10. Diatomic molecules

Suggested reading: **Bernath** Chapter 6 and Chapter 7, Section 1 (vibrations of diatomics). See **Herzberg**, *Diatomics* for great detail.

Potential energy curves:



The potential curve is approximately harmonic (*i.e.*, a parabola) near the bottom. A better, and still analytic, potential is the Lennard-Jones (or 6-12) potential:

$$V(r) = 4a \left[\left(\frac{b}{r}\right)^{12} - \left(\frac{b}{r}\right)^{6} \right].$$

It has nice analytic properties for calculations of spectroscopic and thermodynamic properties. Also note the Morse potential, in Bernath:

 $V(r) = D(1 - e^{-\beta(r-re)})^2$. In general, though, accurate potentials are non-analytic, and are derived either from inversion of spectroscopic data or from complex theoretical calculations.





The potential well pictured on the previous page is for a particular electronic state (a bound state), which could well be the electronic ground state (see O_2 , just above, for a more complex system of potentials). The **rotational energy** *levels* corresponding to this state can be expanded as:

 $E = BJ(J+1) - DJ^{2}(J+1)^{2} + HJ^{3}(J+1)^{3} + \dots +$

Departures from the simple BJ(J+1) are due (primarily) to *centrifugal distortion*. By convention, only the first distortion term, $-DJ^2(J+1)^2$ has a negative sign. *D* is usually (perhaps always?) positive. Higher terms can in principle be positive or negative.

Rotational partition functions for diatomics are approximated by $Q_R \simeq T/c_2 B$, which is accurate for $B \ll kT$ and negligible centrifugal distortion. This is a very occasionally useful approximation but it is mostly a relic of pre-computer days. As an example, the HITRAN Database of line parameters, which we will meet later, tabulates partition functions for every 1°K. Interpolating assuming $Q_R \simeq T/c_2 B$ gives very accurate values in such a case. The **vibrational energy levels** (Hermite polynomials are the wave functions for the harmonic approximation to the vibrational potential) are approximately $E = \omega(v+1/2), v = 0, 1, 2, \cdots$ where ω is the *vibrational constant*. It is related to the Hooke's law restoring force and the reduced mass for vibration, μ_v : $\omega \propto \sqrt{k/\mu_v}$. Note that the lowest vibrational state, for v = 0 has energy $\omega/2$; this is the vibration *zero-point energy*, required by the Heisenberg uncertainty principle. In general, vibrations also have distortions and terms coupling vibration to rotation (*Bernath*, Eq. 7.28): $E = \omega(v+1/2) - \omega_e X_e (v+1/2)^2 + BJ(J+1) - DJ^2(J+1)^2 - \alpha_e (v+1/2)J(J+1) + \cdots$

Vibrational transitions

Torquing between vibration states is accomplished through $d\mu/dr$ instead of μ (see

why?) and intensities are $\propto \left|\frac{d\mu}{dr}\right|_{r_e}^2$ + higher-order terms (μ here is dipole moment; sorry

about any confusion with the μ_v of vibrational energy levels). This explains, for example, why OH vibrational transitions are weak while rotational transitions are strong: μ is large, while $d\mu/dr$ is small near r_e (it has zero derivative near r_e).



Vibration band types: The *fundamental* transition is $v = 0 \rightarrow 1$ in absorption, $v = 0 \leftarrow 1$ in emission. *Overtone* transitions have $\Delta v > 1$. *Hot band* transitions are between states with $v \ge 1$. This gets much more complicated very quickly with polyatomics (discussed later) since the number of fundamentals goes as 3N - 6(3N - 5 for linear molecules), where N is the number of atoms in the molecule.

The vibrational partition function is $Q_v = 1 + e^{-c_2\omega/T} + e^{-2c_2\omega/T} + \cdots$, ignoring distortion. (Where did the $\omega/2$ zero-point energy go? Got it? *The choice of the zero of energy for Boltzmann statistics is arbitrary, as long as it is consistent, since any offset occurs in both numerators and denominators.*) Often there is only one upper vibrational state that is thermally accessible (*i.e.*, which has a non-negligible Boltzmann factor).^{*} Then

 $Q_{\nu} \simeq \frac{1}{1 - e^{-c_2 \omega/T}}$. Again, this is now mostly a pre-computer curiosity, except for narrow range interpolation.

^{*}This is less true for polyatomics, with ozone and CO₂ being important cases. It is also less true for hot conditions (astrophysical, combustion, plasmas).