

MEASURING TRACE GAS PROFILES FROM SPACE

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Profiling the Atmosphere

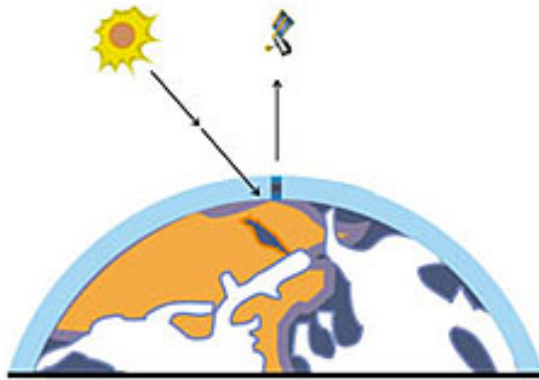
- Active
 - LIDAR
 - RADAR
- Passive
 - Occultation
 - Nadir backscatter
 - Nadir thermal emission (IASI, TES)
 - Limb scattering (OSIRIS, SCIAMACHY)
 - Limb thermal emission (MIPAS)

Profiling the Atmosphere

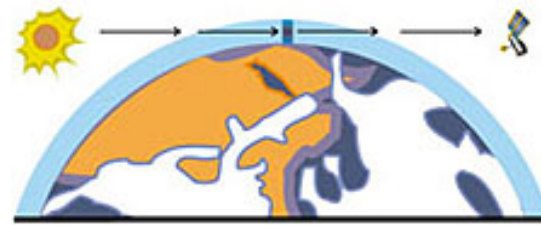
- Active
 - LIDAR
 - RADAR
- **Passive**
 - **Occultation**
 - **Nadir backscatter**
 - Nadir thermal emission (IASI, TES)
 - Limb scattering (OSIRIS, SCIAMACHY)
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Passive Measurement Geometries

A. Backscatter Ultraviolet



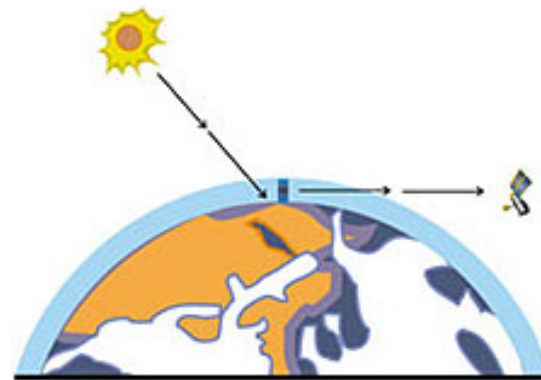
B. Occultation



C. Limb Emission



D. Limb Scattering



Nadir Backscatter Instruments

Instrument	Dates	Profiles retrieved from nadir?
TOMS	1978 – 2006	
GOME	1995 – 2011	O ₃ (Liu et al., 2005)
SCIAMACHY	2002 – 2012	
OMI	2004 –	O ₃ (Liu et al., 2010) SO ₂ (Yang et al., 2010)
GOME-2A GOME-2B GOME-2C	2006 – 2012 – 2017 –	O ₃ (Cai et al., 2012) SO ₂ (Nowlan et al., 2011)
OMPS	2011 –	
TROPOMI	2015 –	
Sentinel-4	2017 –	
TEMPO	2018 –	
GEMS	2018 –	

Remote Sounding: Inverse Problem

$$\mathbf{y} = \mathbf{F}(\mathbf{x}) + \boldsymbol{\varepsilon}$$

- We have \mathbf{y} (spectral data)
- We want \mathbf{x} (profile & other fitted parameters)
- Requirements for inversion
 - Forward model (\mathbf{F})
 - Retrieval algorithm

Nadir Backscatter

- Global coverage
- Limited altitude information, and only for certain molecules (Ozone, volcanic SO₂)
- Almost always need *a priori* information on state of atmosphere (i.e., ozone profile climatology)
- Altitude information on ozone and SO₂ has so far been derived from OMI, GOME, and GOME-2 using the UV

Optimal Estimation Approach

- Combine *a priori* knowledge with measurements
- Iterate until convergence

$$\Delta \mathbf{x} = (\mathbf{K}^T \mathbf{S}_\varepsilon^{-1} \mathbf{K} + \mathbf{S}_a^{-1})^{-1} [\mathbf{K}^T \mathbf{S}_\varepsilon^{-1} \Delta \mathbf{y} - \mathbf{S}_a^{-1} (\mathbf{x} - \mathbf{x}_a)]$$

\mathbf{x} = current guess of retrieved parameter

\mathbf{y} = measurements

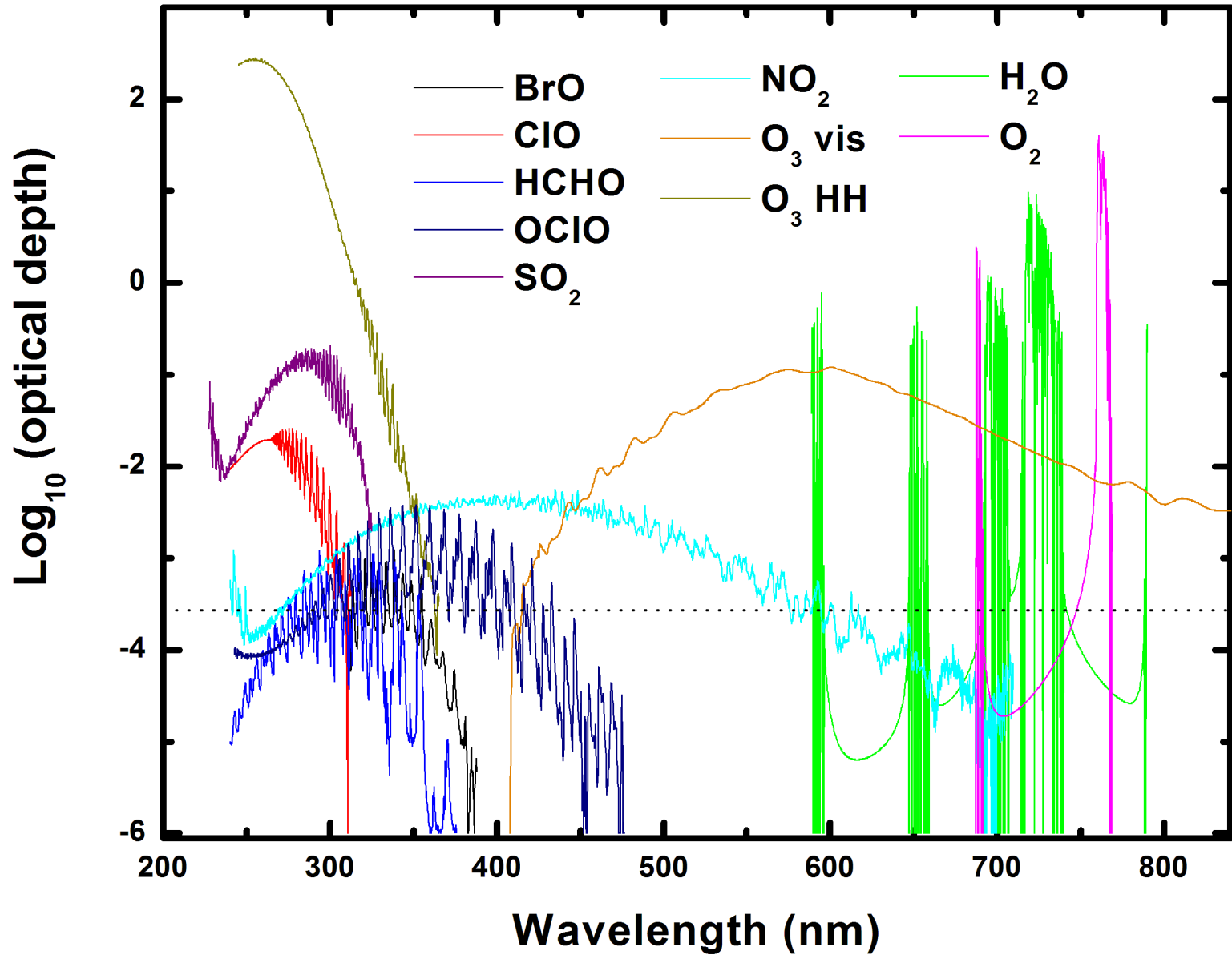
\mathbf{S}_ε = measurement error covariance matrix

$\mathbf{K} = d\mathbf{y}/d\mathbf{x}$

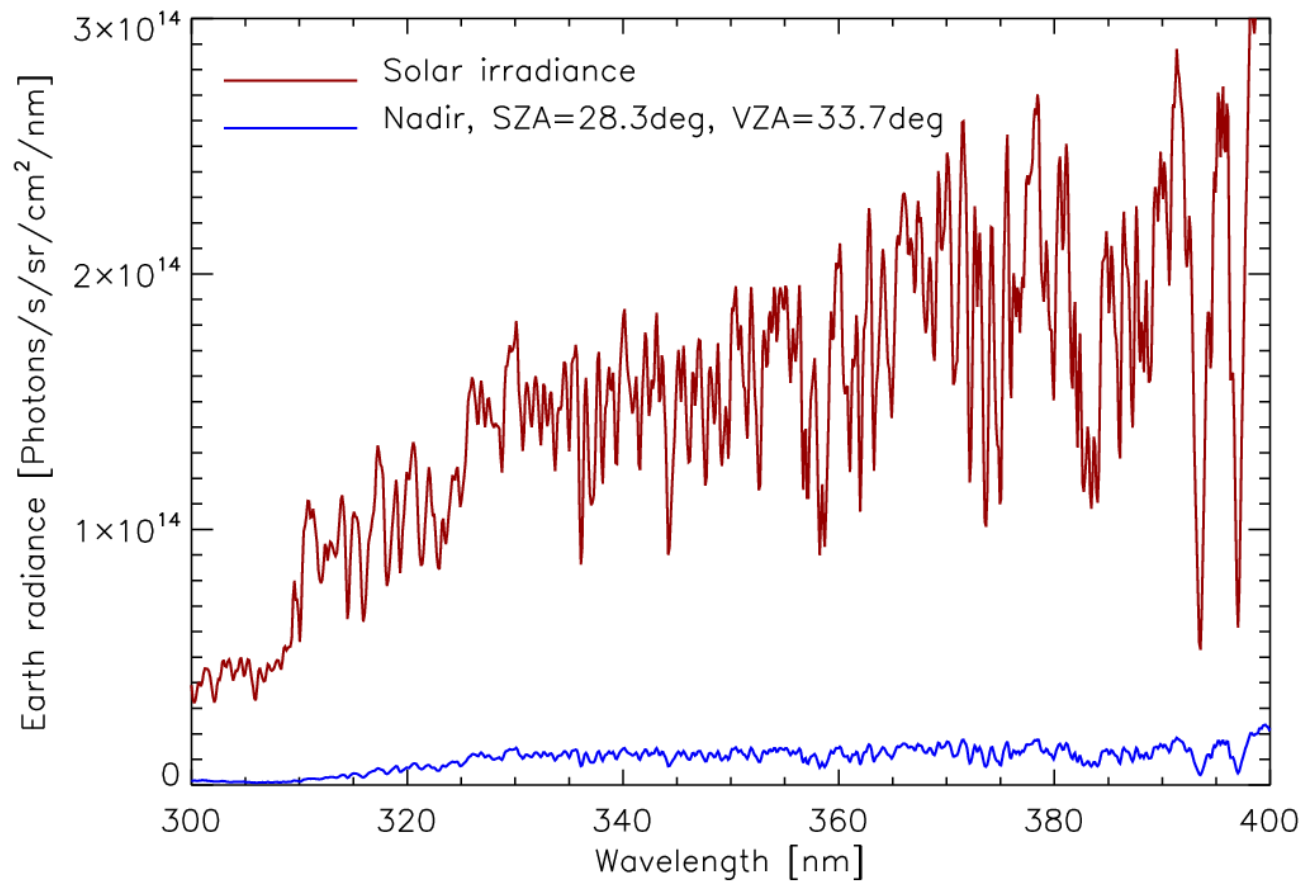
\mathbf{x}_a = a priori profile (climatology or from a model)

\mathbf{S}_a = a priori error covariance matrix

Optical Depths for Typical GOME Measurement Geometry

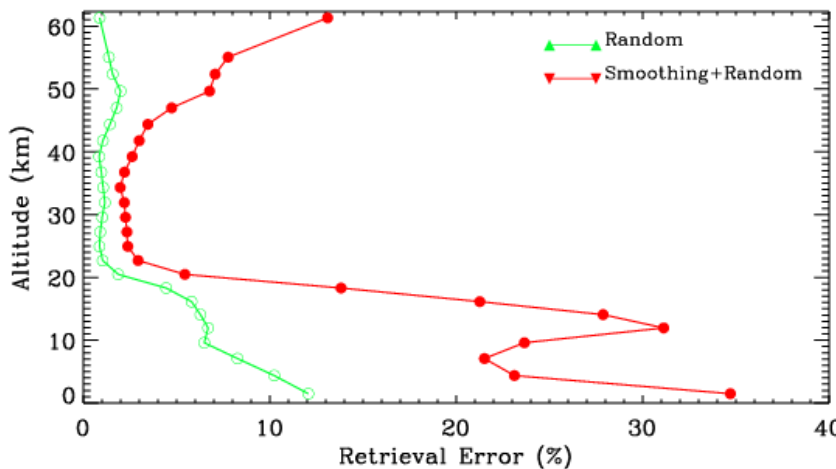
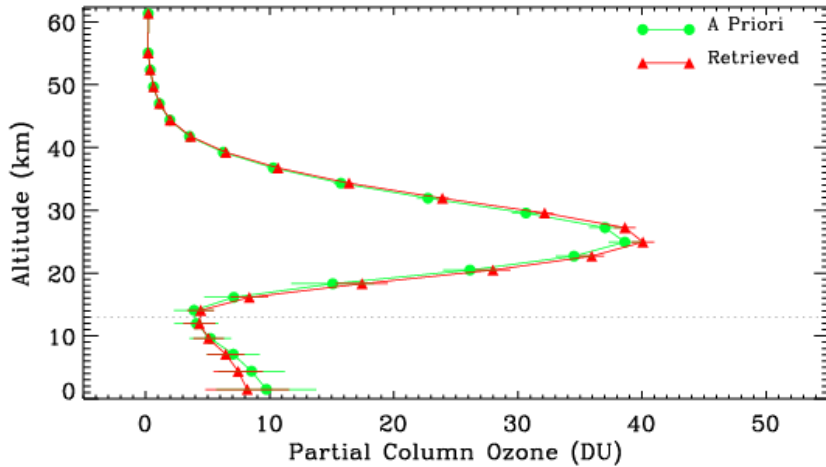


GOME-2 Spectra, Channel 2

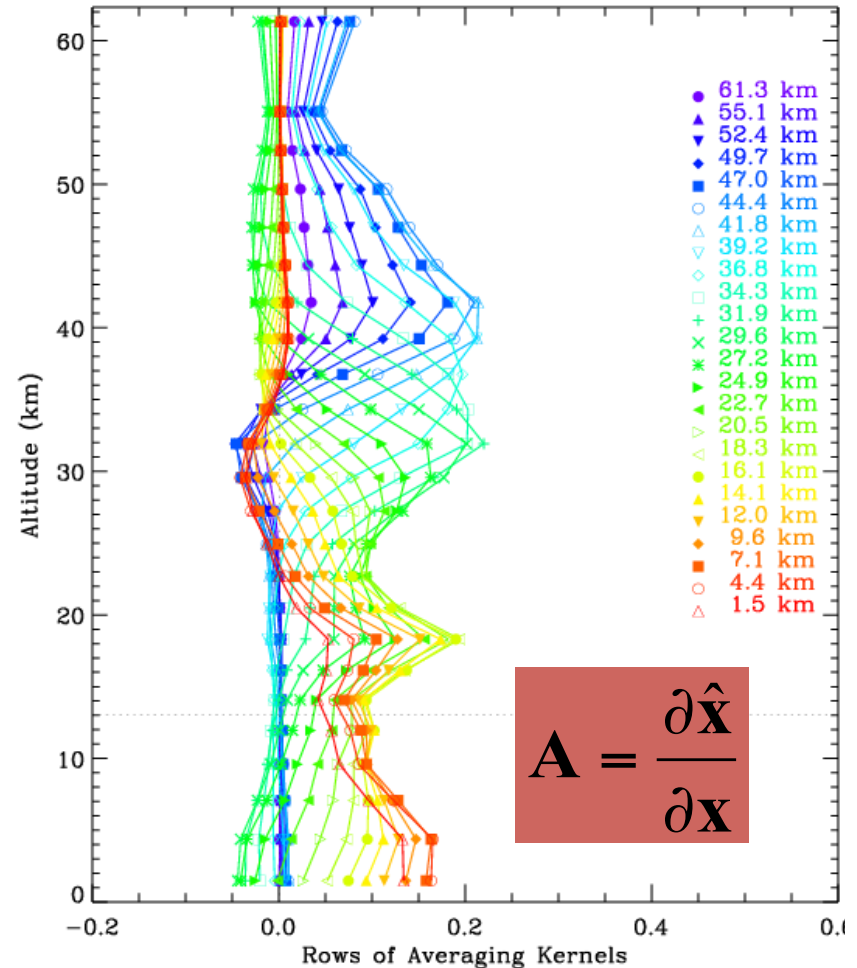


GOME-2 Ozone Profile (Backscatter)

Orbit: 9265, Xtrack: 01, Scan: 0137
 SZA=28, VZA=41, AZA= 5, $f_c=0.20$, $P_c= 895\text{mb}$

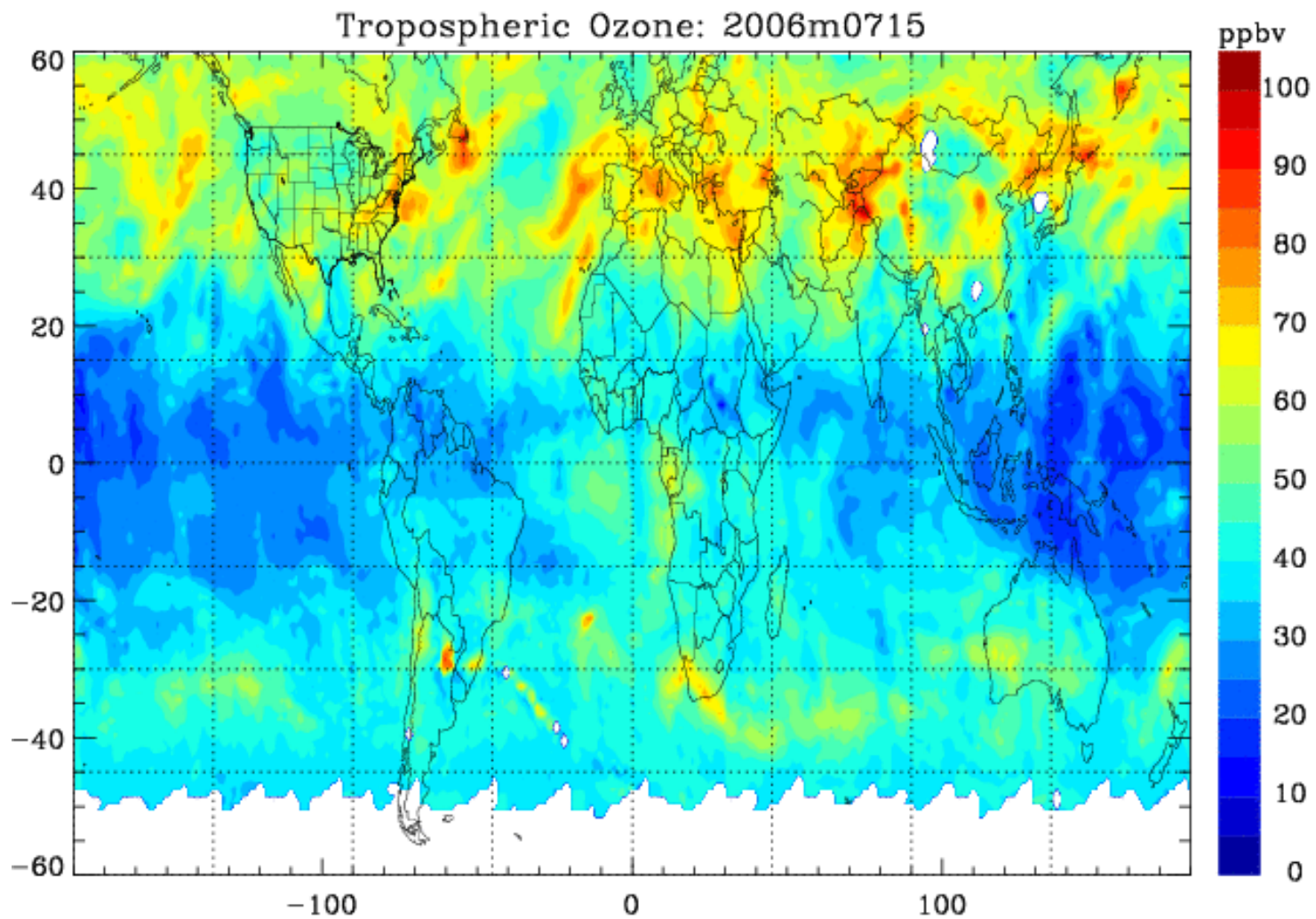


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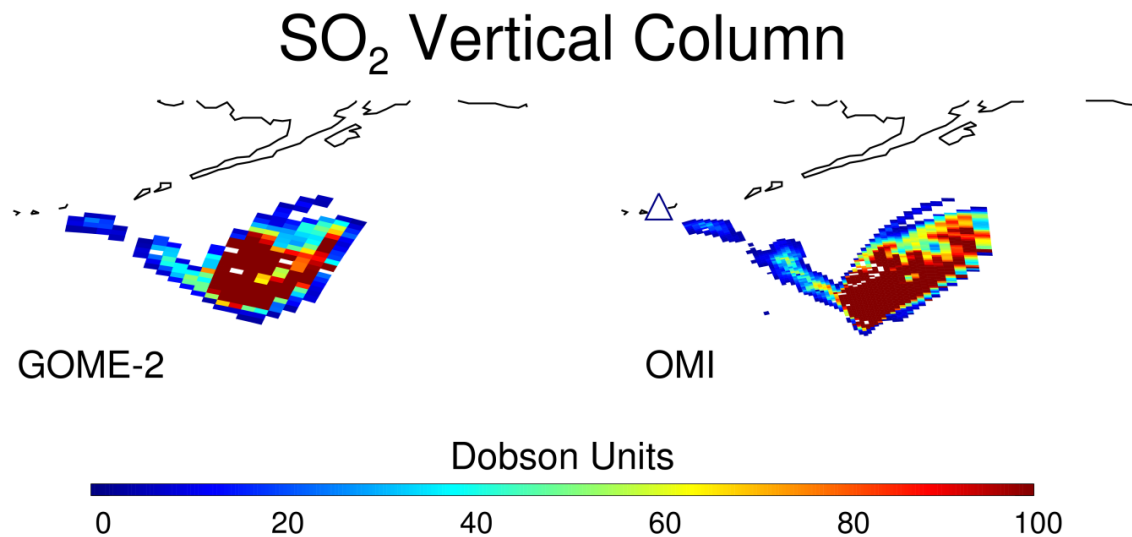
$$\mathbf{A} = \frac{\partial \hat{\mathbf{x}}}{\partial \mathbf{x}}$$

OMI Tropospheric Ozone

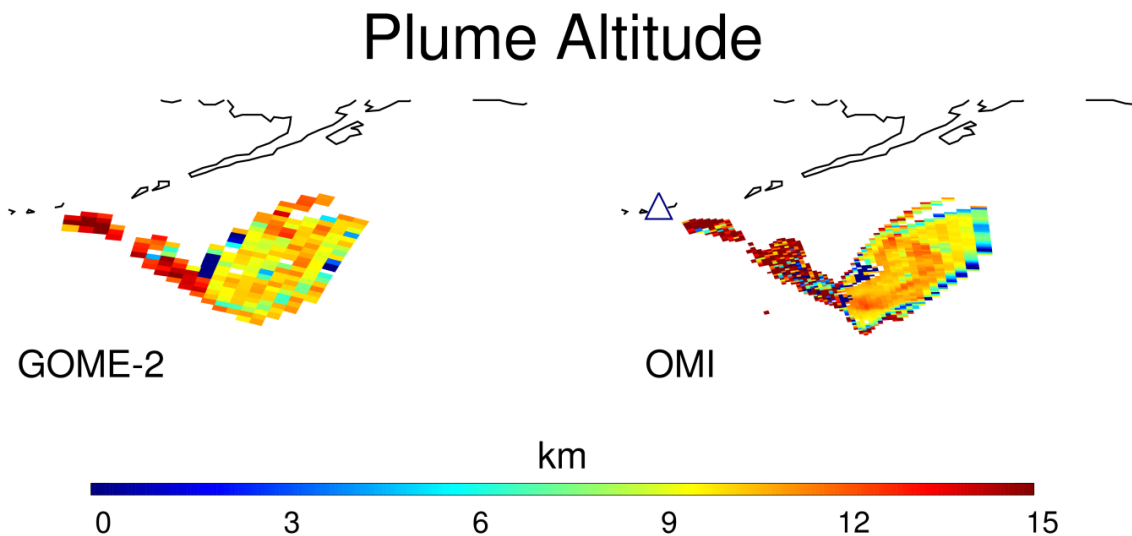


GOME-2 and OMI SO₂

Mt. Kasatochi
Alaska
9 August 2008

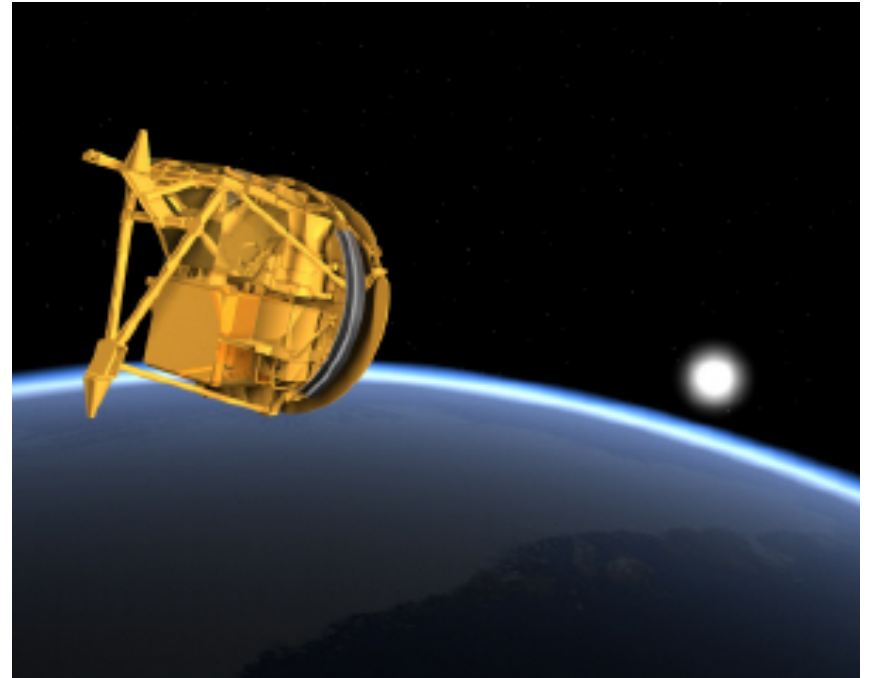


A priori SO₂ altitude = 10 km
Uncertainty = 2 km



Occultation Measurements

- Self-calibrating
- High vertical resolution
- Sparse global coverage



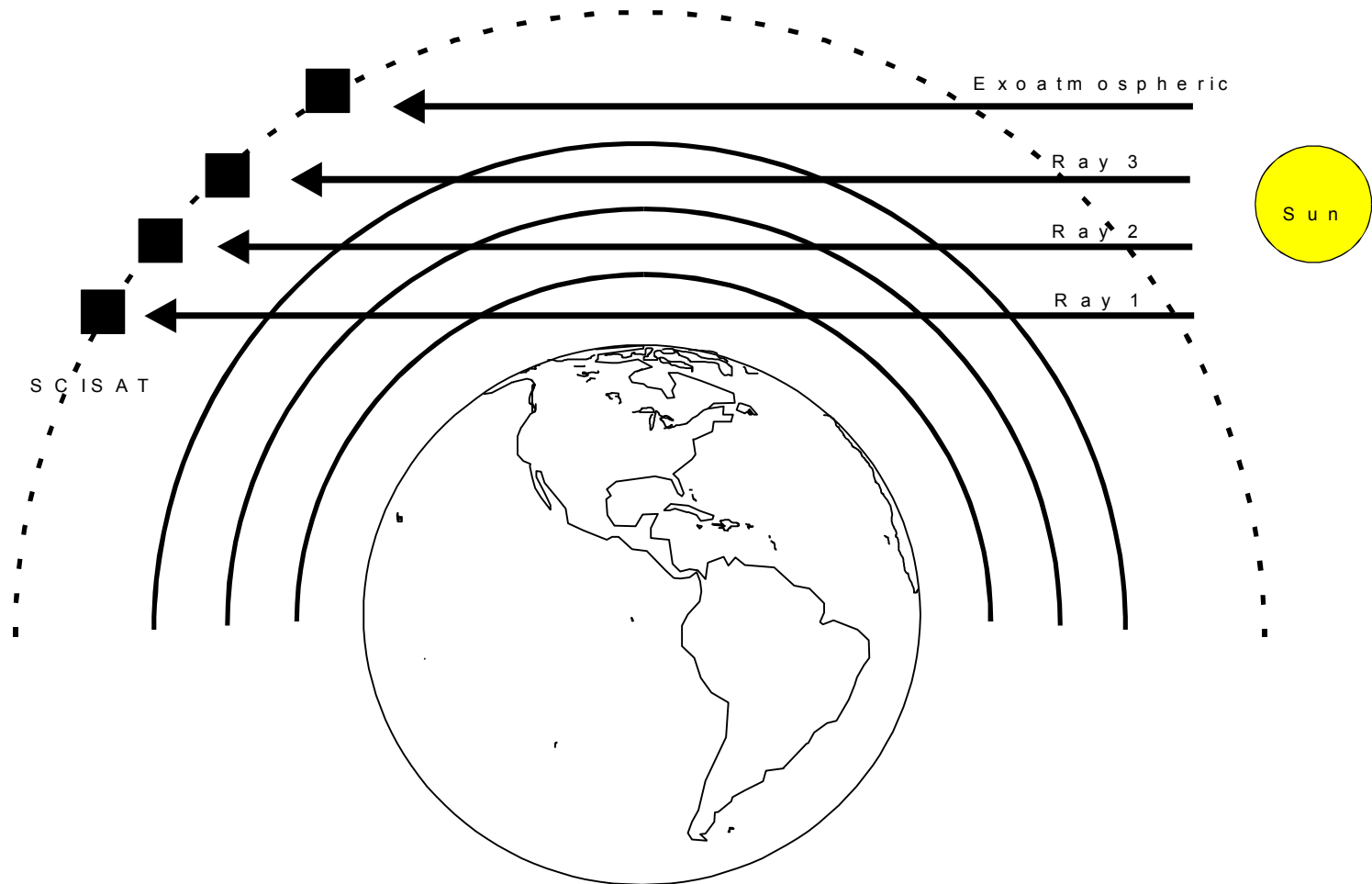
Occultation Instruments

Instrument	Dates	Spectral Region	Source
ATMOS (shuttle)	1985 – 1994	infrared	Sun
HALOE	1991 – 2005	NIR, infrared	Sun
SAGE I	1979 – 1981	UV, visible, NIR	Sun
SAGE II	1984 – 2005		
SAGE III	2001 – 2005		
SAGE III ISS	2015 –		
GPS technique	1995 –	radio	GPS
ILAS	1996 – 1997	NIR, infrared	Sun
ILAS II	2002 – 2003		
SCIAMACHY	2002 – 2012	UV – infrared	Sun and moon
GOMOS	2002 – 2012	UV, visible, NIR	Stars
ACE-FTS	2003 –	infrared	Sun
ACE-MAESTRO	2003 –	visible, NIR	Sun

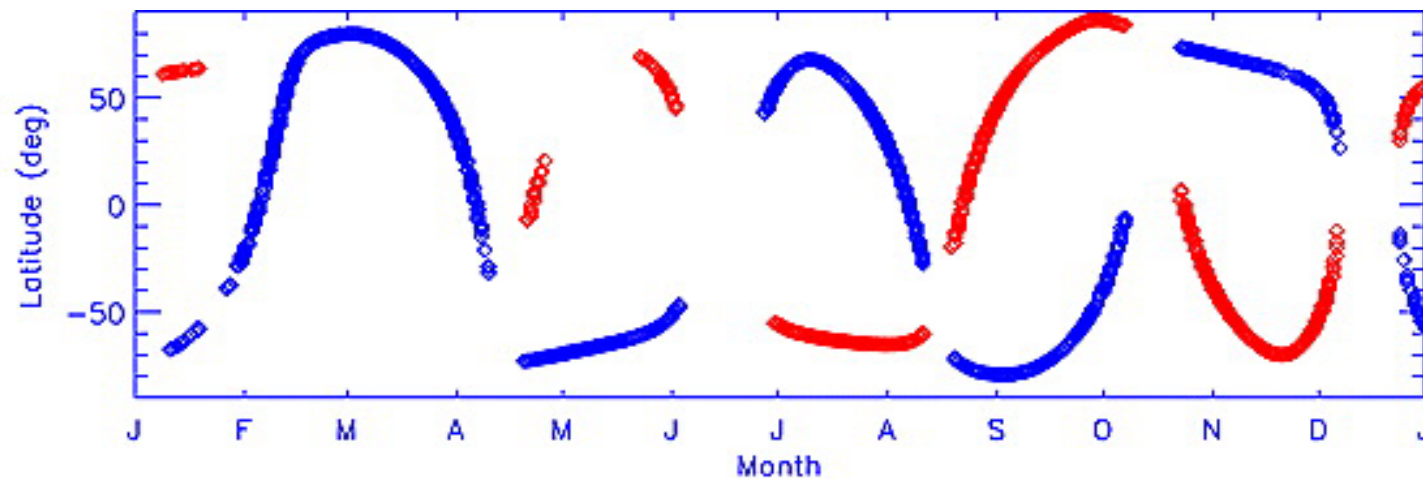
Atmospheric Chemistry Experiment

- ACE is on SCISAT satellite
- Launched August 12, 2003
- Two primary instruments on-board
 - **ACE-FTS** (Fourier Transform Spectrometer)
 - → **INFRARED**
 - **ACE-MAESTRO**
 - → **VISIBLE-NIR**
- ACE-FTS measures:
 - H₂O, O₃, N₂O, NO, NO₂, HNO₃, N₂O₅, H₂O₂, HO₂NO₂, N₂, HCl, HF, ClONO₂, CFC-11, CFC-12, CFC-113, COF₂, COCl₂, COFCl, CF₄, SF₆, CH₃Cl, CCl₄, HCFC-22, HCFC-141b, HCFC-142b, CO, CH₄, CH₃OH, H₂CO, HCOOH, C₂H₂, C₂H₄, C₂H₆, OCS, HCN, ClO, acetone, PAN, aerosols

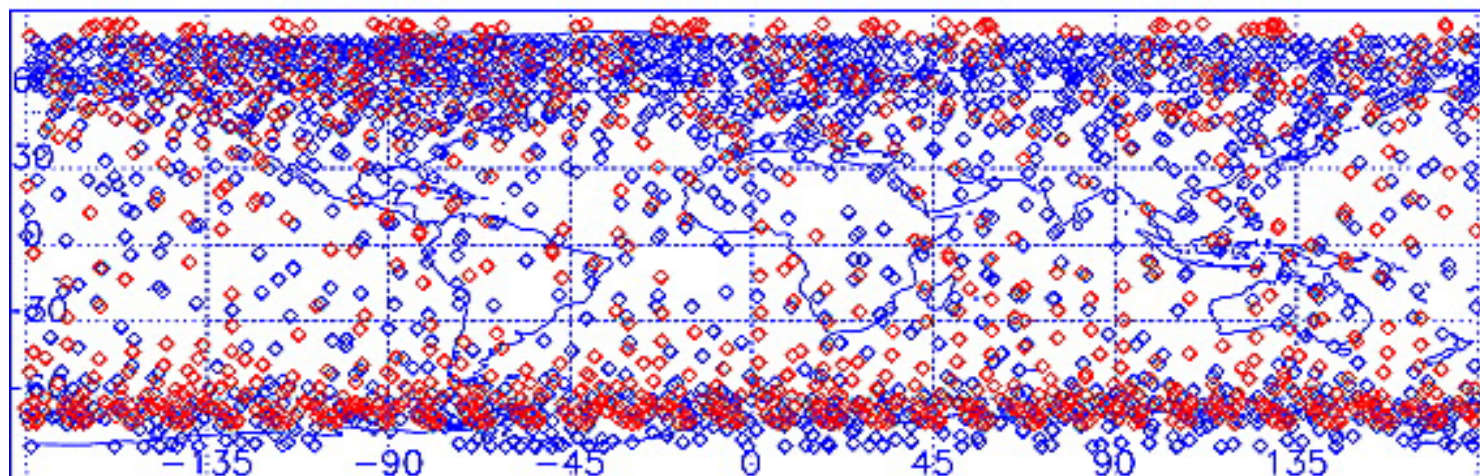
Solar Occultation Measurements



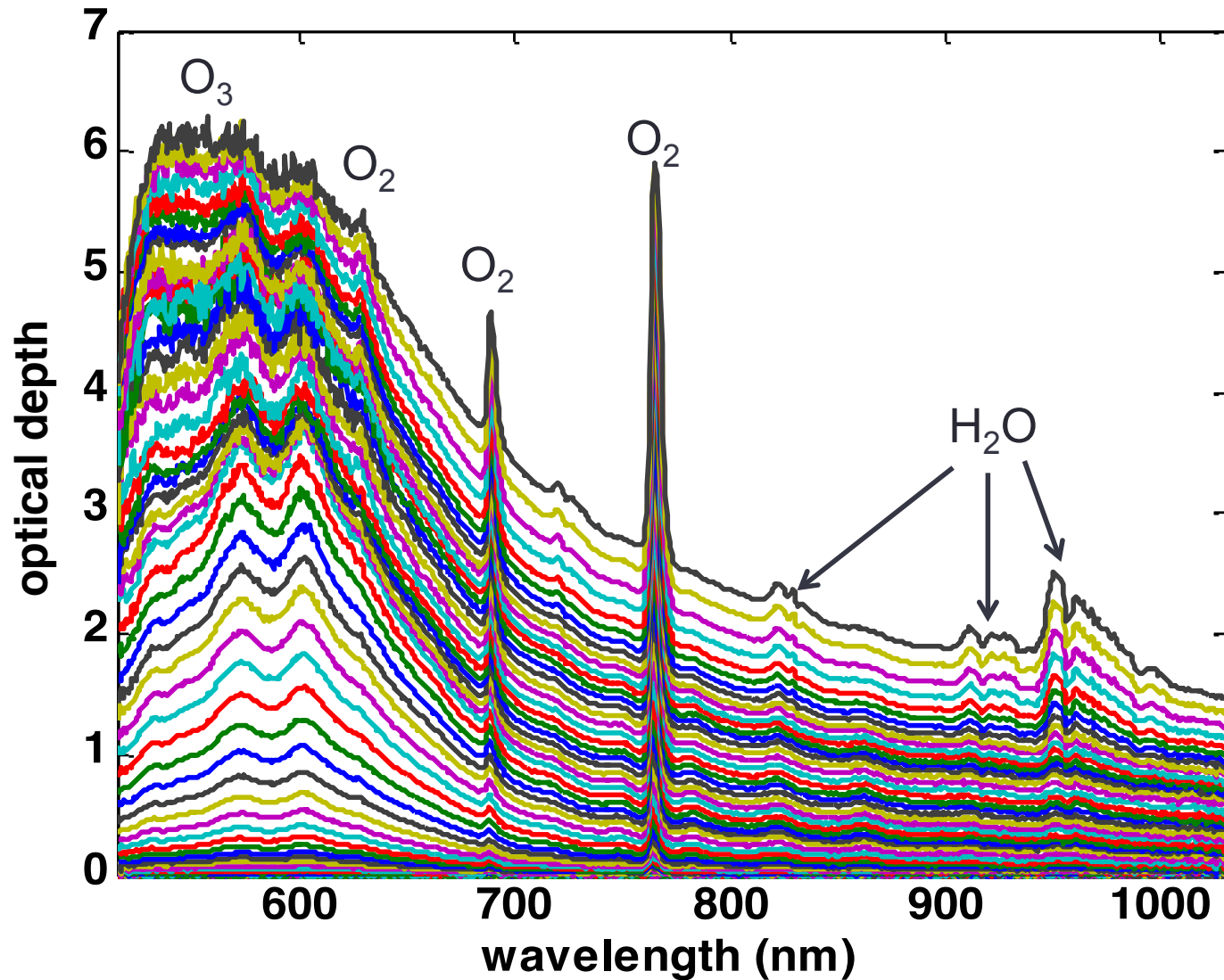
ACE Occultation Coverage: 2004



SUNRISE
SUNSET



MAESTRO Optical Depth Spectra



Occultation Retrievals

- Traditional method: “onion peeling”
- Newer approach: global fitting
 - Simultaneous fitting of every spectrum in an occultation
 - Arrange all spectra into one giant measurement vector

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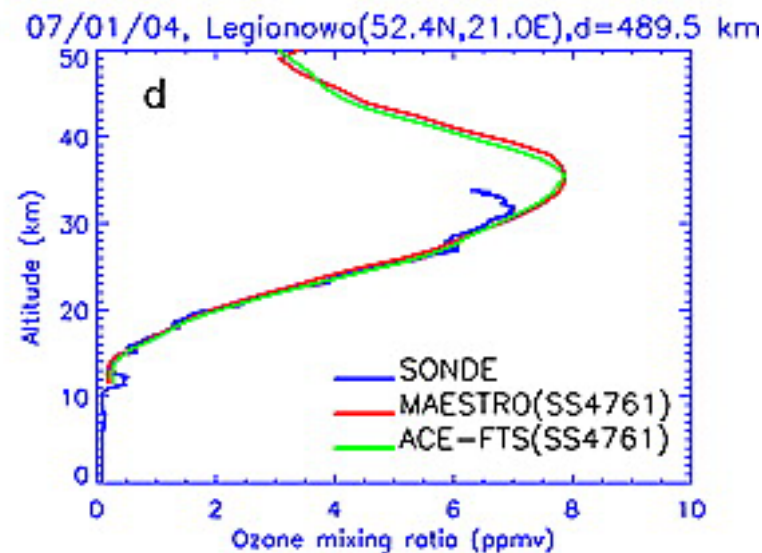
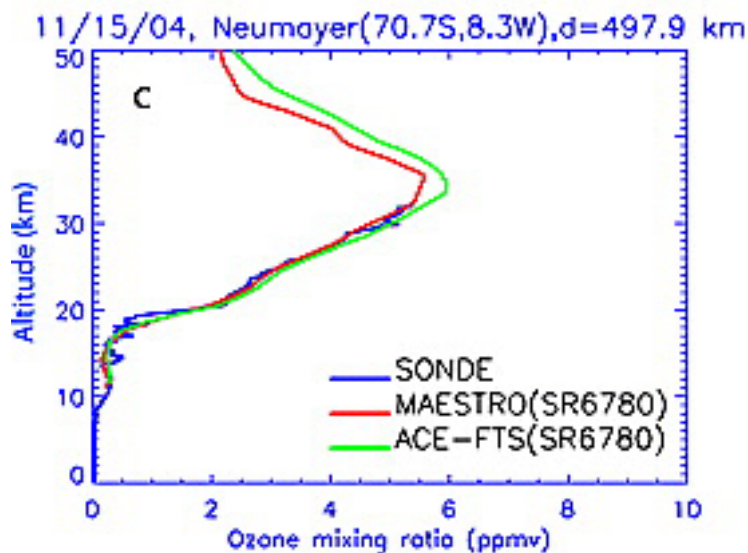
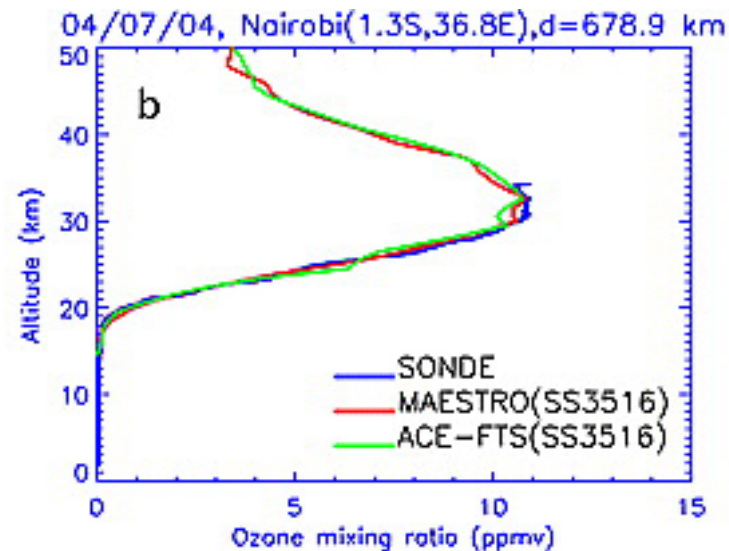
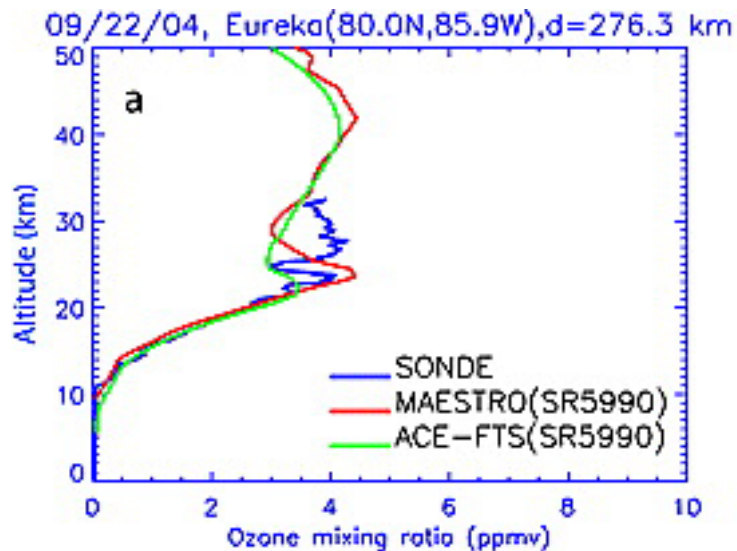
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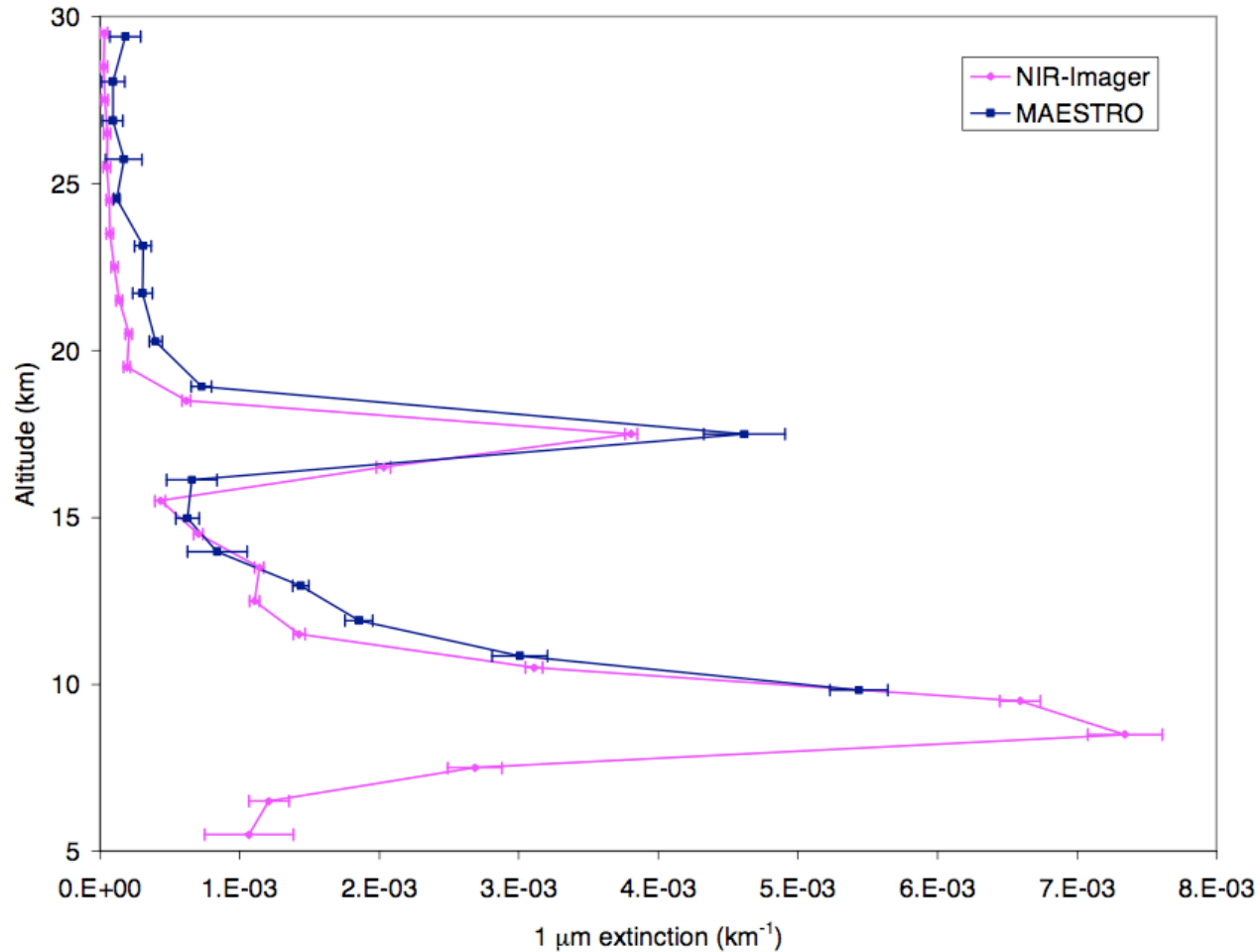
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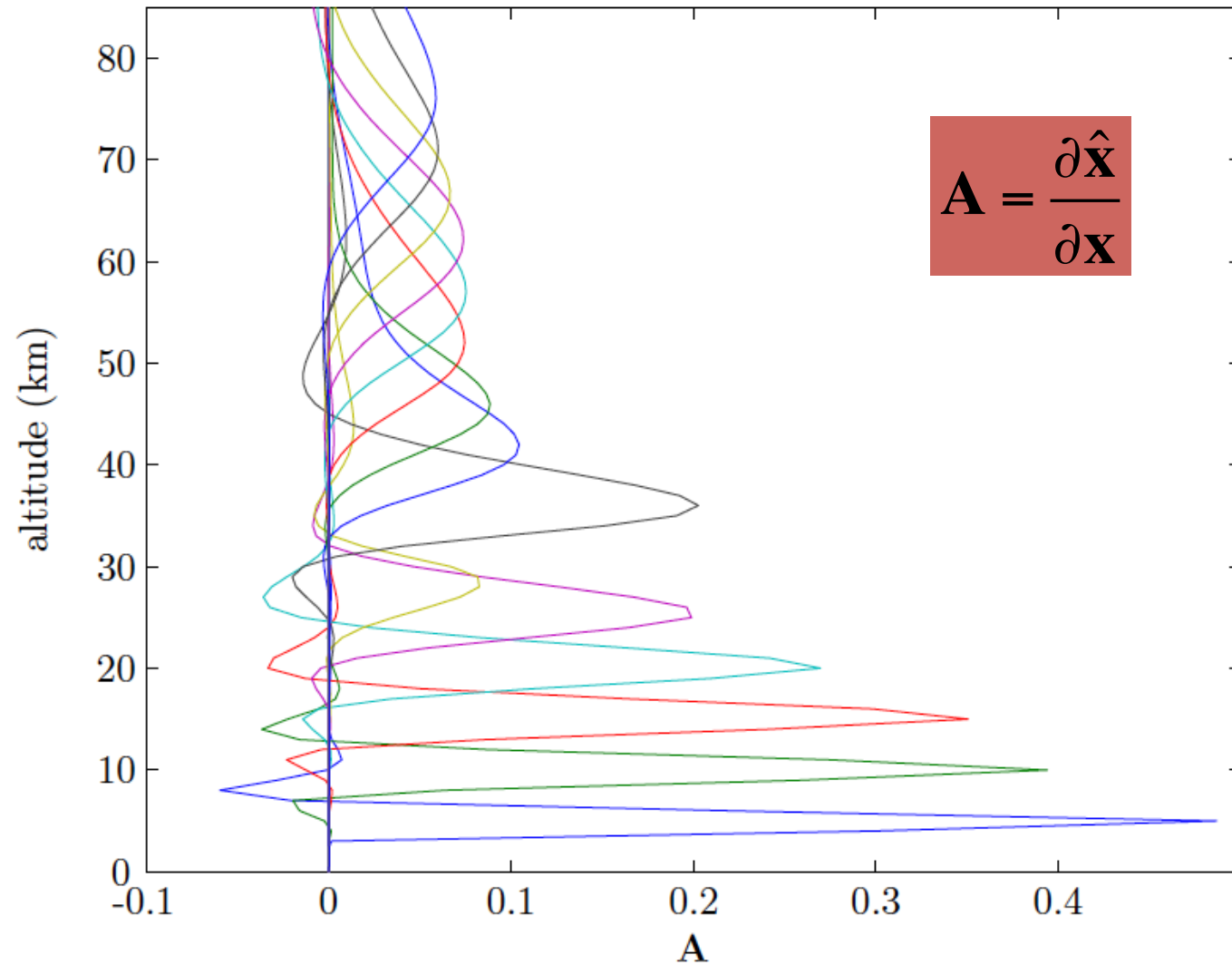
ACE-MAESTRO: Ozone Profiles



ACE-MAESTRO: Mt. Kasatochi Aerosols

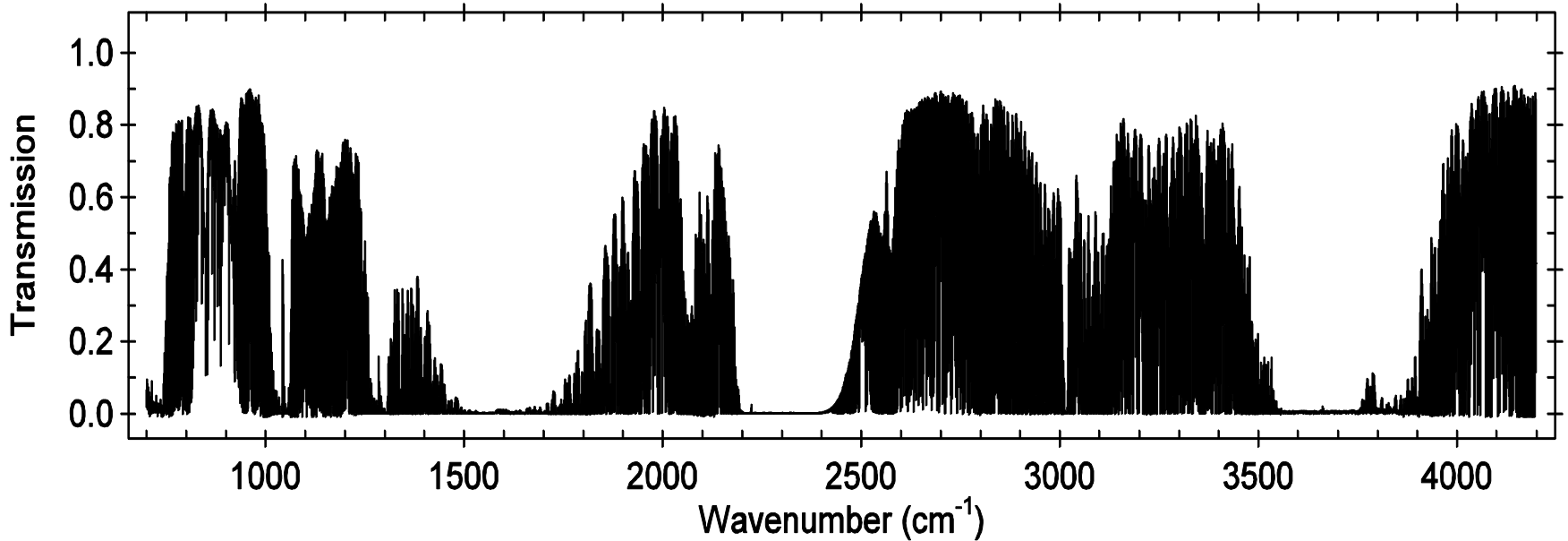


Density Averaging Kernels from O₂ Bands

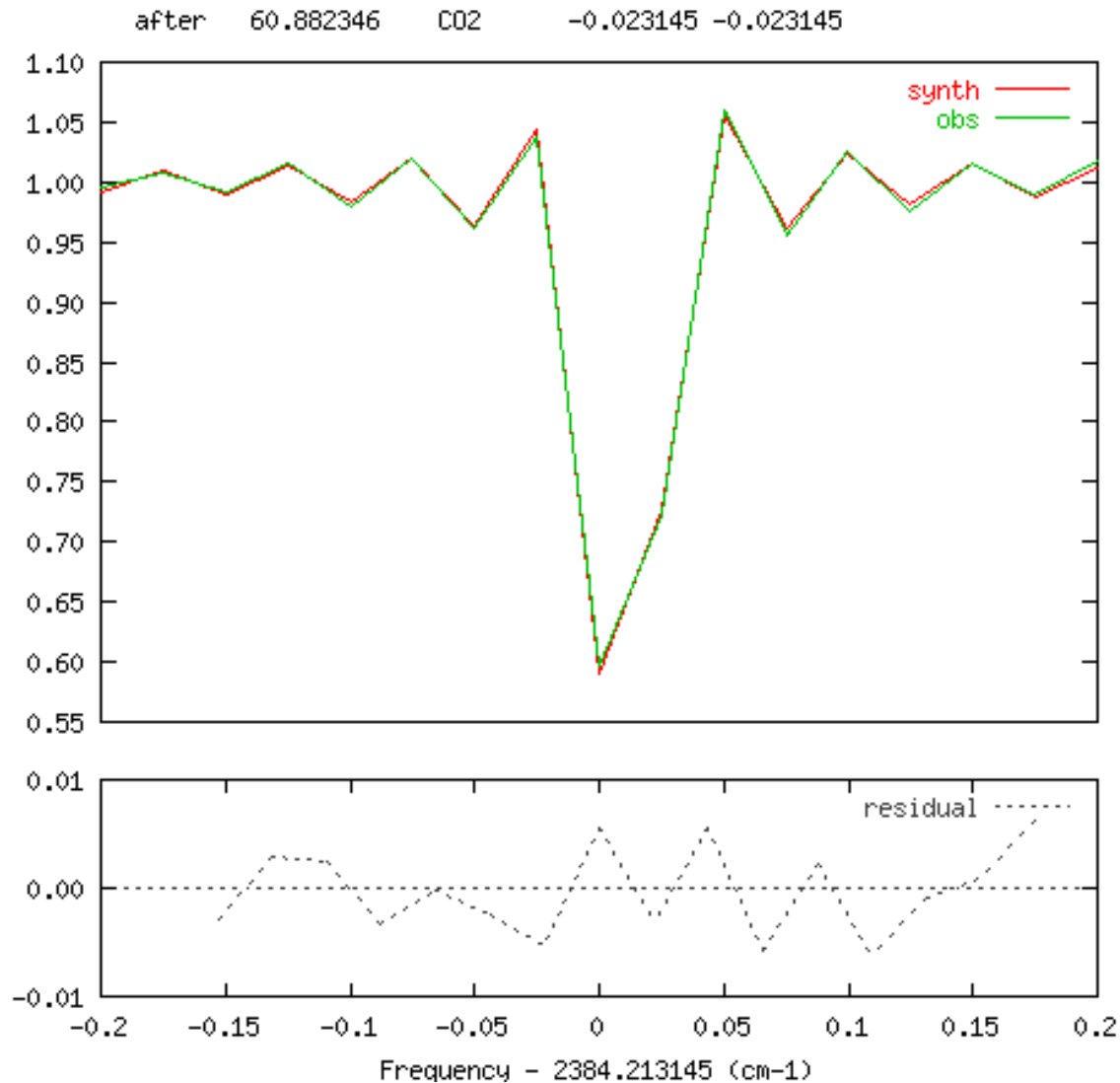


ACE-FTS Spectra

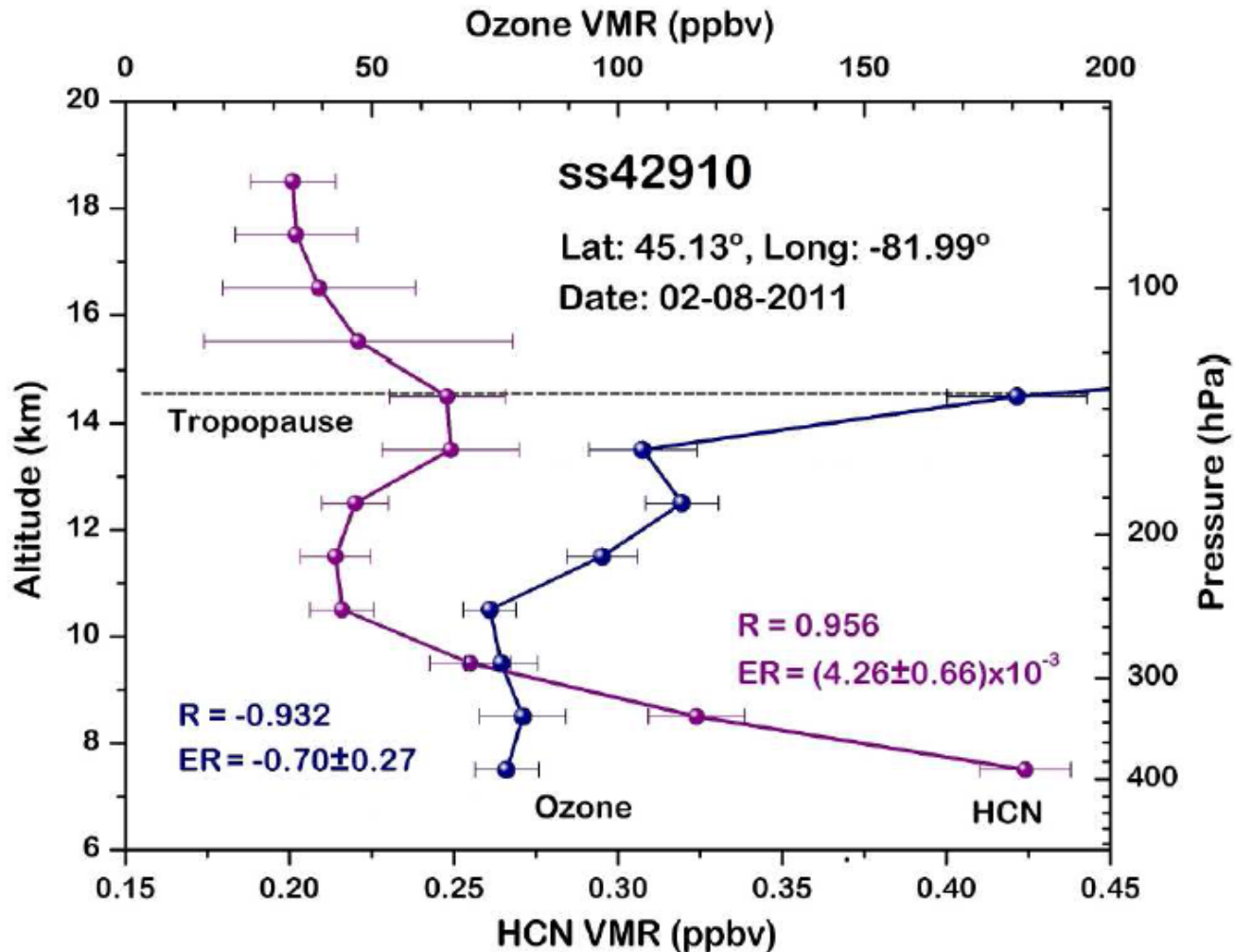
ACE Tropospheric Spectrum (8-12 km tropics)



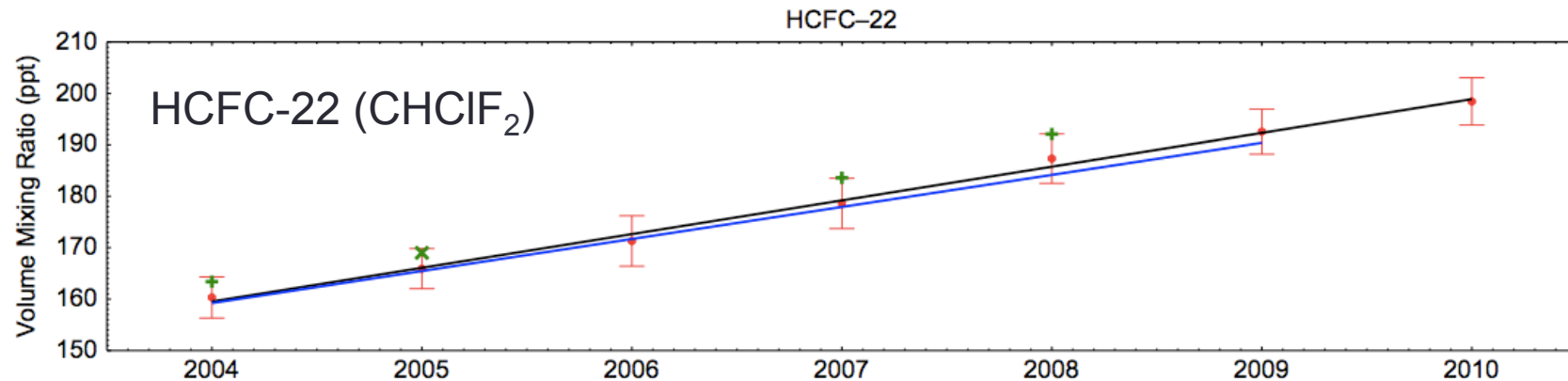
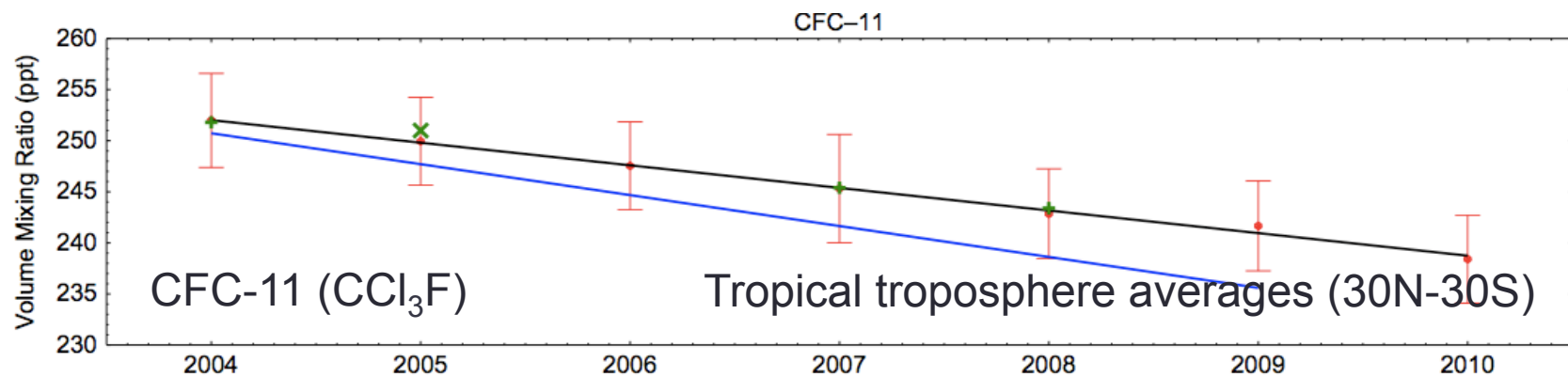
ACE-FTS CO₂ line (near 61 km)



ACE-FTS: Canadian Biomass Burning



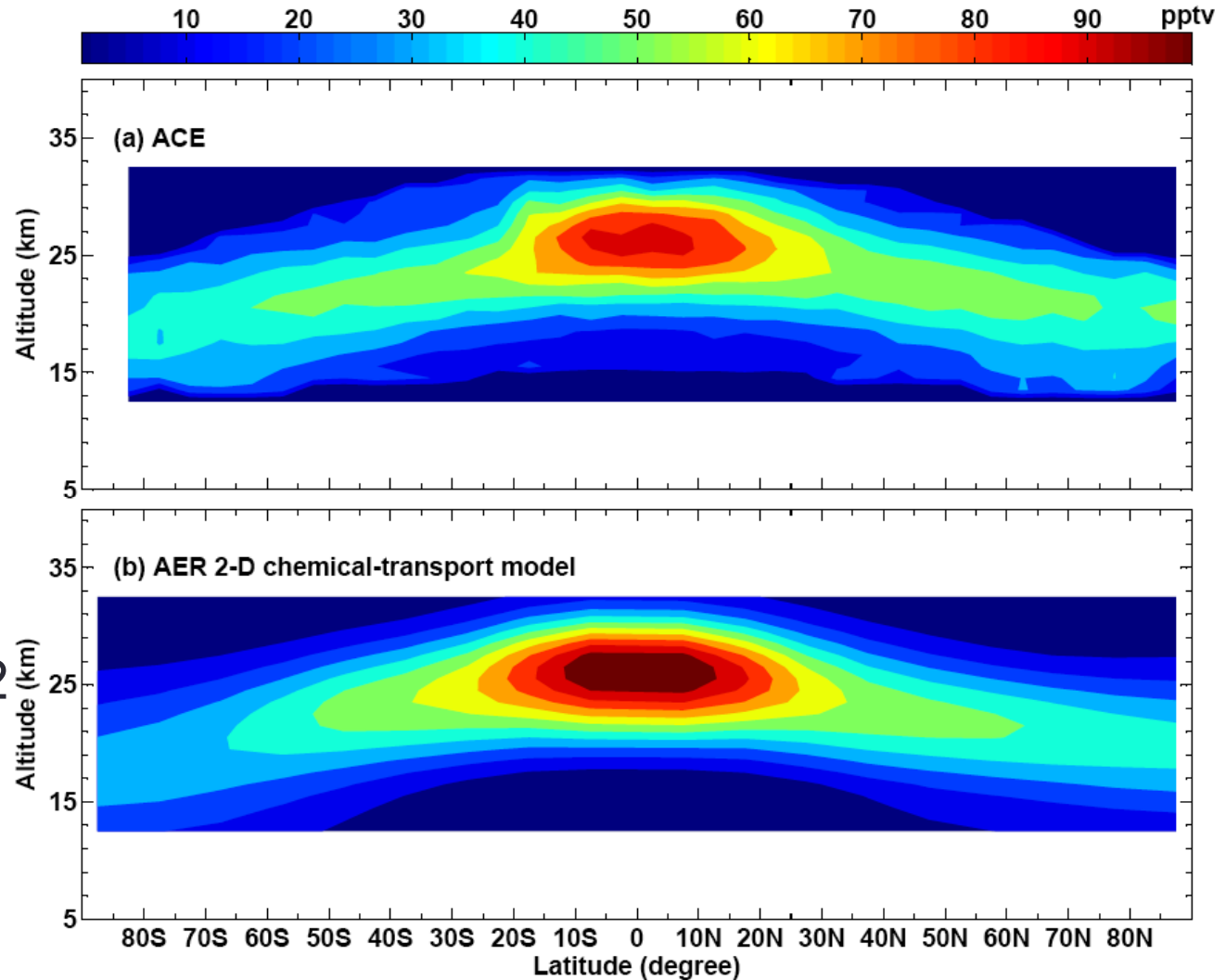
Halogen-containing Species Trends



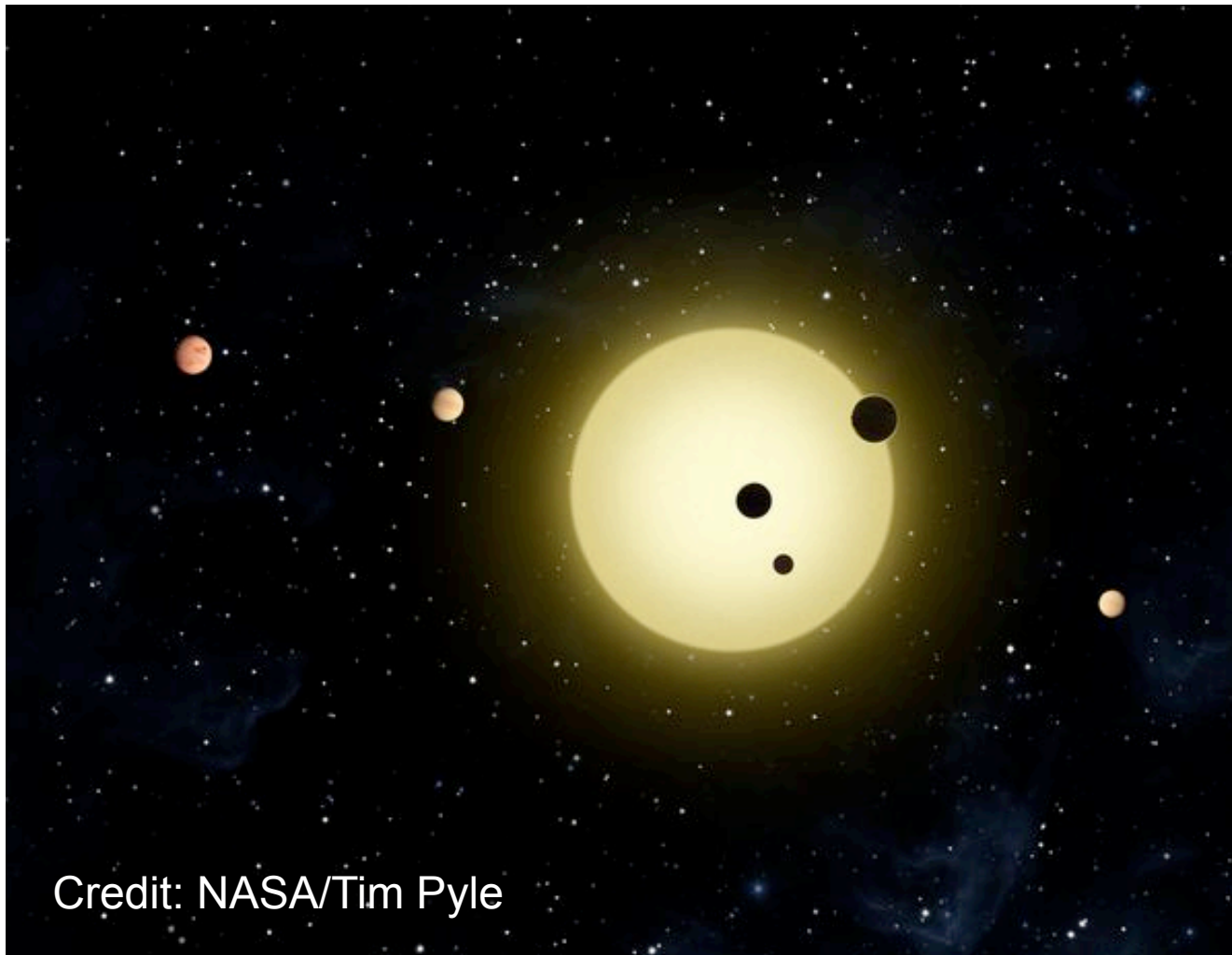
ACE-FTS meas. (~7-16 km)
SLIMCAT model (~7-16 km)
AGAGE meas. (surface)

Distribution of COClF

- Carbonyl chlorofluoride is a product of chlorofluorocarbon (CFC-11 mainly) decomposition
- Previously studied by aircraft instruments (5 - 12 km)
- First global picture obtained from ACE-FTS



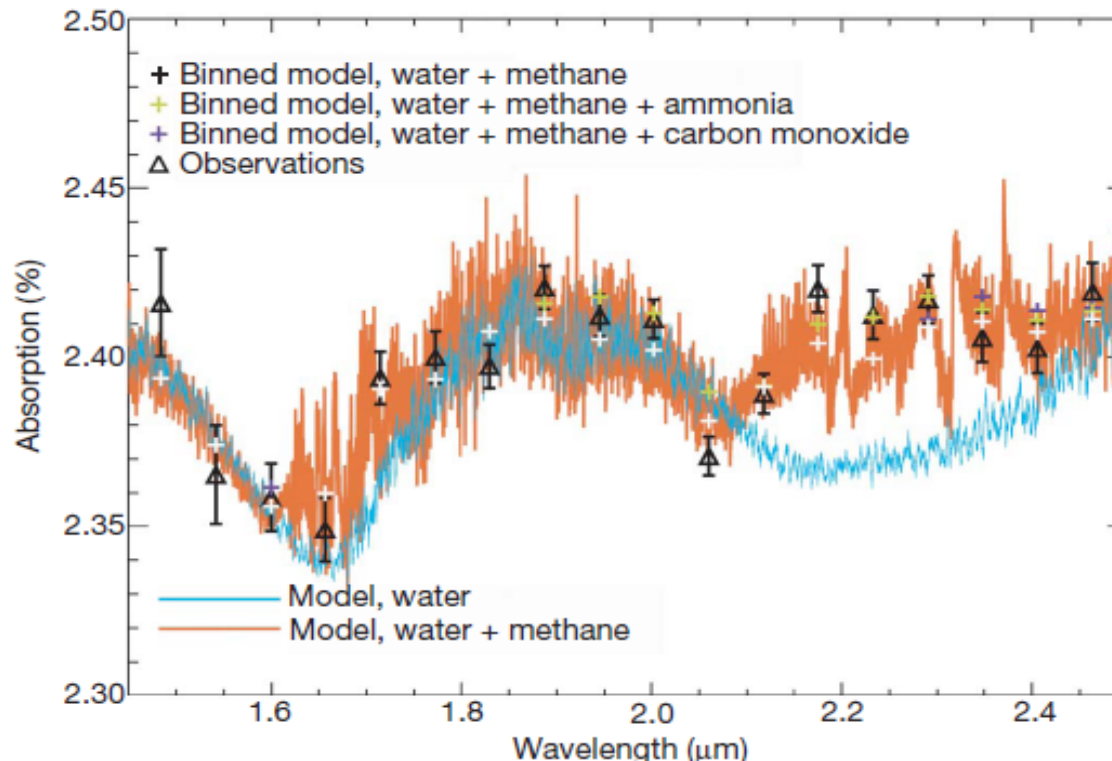
Exoplanet atmospheres from occultation?



LETTERS

The presence of methane in the atmosphere of an extrasolar planet

Mark R. Swain^{1*}, Gautam Vasisht^{1*} & Giovanna Tinetti^{2*}



The End