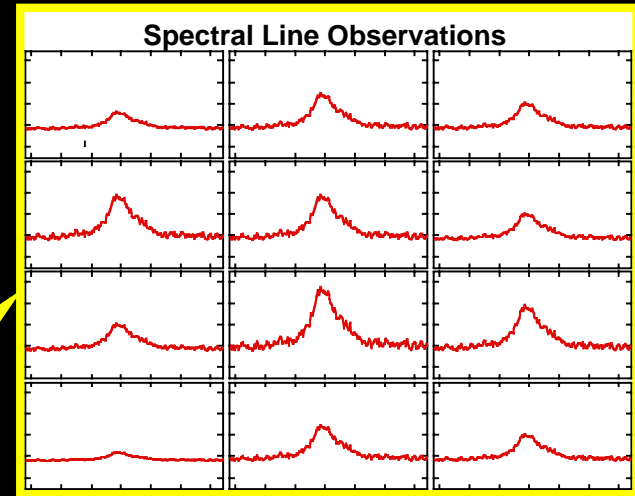
"Star-counting"

Counts of stars per
unit area measure
how much material
must be producing
obscuration.

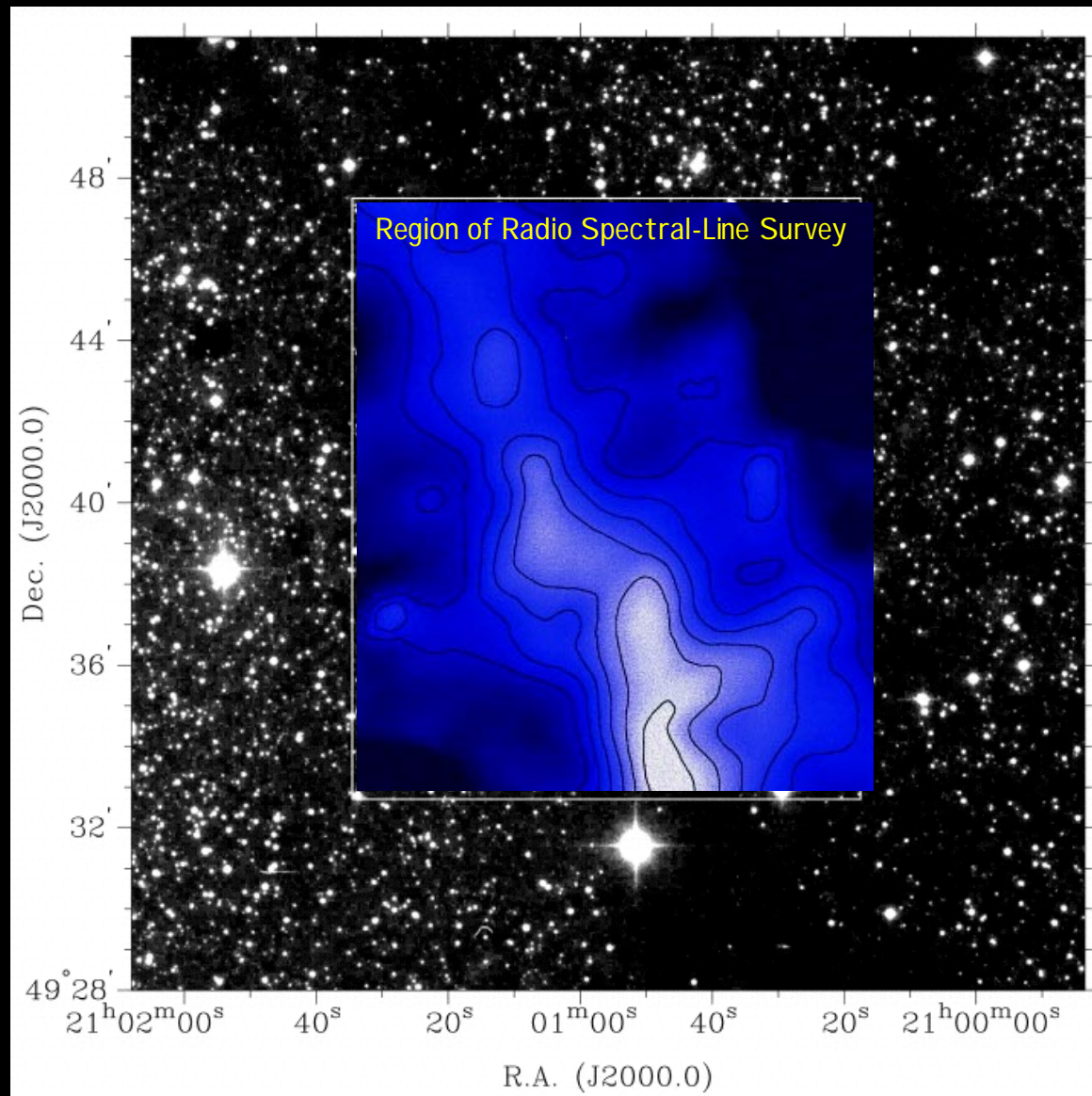


Observations by Alves, Lada & Lada 1999

Radio Spectral-line Observations of Molecular Clouds

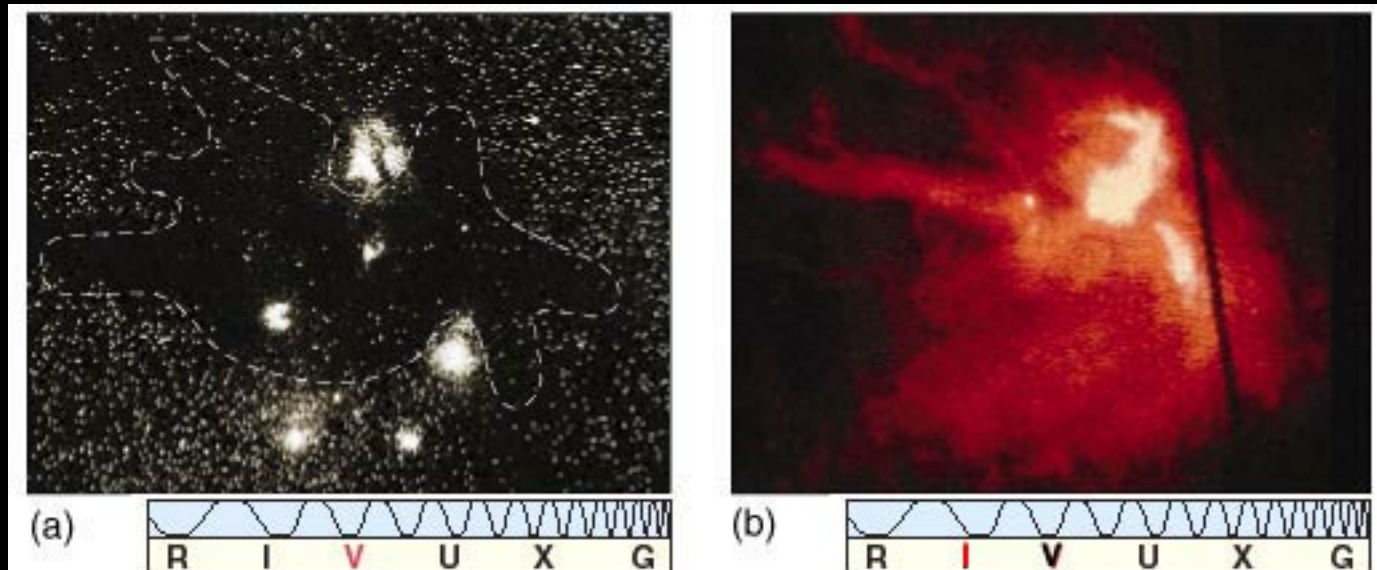


How do Optical & Radio Views Compare?



Observations by Alves, Lada & Lada 1999

Cold, Dark & Dusty



- Gas and Dust are Very Cold in Molecular Clouds, $T \sim 10-100$ Kelvin
- **Dust** at 10 K **"Glows"** in the **Far-Infrared**



Recycling in the Universe

Storage & Collection
in Interstellar Medium



Processing, Production



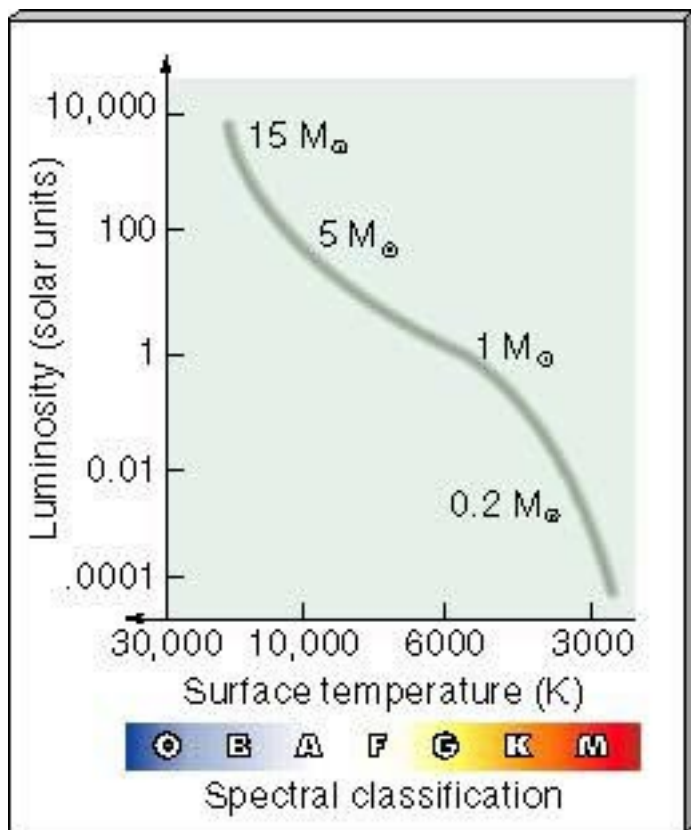
Discarding



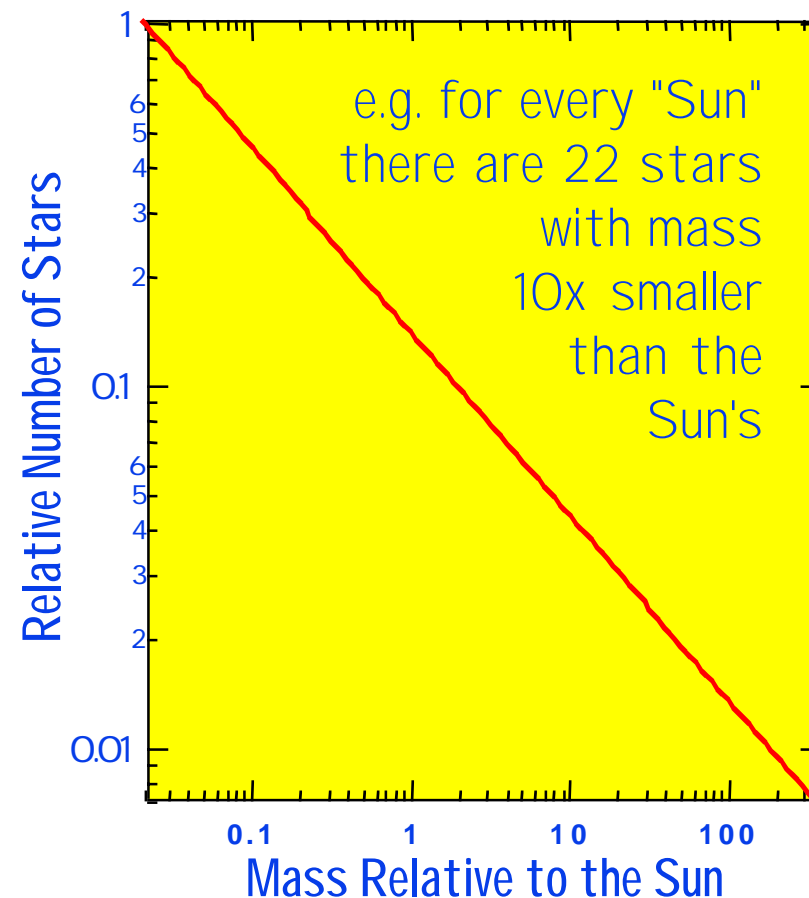
Consumption

Consumption of Recyclables

The Hertzsprung-Russell Diagram

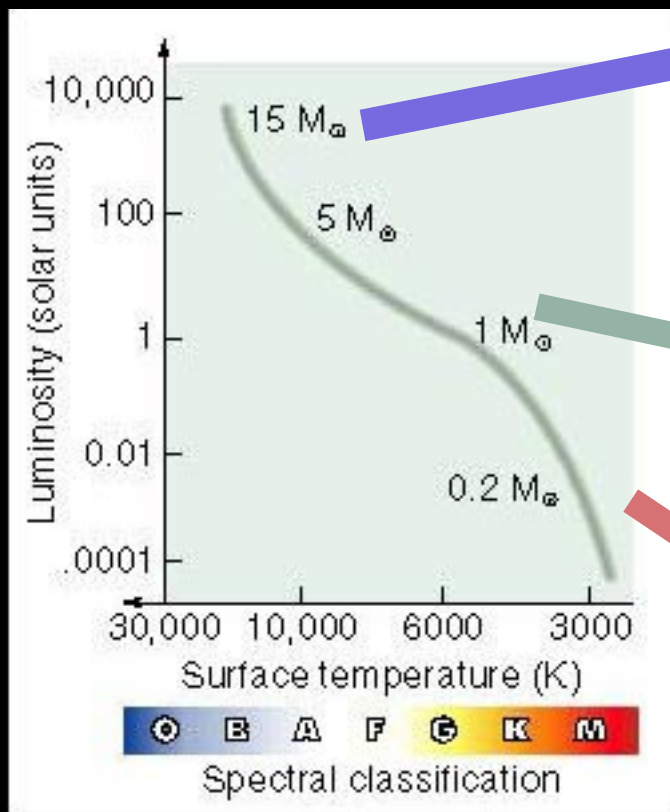


The "Initial Mass Function" (IMF)



Stellar Recyclables

The Hertzsprung-Russell Diagram



Supernova,
then neutron
star/pulsar or
black hole

Spectacular
contribution, and
collection. Explosion
injects, and "sweeps
up" interstellar
material.

Red giant then
white dwarf

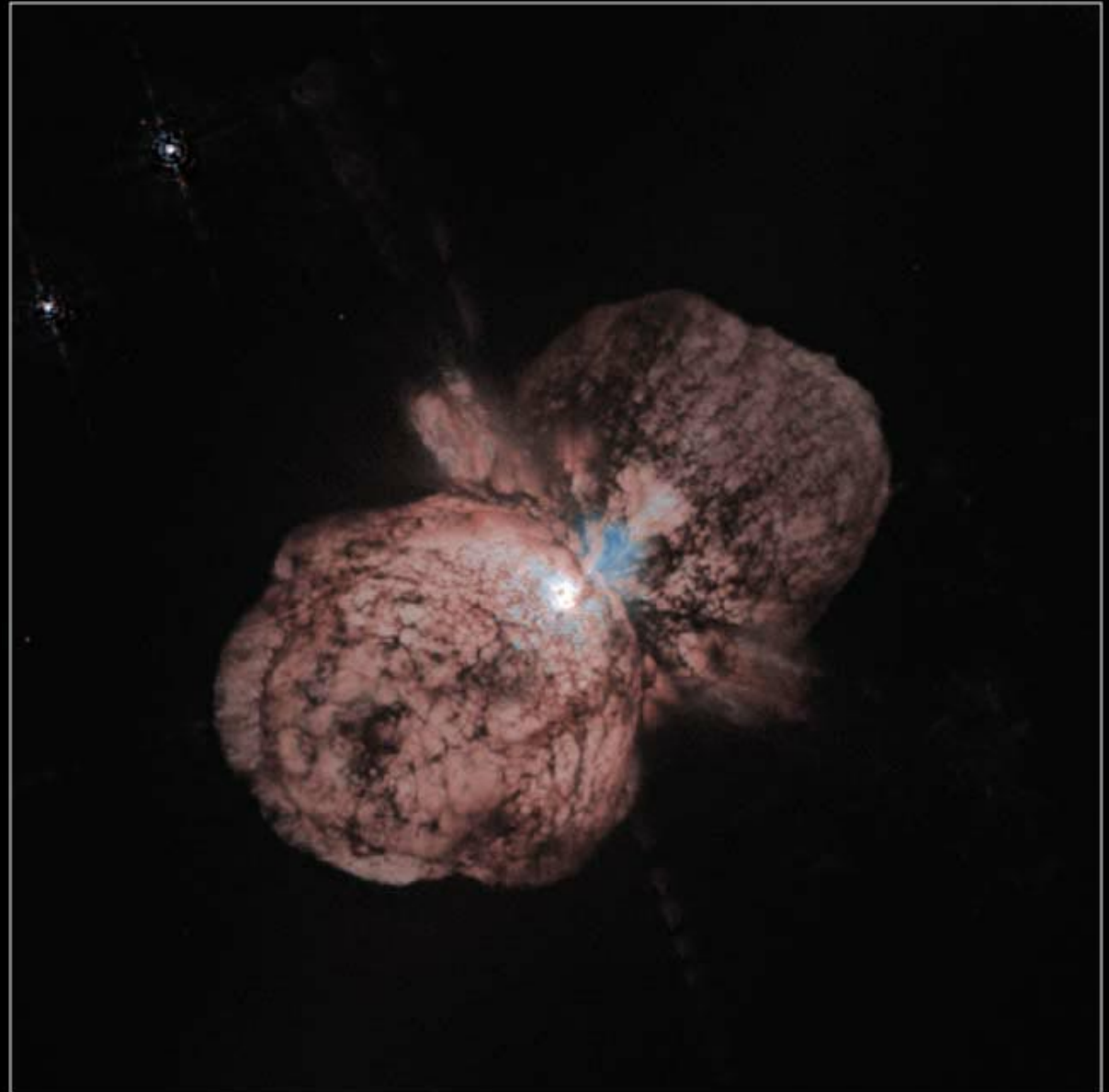
Good recyclables.
Red-giant wind main
dust injection in ISM.

Long-lived
brown dwarfs

"Styrofoam"

Stellar Winds: Discarding the Recyclables

Mass=100 x Sun



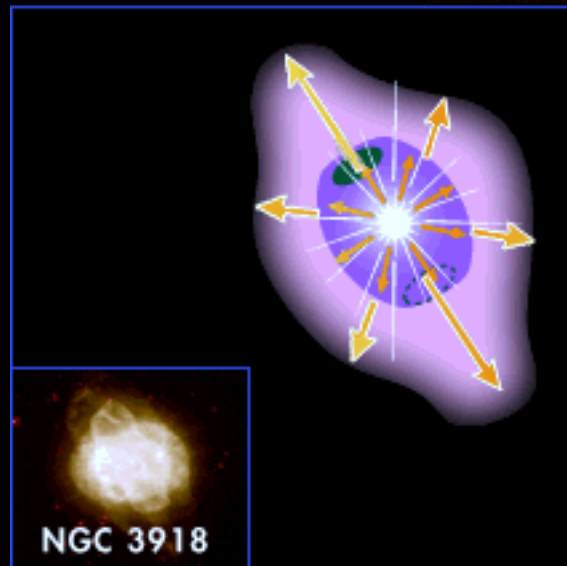
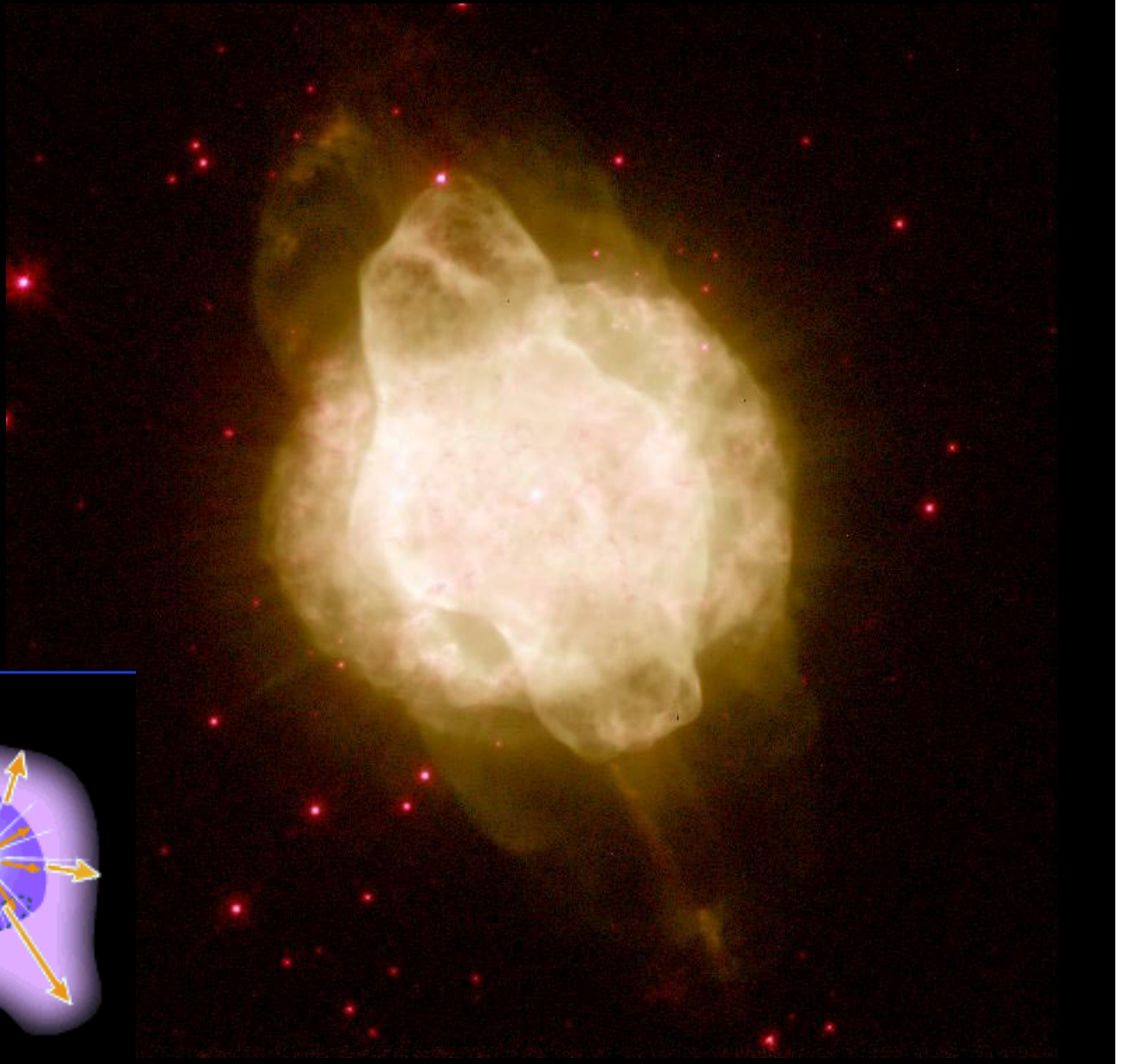
Eta Carinae

HST • WFPC2

PRC96-23a • ST ScI OPO • June 10, 1996
J. Morse (U. CO), K. Davidson, (U. MN), NASA

"Excess Gas?"

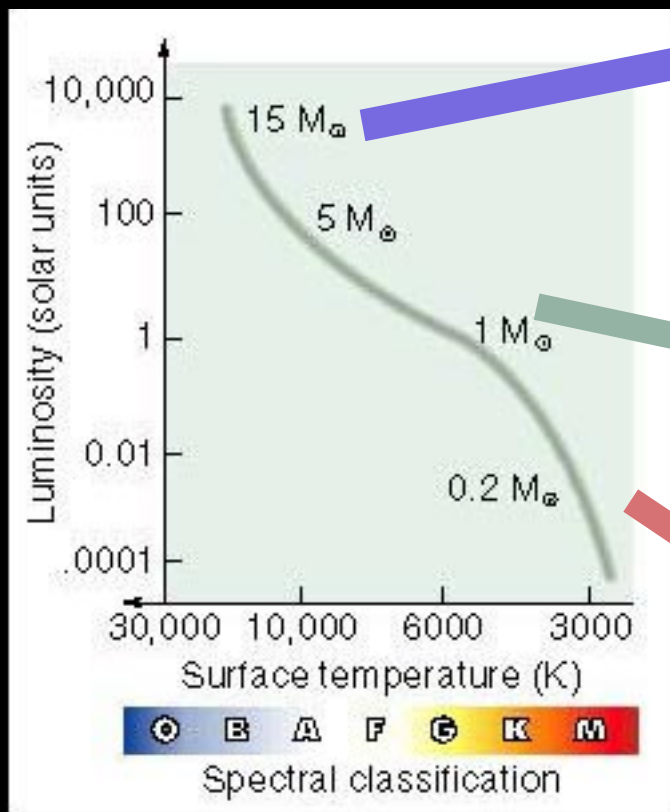
(Post-red-giant planetary nebula)



NGC 3918

Stellar Recyclables

The Hertzsprung-Russell Diagram



Supernova,
then neutron
star/pulsar or
black hole

Spectacular
contribution, and
collection. Explosion
injects, and "sweeps
up" interstellar
material.

Red giant then
white dwarf

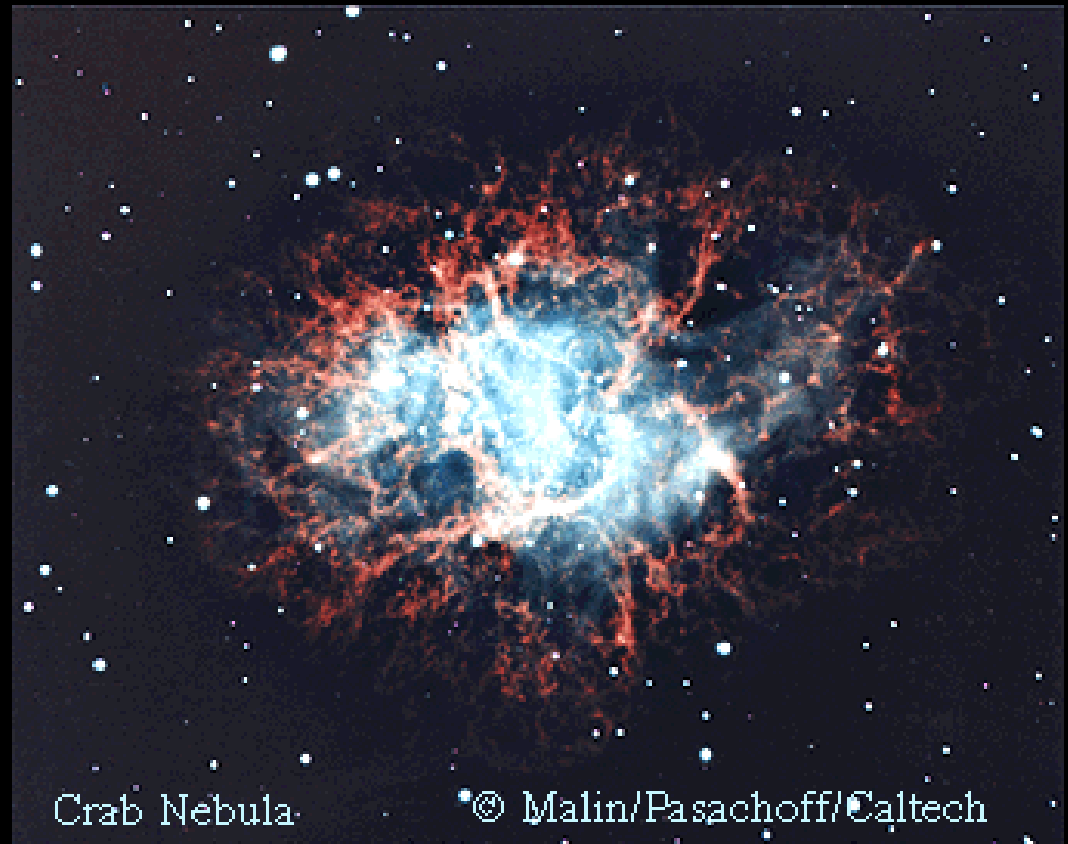
Good recyclables.
Red-giant wind main
dust injection in ISM.

Long-lived
brown dwarfs

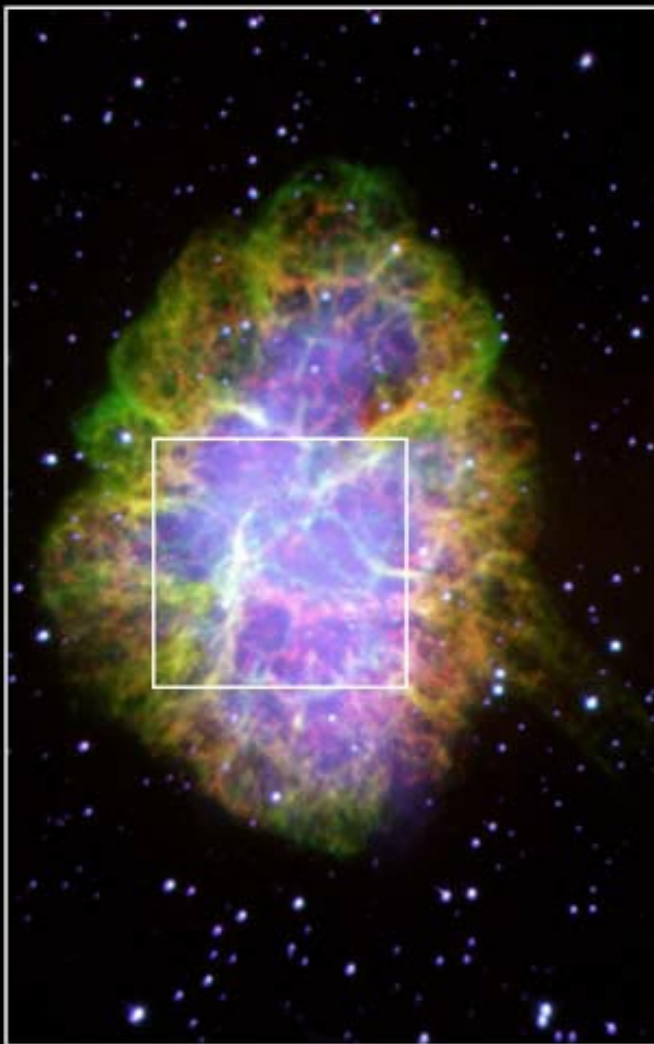
"Styrofoam"

Massive Stars & Supernovae

- **Winds** from O stars account for 30% of recyclable input to ISM
- **Supernovae** from O stars throw out much of the remaining mass
- Biggest contribution of **(correlated) supernovae** is to "collection"

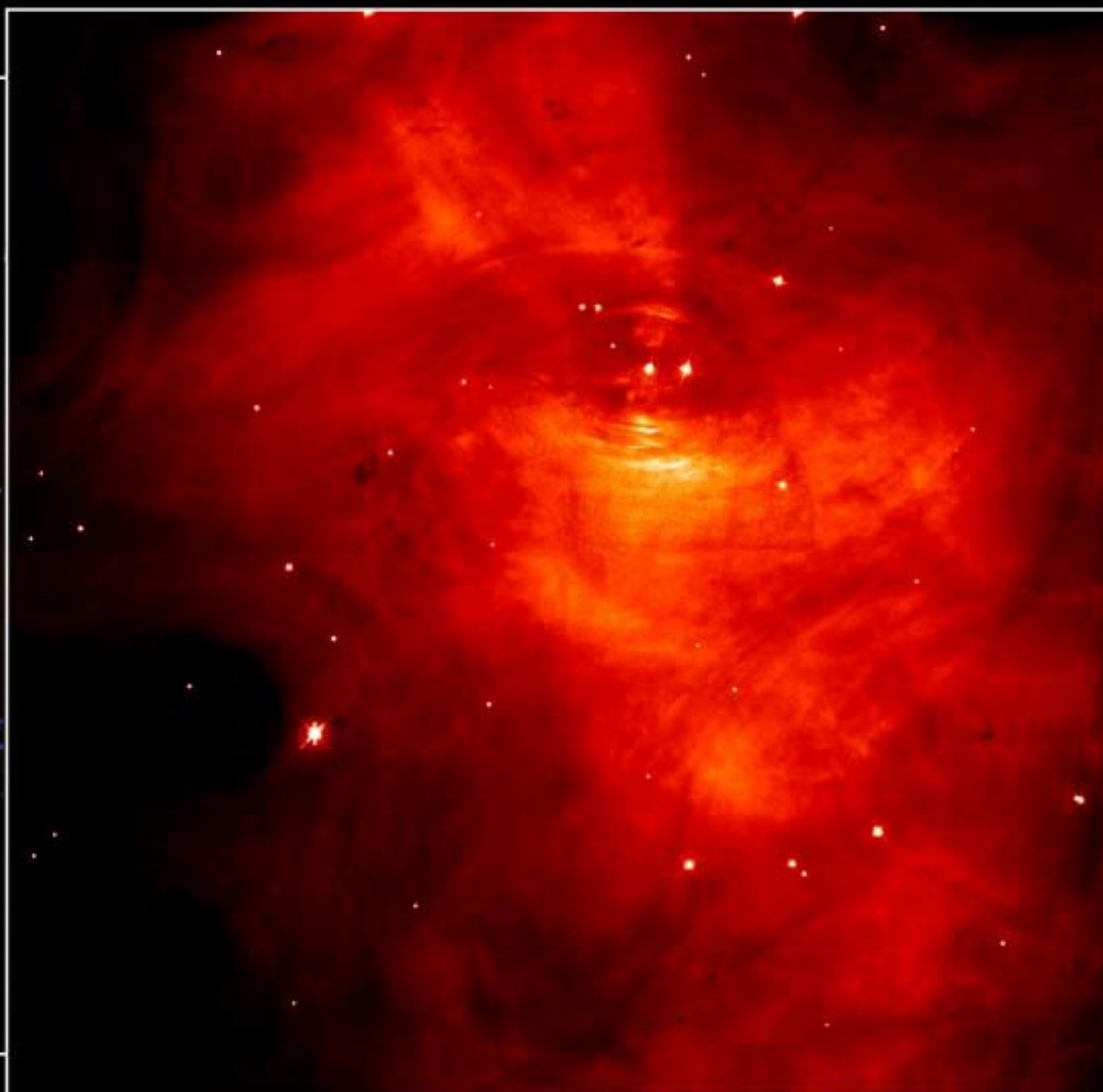


Crab Nebula



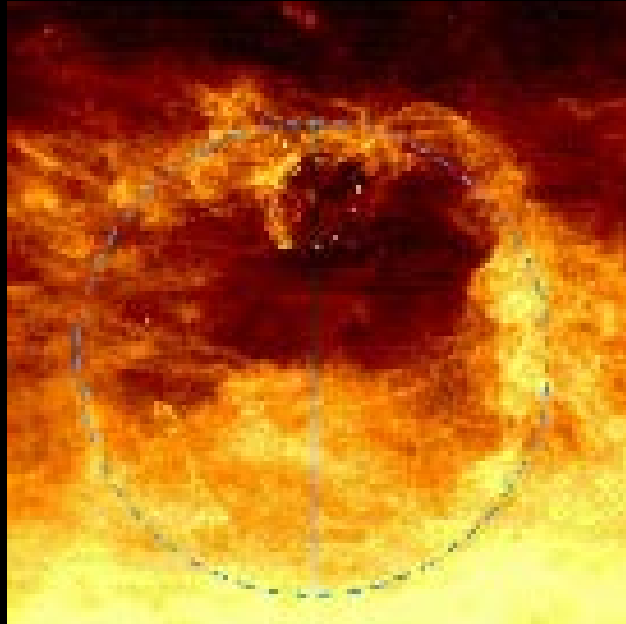
Palomar

PRC96-22a · ST ScI OPO · May 30, 1996
J. Hester and P. Scowen (AZ State Univ.) and NASA



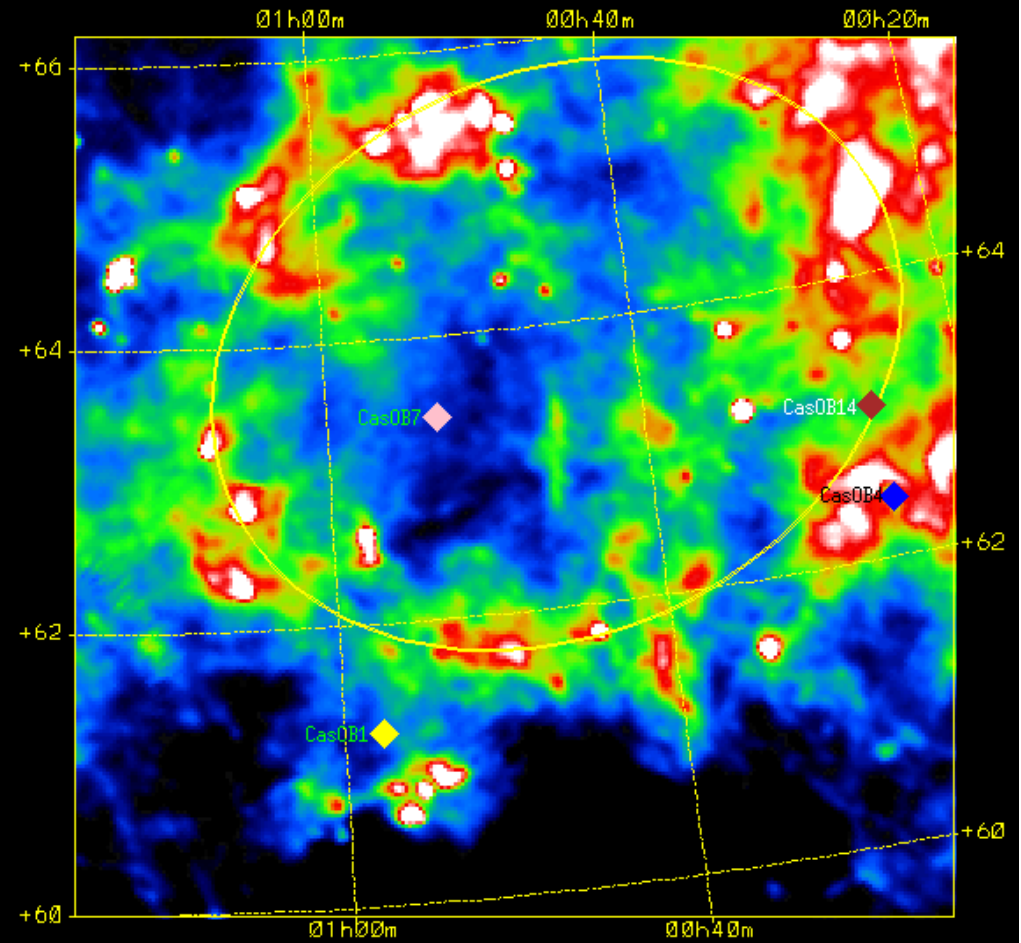
HST · WFPC2

Swept-up Gas: The Next Generation



Far-infrared dust emission
map of North Celestial Pole Loop,

Pound & Goodman 1997



Spectral-line Emission from
Gas In Cassiopeia
Tóth et al. 1995

(At least) How much Gas is Swept-Up?



$$\text{Radius} = R = 20 \text{ pc}$$

$$\text{Volume} = V = \frac{4\pi}{3} R^3$$

$$\text{Density} = \rho$$

$$\rho_{ISM} \approx 1 \text{ atom / cc} = 1.67 \times 10^{-24} \text{ g / cc}$$

$$\text{Mass}_{\text{swept-up}} = \frac{4\pi}{3} R^3 \rho = 1.5 \times 10^{36} \text{ g} = 800 \text{ Solar Masses}$$

Warning Globe Readers: Simple Algebra to Follow!

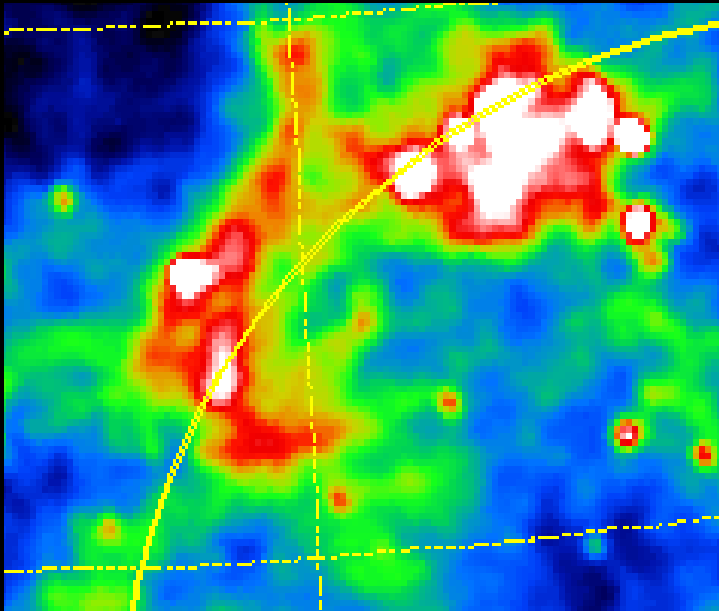


Recycling in the Universe

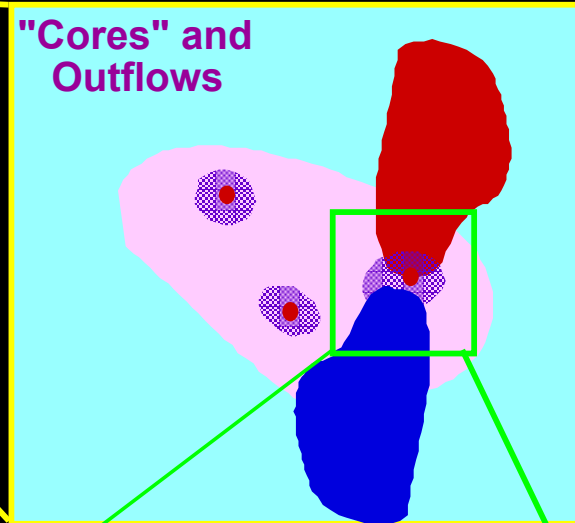
Storage & Collection
in Interstellar Medium



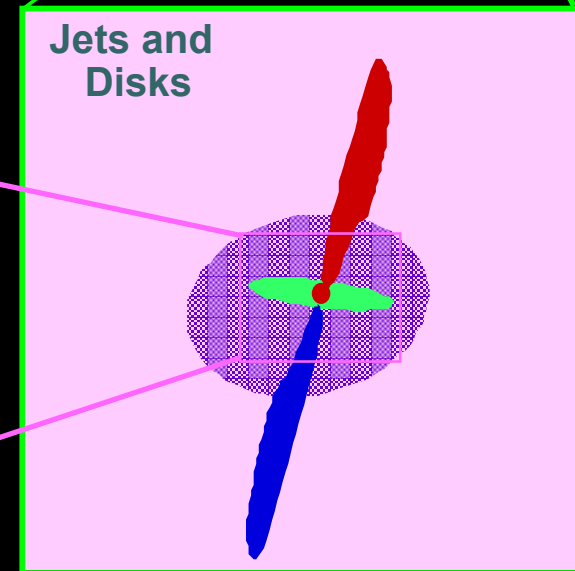
Young Stars do Their Share Too



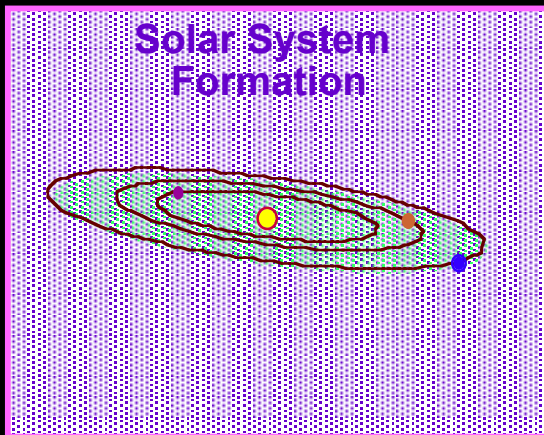
"Cores" and Outflows

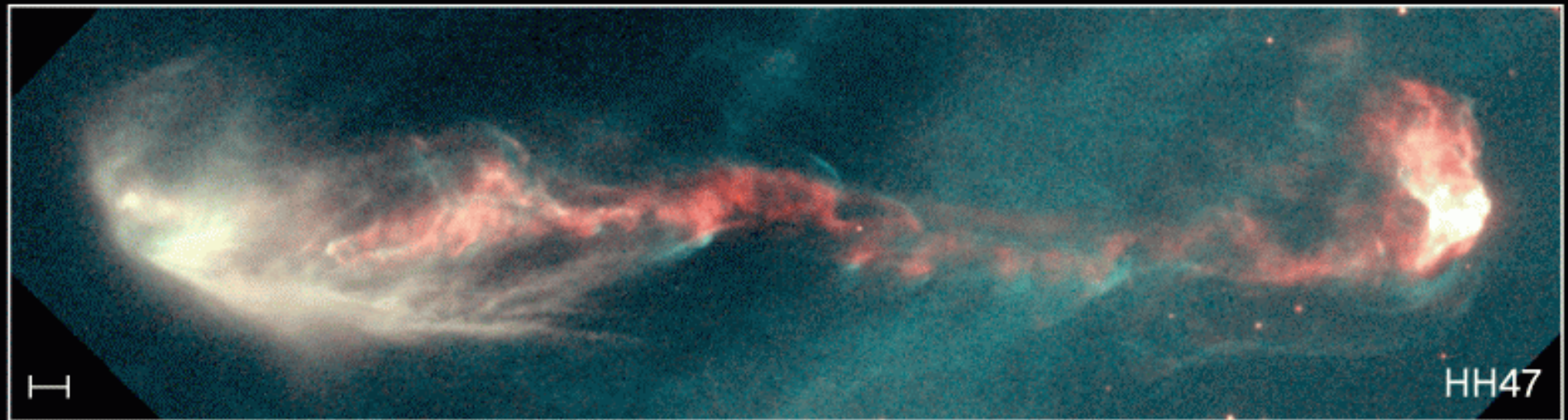
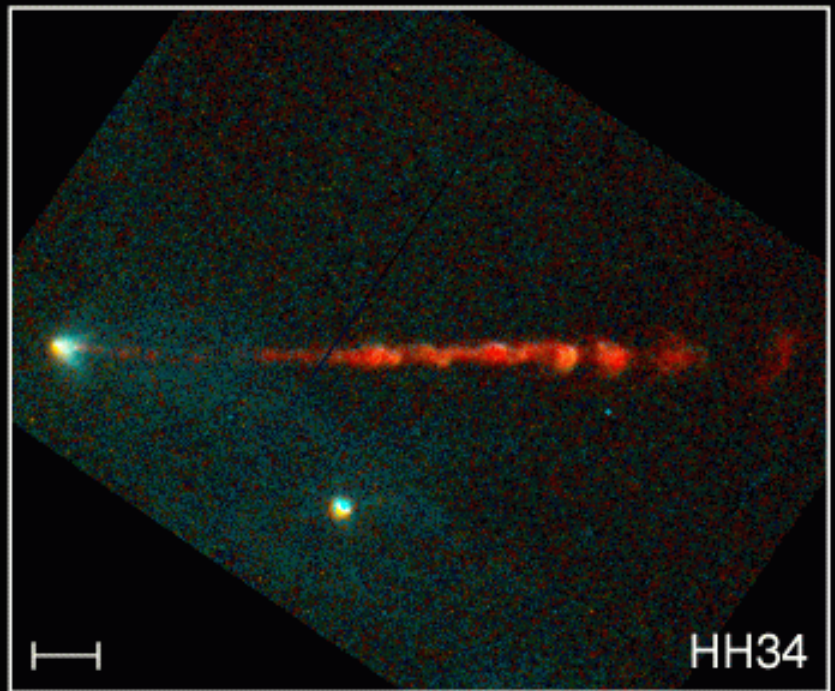
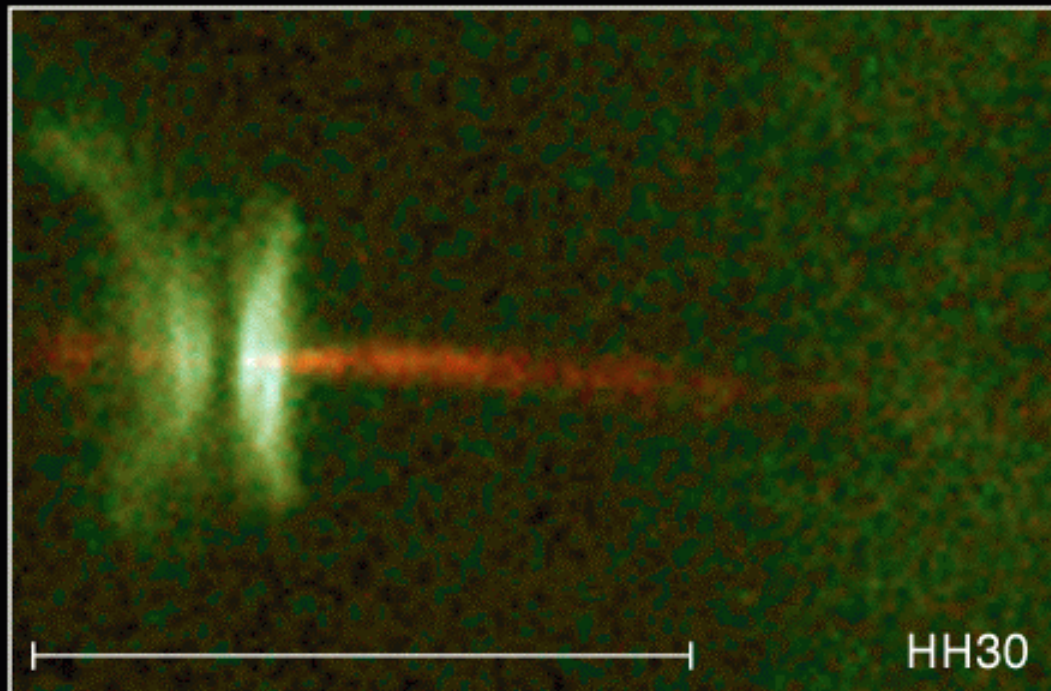


Jets and Disks



Solar System Formation





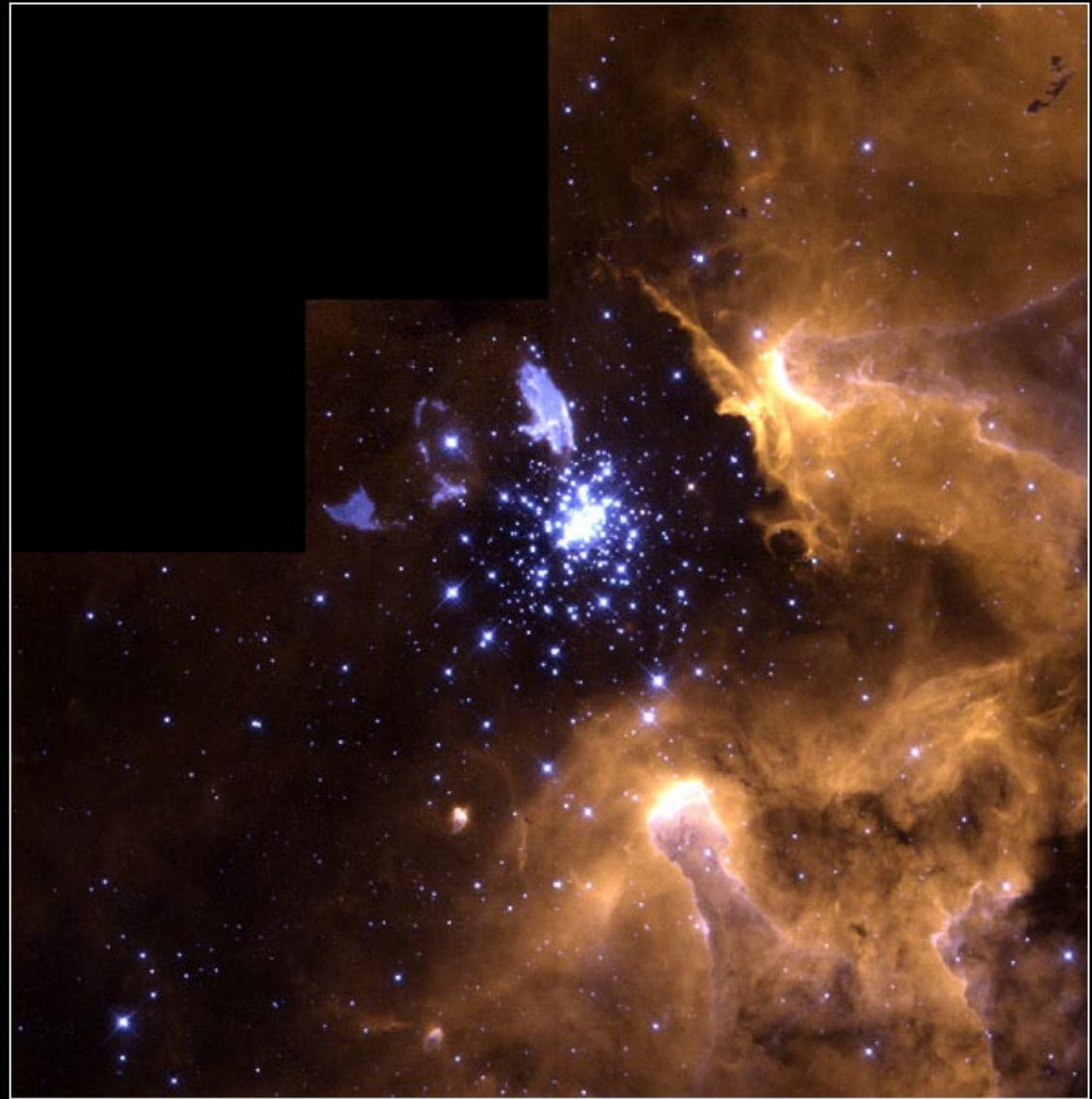
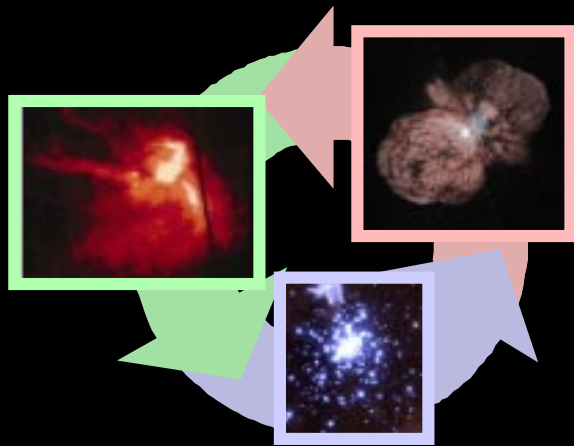
Jets from Young Stars

HST • WFPC2

PRC95-24a • ST ScI OPO • June 6, 1995

C. Burrows (ST ScI), J. Hester (AZ State U.), J. Morse (ST ScI), NASA

One Picture with the Whole Story



NGC 3603

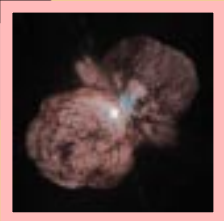
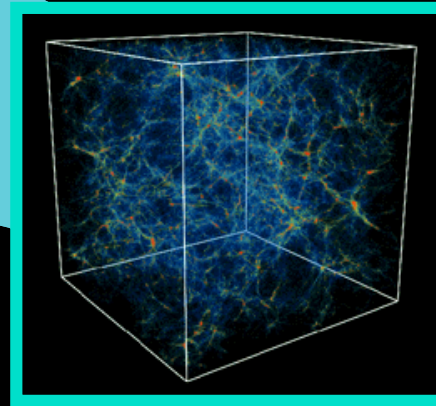
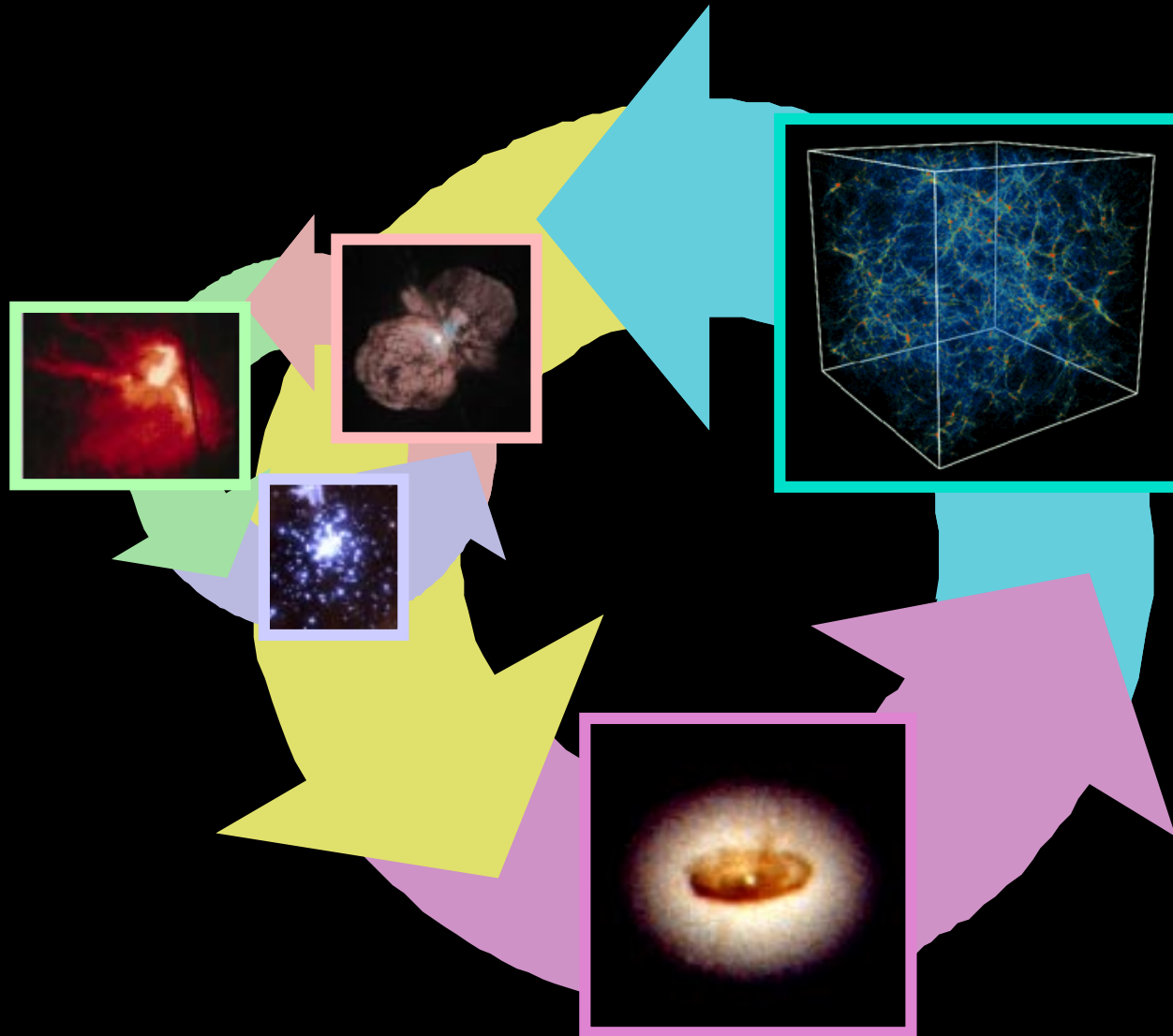
HST • WFPC2

PRC99-20 • STScI OPO • June 1, 1999

Wolfgang Brandner (JPL/IPAC), Eva K. Grebel (Univ. Washington),
You-Hua Chu (Univ. Illinois, Urbana-Champaign) and NASA



Recycling in the Universe(?)



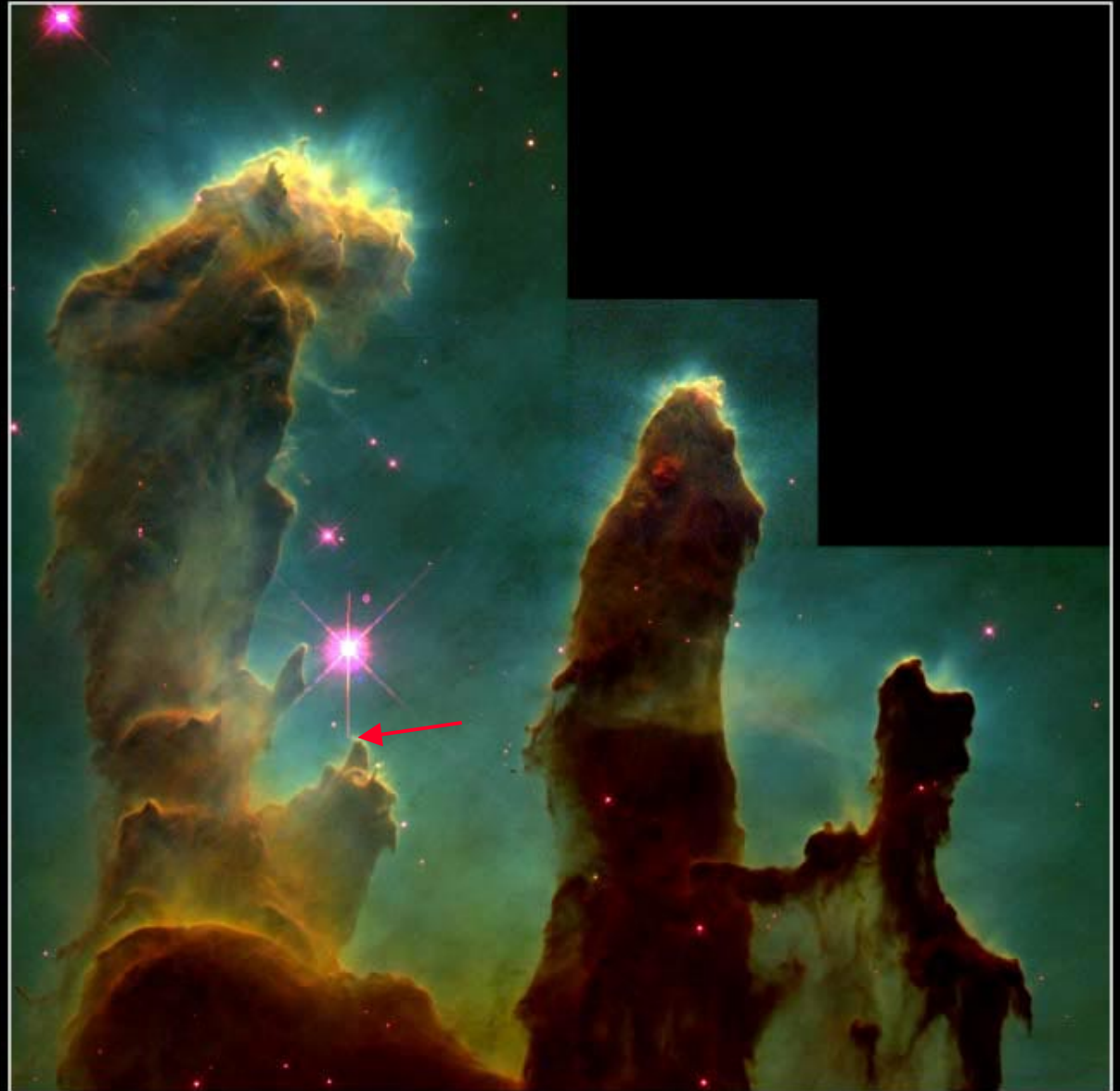
For more information...

cfa-www.harvard.edu/~agoodman

and

Alyssa Goodman's upcoming article in
Sky & Telescope Magazine

(Unusual?)
Stellar
Nursery
in the Eagle
Nebula



Gaseous Pillars • M16

HST • WFPC2

PRC95-44a • ST Sci OPO • November 2, 1995
J. Hester and P. Scowen (AZ State Univ.), NASA

Star Formation Caused by A Galaxy Collision (a.k.a. igniting the trash)

