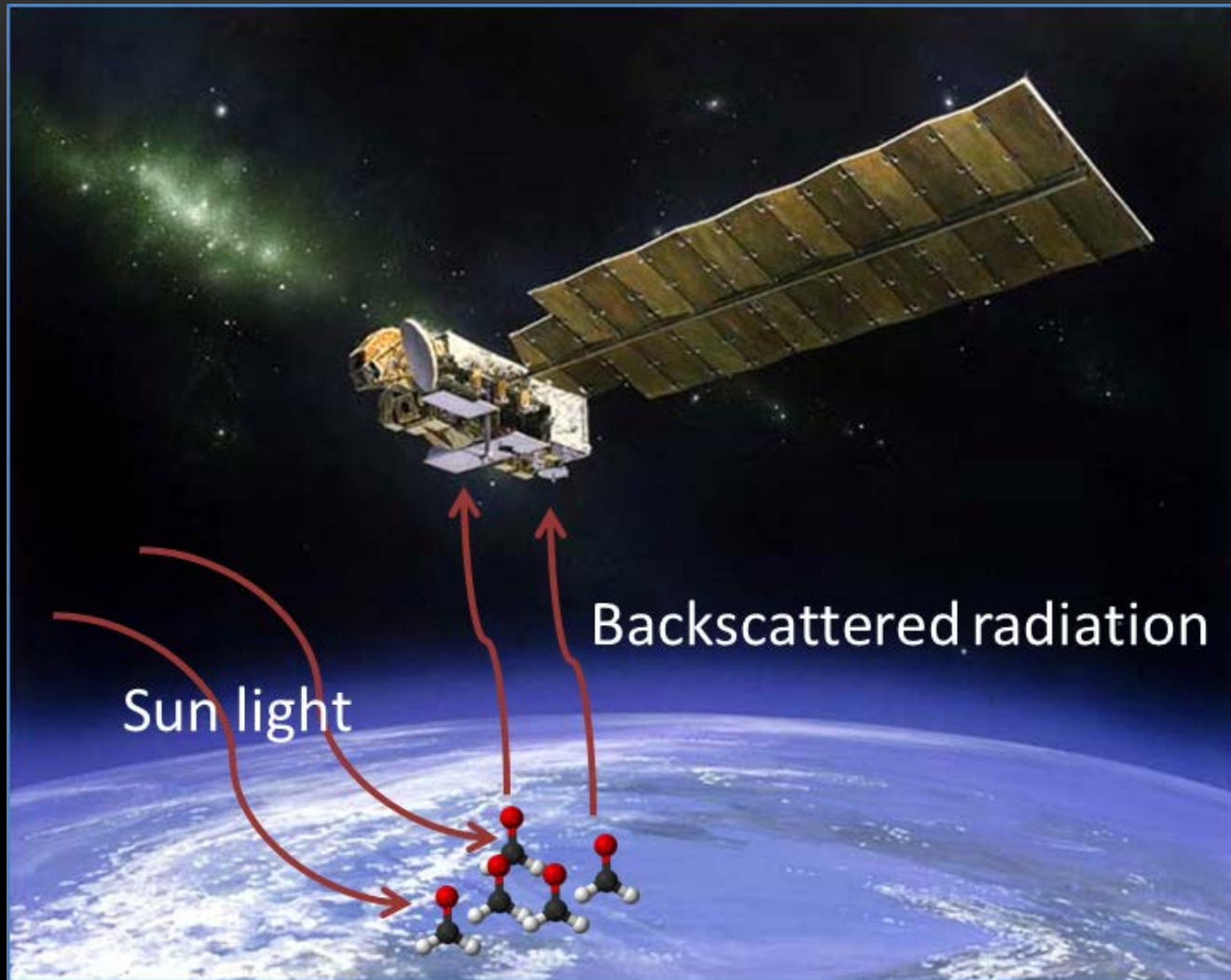


UV/VIS BACKSCATTERED SUN LIGHT RETRIEVALS FROM SPACE BORN PLATFORMS

G. Gonzalez Abad

With a lot of help from K. Chance, X. Liu, C. Miller and others



Sun light

Backscattered radiation

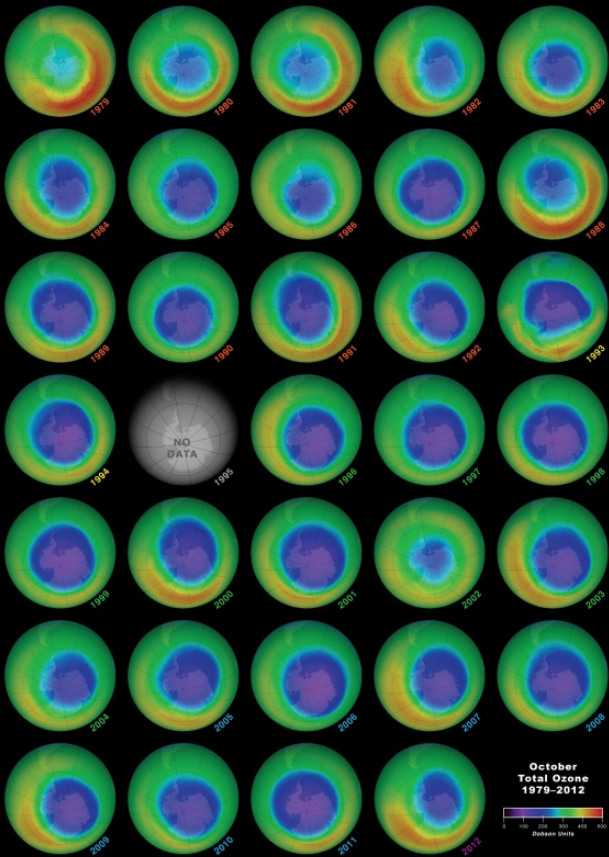
OVERVIEW

- A bit of history on UV-VIS backscatter spectrometers
 - Past, present and future instruments (as far as I know)
 - TOMS vs. OMI
 - GOME vs. OMI
 - The GEO constellation
 - Slant column retrievals
 - The equations: BOAS
 - To have in mind (solar spectrum, cross sections, ring, undersampling, ...)
 - Examples
 - Slant column to vertical column
 - Equations
 - Vertical columns: examples from OMI
-

HISTORY: THE DISCOVERY OF THE OZONE HOLE FUELLED THE INTEREST IN THE ATMOSPHERIC CHEMISTRY AND COMPOSITION DURING THE 1970S

USA

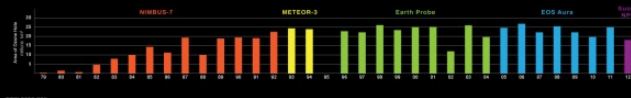
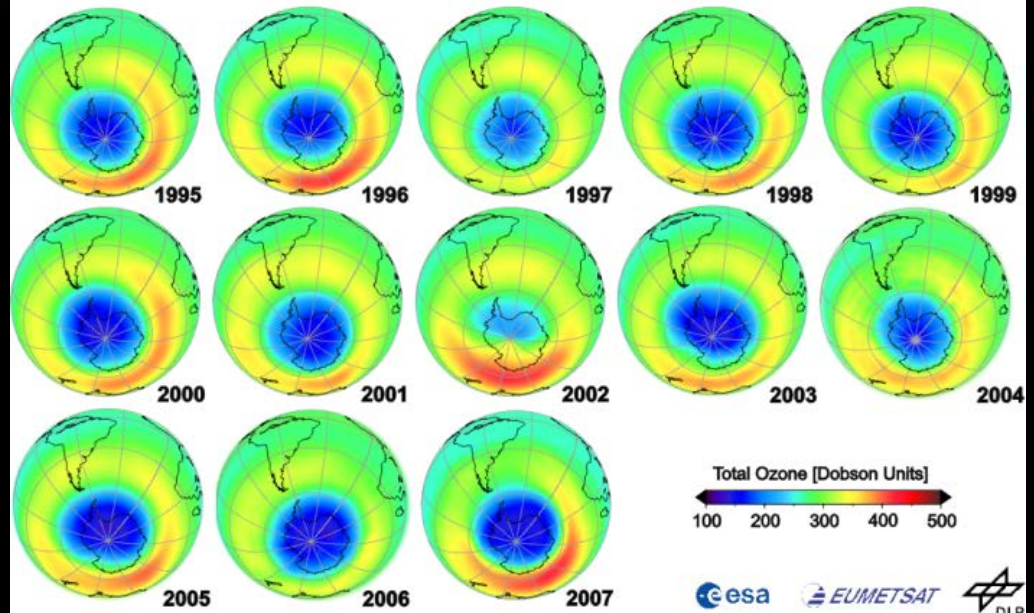
The Ozone Hole Over 30 Years of Satellite Observations



Europe

Monitoring the Antarctic Ozone Hole by GOME, SCIAMACHY and GOME-2

Total Ozone Monthly Mean, September 1995 - 2007



HISTORY & PRESENT: USA

Instrument	Satellite (Agency)	Period
BUV	Nimbus-4 (NASA)	1970-1980
SBUV	Nimbus-7 (NASA)	1978-1993
TOMS	Nimbus-7 (NASA)	1978
	Meteor-3 (NASA) (Russian)	1993
	Earth Probe (NASA)	1996-2007
SBUV/2	NOAA 9,10,11,13,14,17,16,18,19 (NOAA)	1984 – present
OMI	AURA (NASA/KNMI)	2004-present
OMPS	SUOMI-NPP (NOAA)	2011-present

HISTORY & PRESENT: EUROPE & CHINA

Instrument	Satellite (Agency)	Period
GOME	ERS-2 (ESA)	1995-2011
SCIAMACHY	ENVISAT (ESA)	2002-2012
OMI	AURA (KNMI/NASA)	2004-present
GOME2	Metop-A (EUMETSAT) Metop-B (EUMETSAT)	2006-present 2012-present
TOU/SBUS	FY-3A (NRSCC) FY-3B (NRSCC) FY-3C (NRSCC)	2008-present



FUTURE INSTRUMENTS

Instrument	Satellite (Agency)	Orbit
OMPS	JPSS1 (NOAA)	LEO
	JPSS2 (NOAA)	LEO
TROPOMI	SENTINEL 5P (ESA)	LEO
GOME-2	Metop-C (EUMETSAT)	LEO
UVAS	Seosat-Ingenio (CDTI)	LEO
UVN	SENTINEL 4 (ESA)	GEO
GEMS	MP-GEOSAT (KORA)	GEO
TEMPO	TBD (NASA/SMITHSONIAN)	GEO

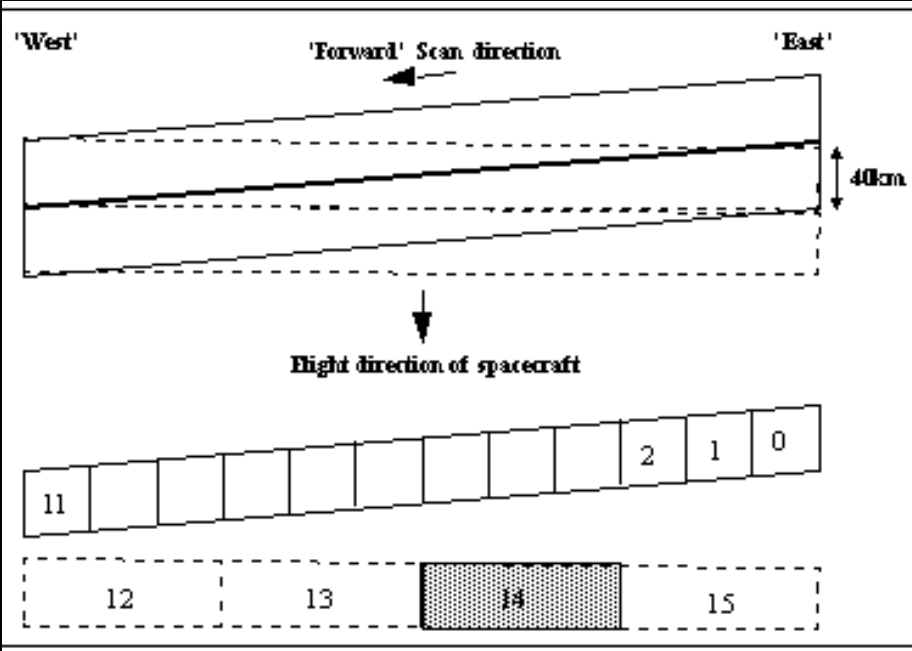
TOMS VS. OMI

	TOMS (Nimbus-7)	OMI
Spectral coverage	312.5 nm 317.5 nm 331.3 nm 339.9 nm 360.0 nm 380.0 nm	UV1: 270-314 nm UV2: 306-380 nm VIS: 350-500 nm
Bandwidth, resolution	1.1 nm	0.63 – 0.45 nm
Ground pixel size (nadir)	50 x 50 km ²	13 x 24 km ²
Global coverage	Daily	Daily

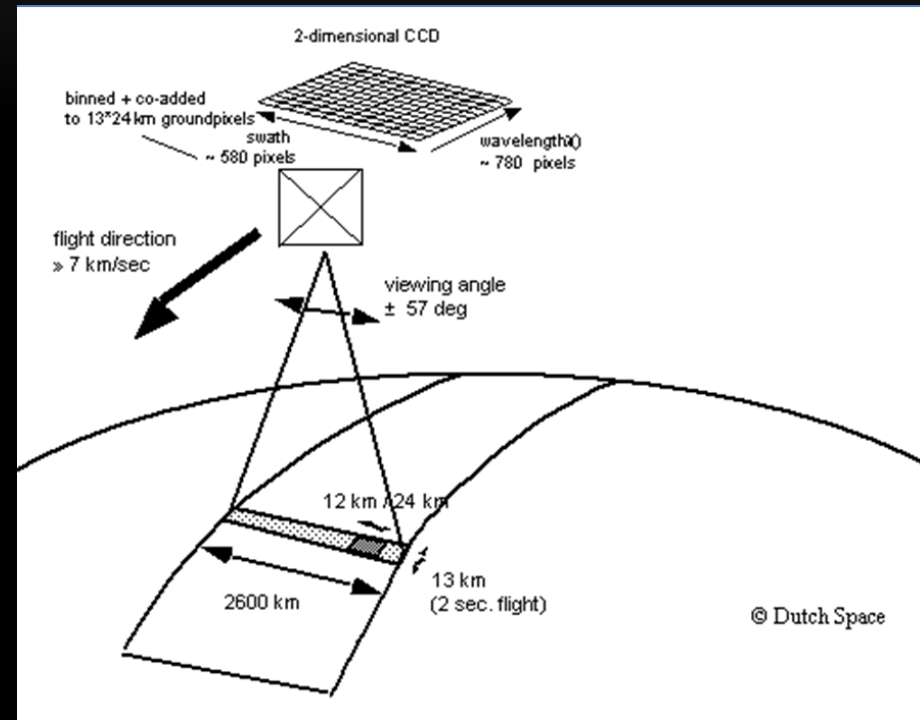
GOME VS. OMI

GOME

OMI

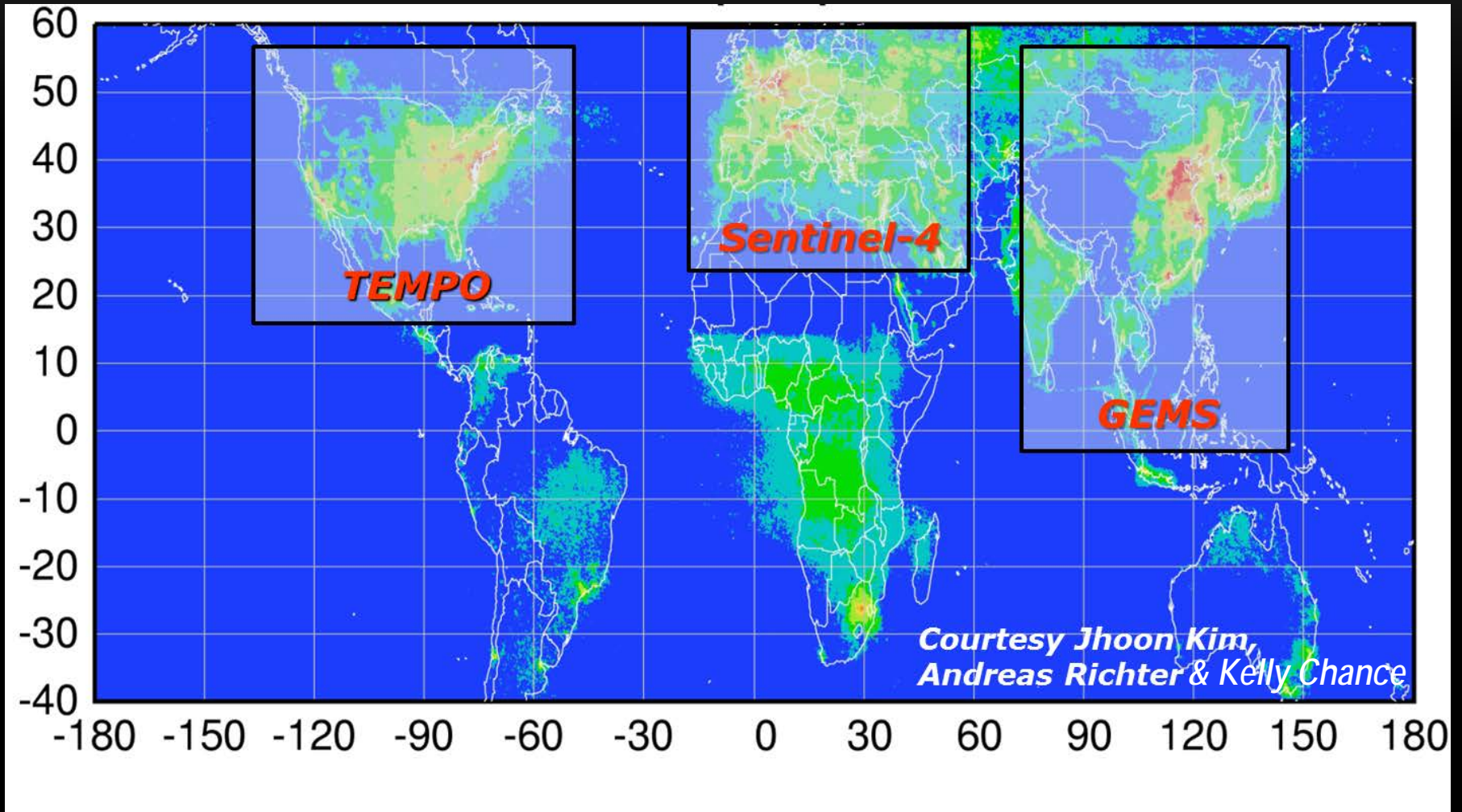


Photodiode detectors 1024 pixels

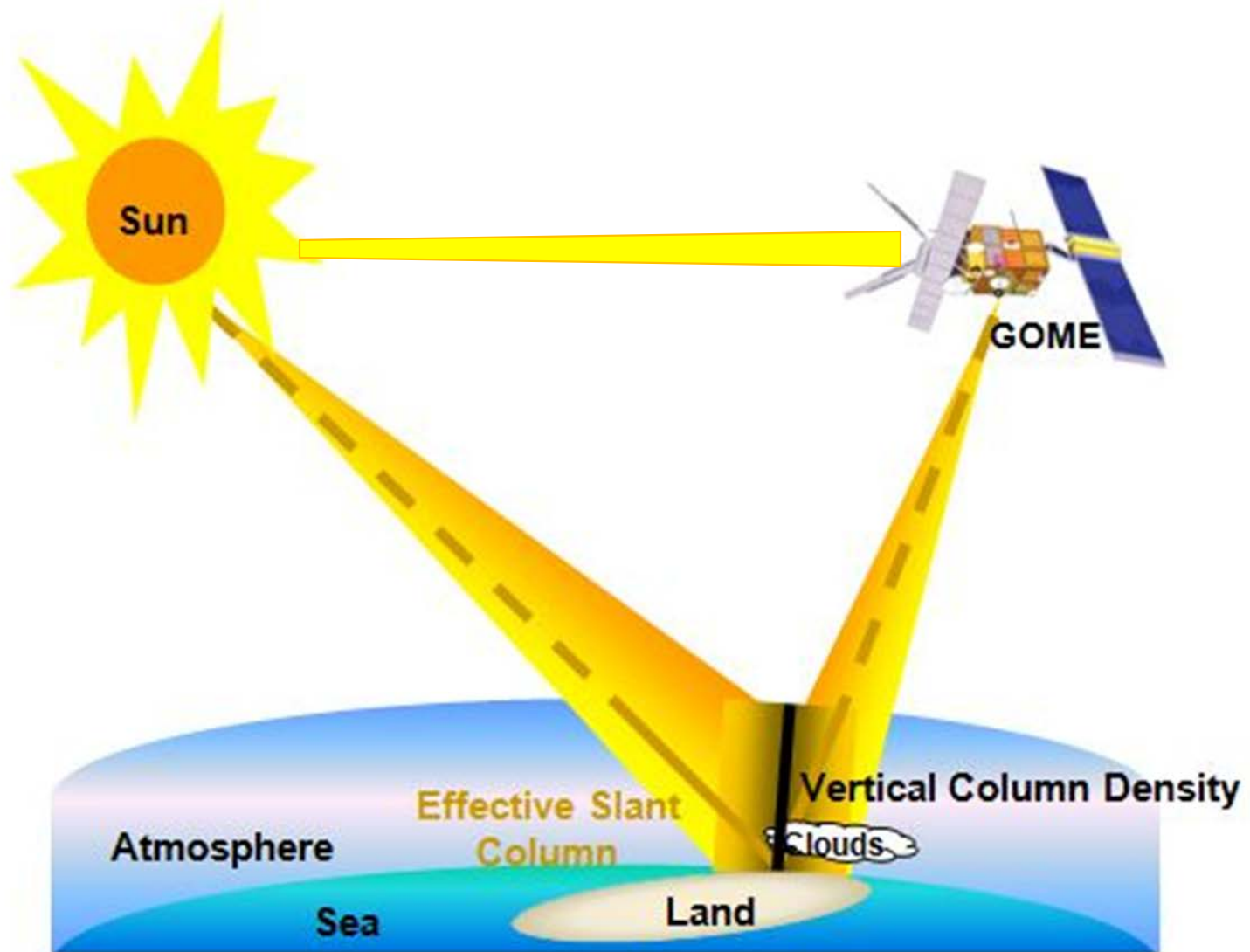


2D CCD detector

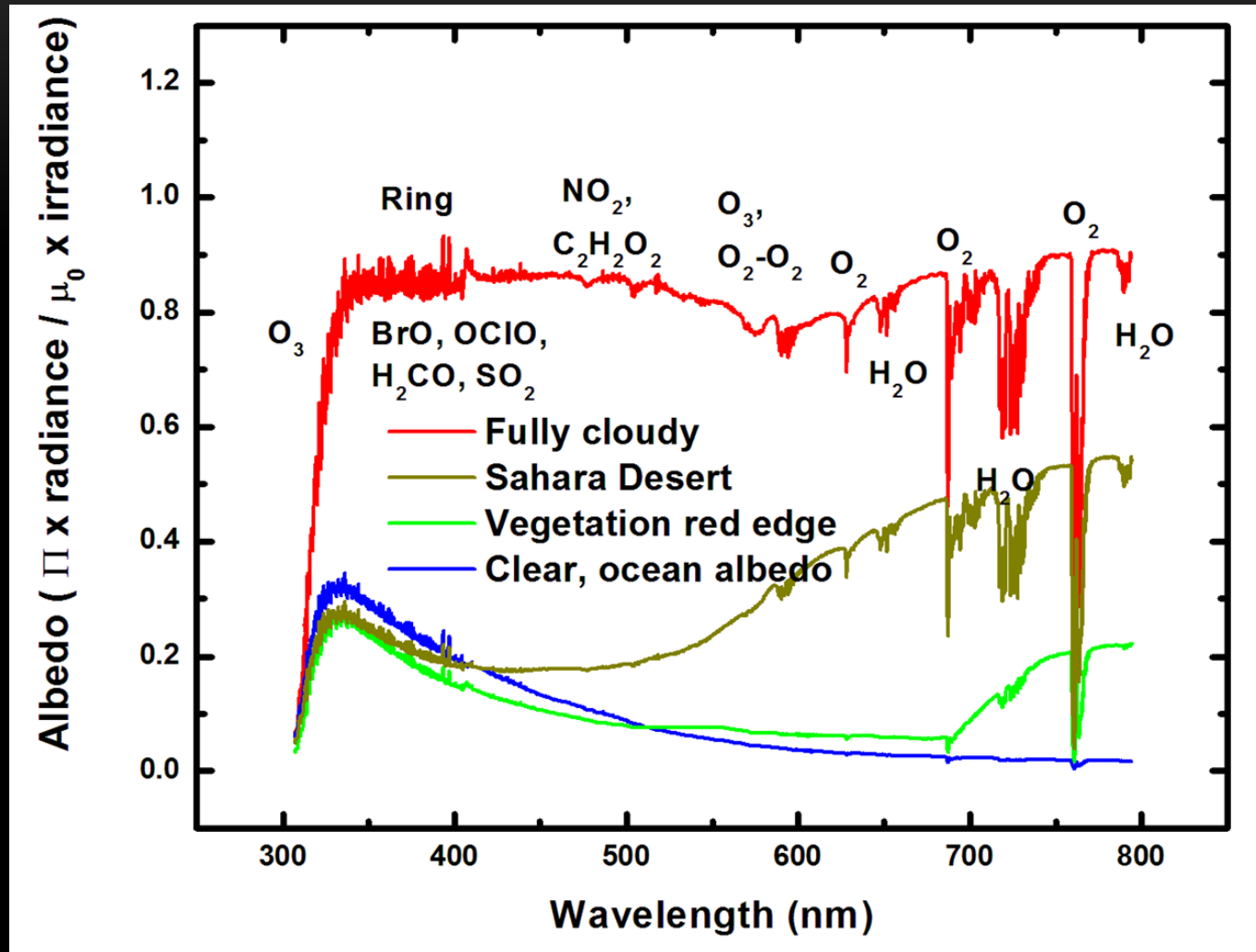
THE GEO CONSTELLATION



SLANT COLUMN RETRIEVAL



SLANT COLUMN RETRIEVAL



SLANT COLUMN RETRIEVAL: BOAS

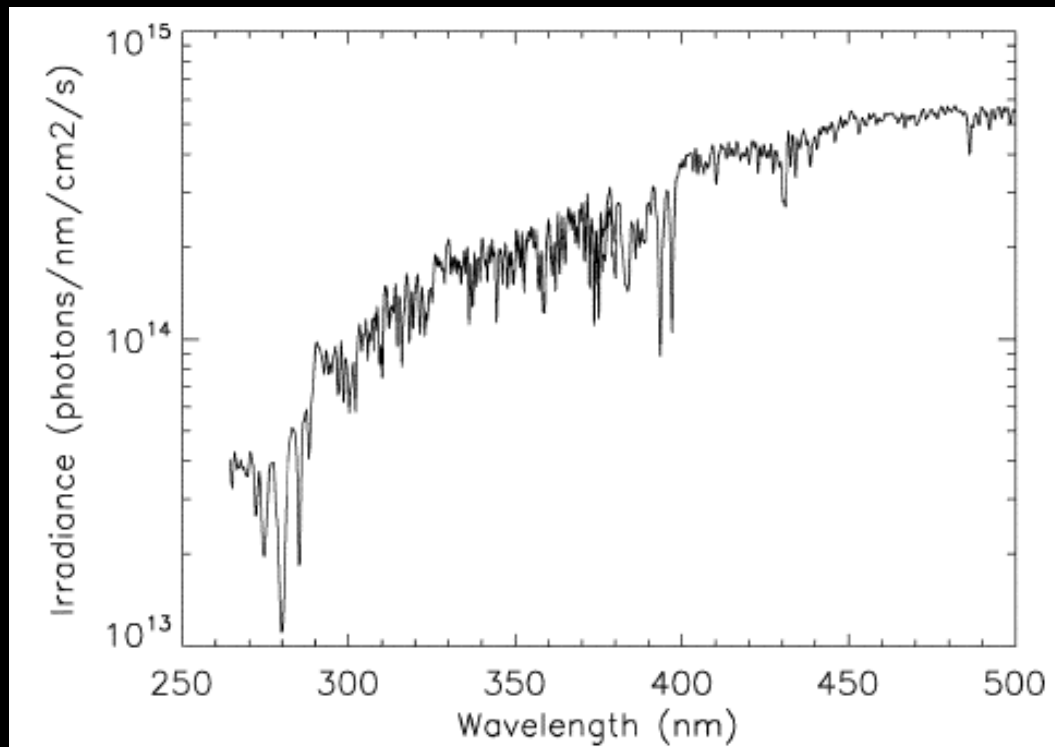
$$\chi^2 = \sum_i (y_i - f(\lambda_i, \beta))^2$$

$$I = \left[(aI_o + \sum_i \alpha_i X_i) e^{-\sum_j \alpha_j X_j} + \sum_k \alpha_k X_k \right] \text{ScalPoly} + \text{BasePoly}$$

$$\epsilon = \beta - \hat{\beta}$$

SLANT COLUMN RETRIEVAL: IRRADIANCE

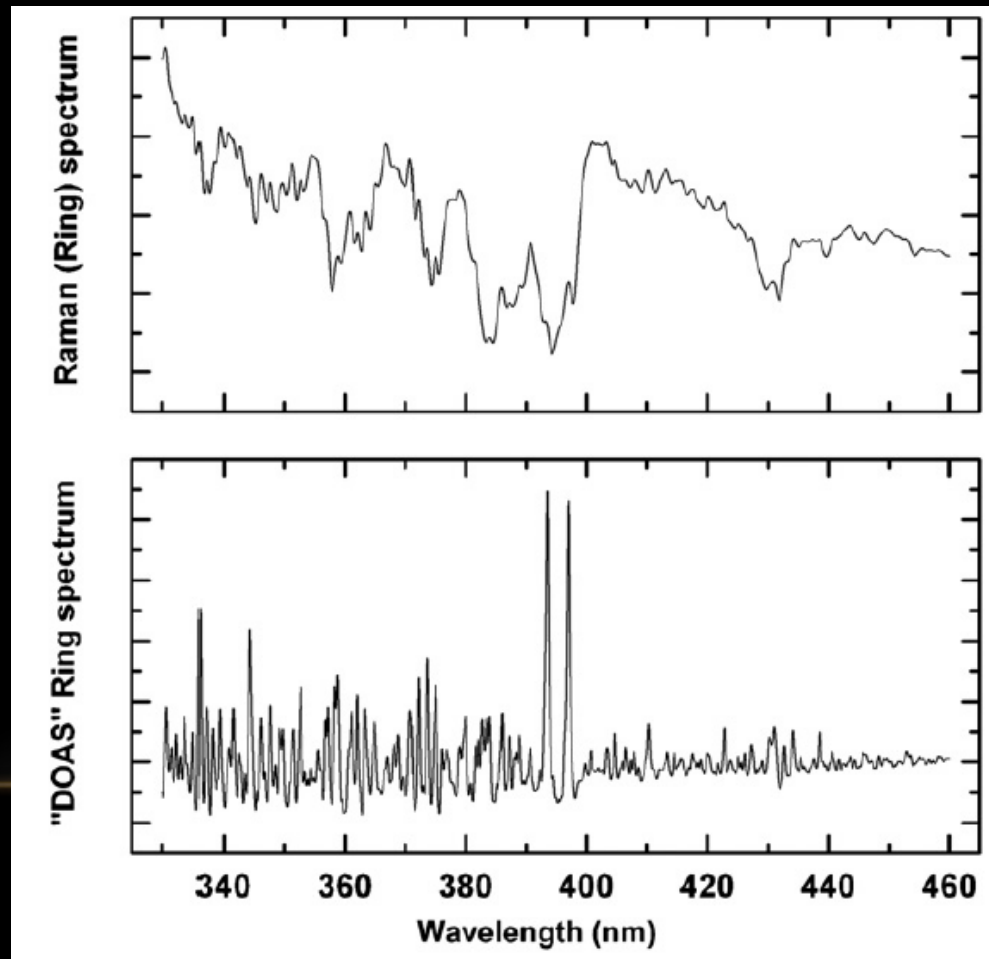
$$I = \left[(aI_0) + \sum_i \alpha_i X_i \right] e^{-\sum_j \alpha_j X_j} + \sum_k \alpha_k X_k \quad \text{ScalPoly} + \text{BasePoly}$$



OMI irradiance, Dobber at al. 2006

SLANT COLUMN RETRIEVAL: RING SPECTRUM

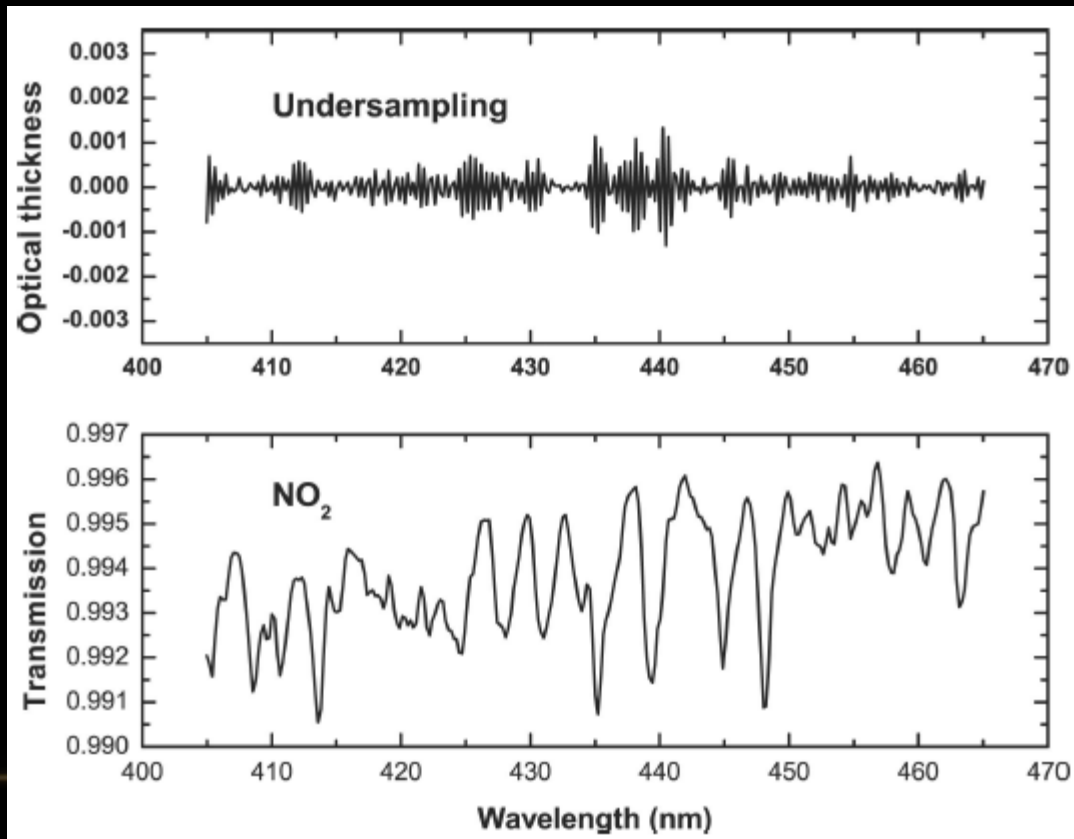
$$I = \left[(aI_o + \sum_i \alpha_i X_i) e^{-\sum_j \alpha_j X_j} + \sum_k \alpha_k X_k \right] ScalPoly + BasePoly$$



Chance and Kurucz
2010

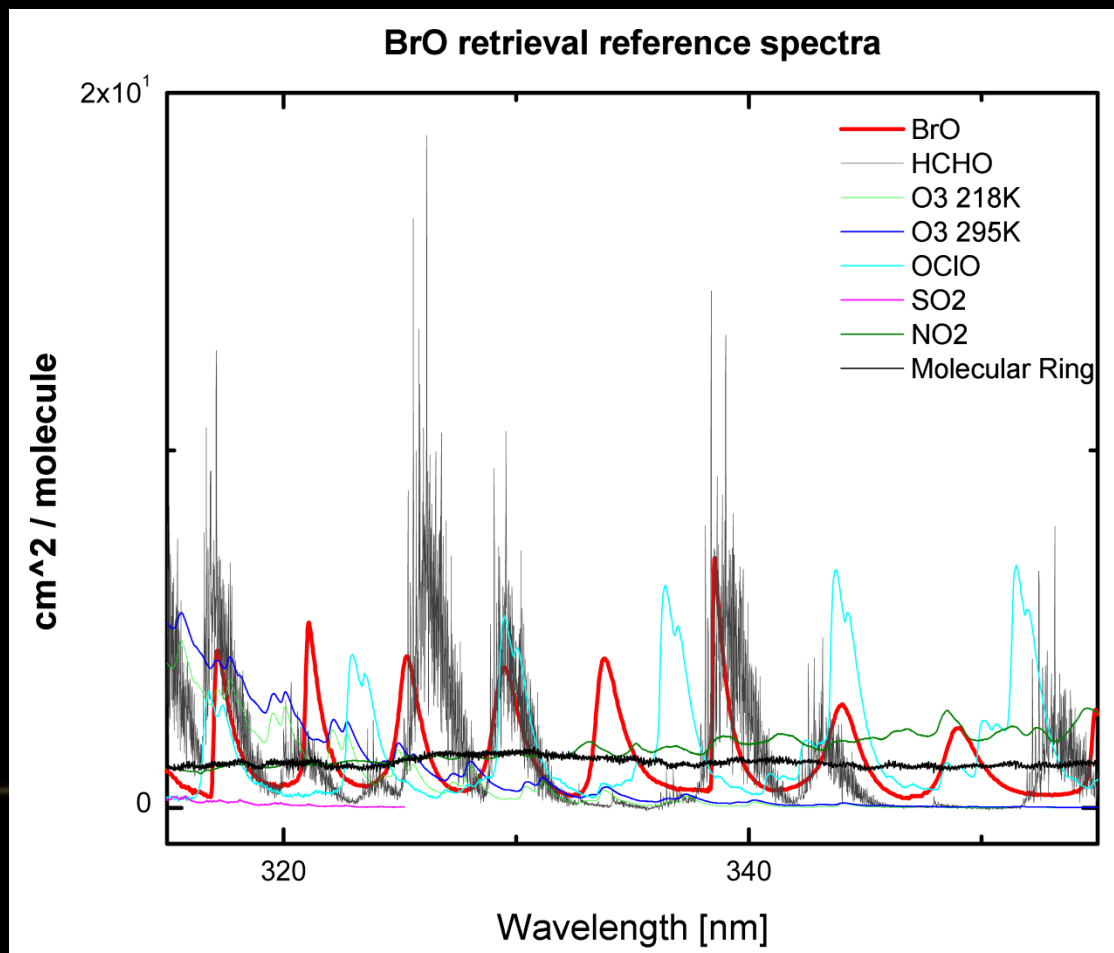
SLANT COLUMN RETRIEVAL: UNDERSAMPLING

$$I = \left[(aI_o + \sum_i \alpha_i X_i) e^{-\sum_j \alpha_j X_j} + \sum_k \alpha_k X_k \right] ScalPoly + BasePoly$$



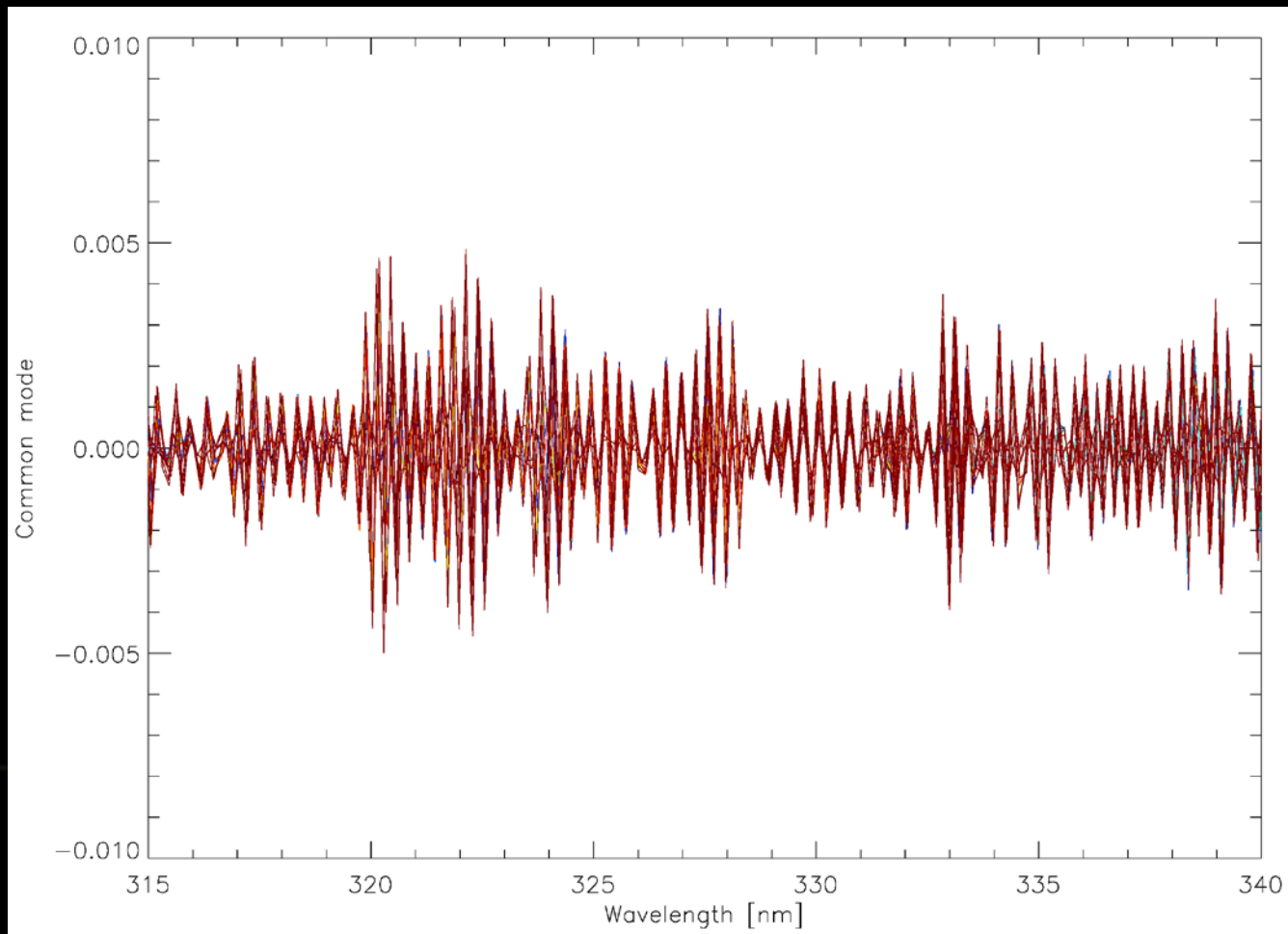
SLANT COLUMN RETRIEVAL: CROSS SECTIONS

$$I = \left[(aI_o + \sum_i \alpha_i X_i) e^{-\sum_j \alpha_j X_j} + \sum_k \alpha_k X_k \right] ScalPoly + BasePoly$$

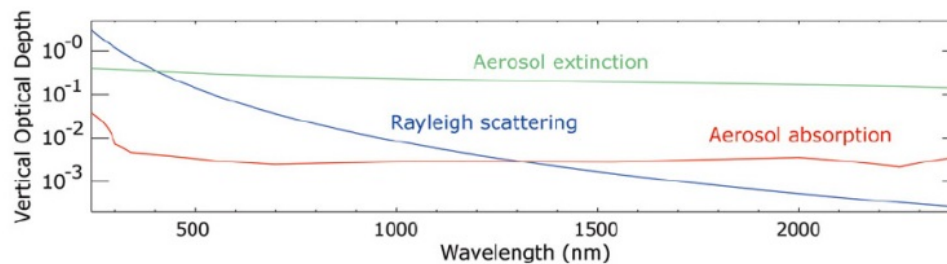
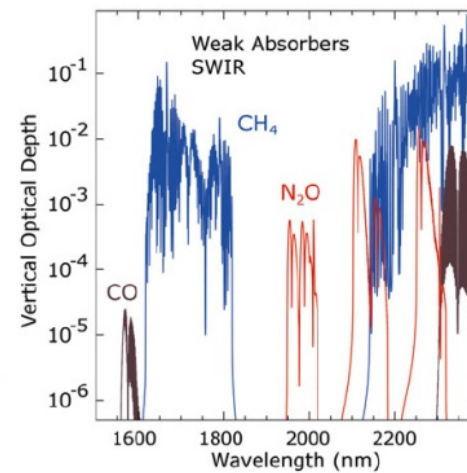
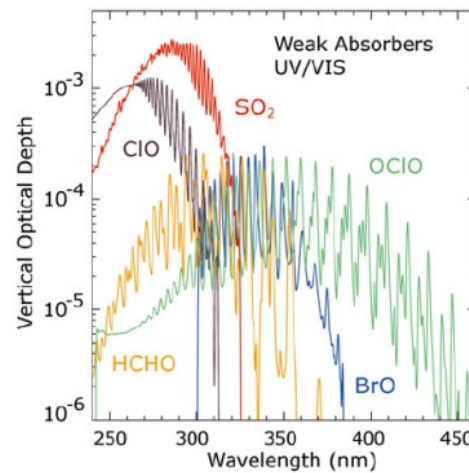
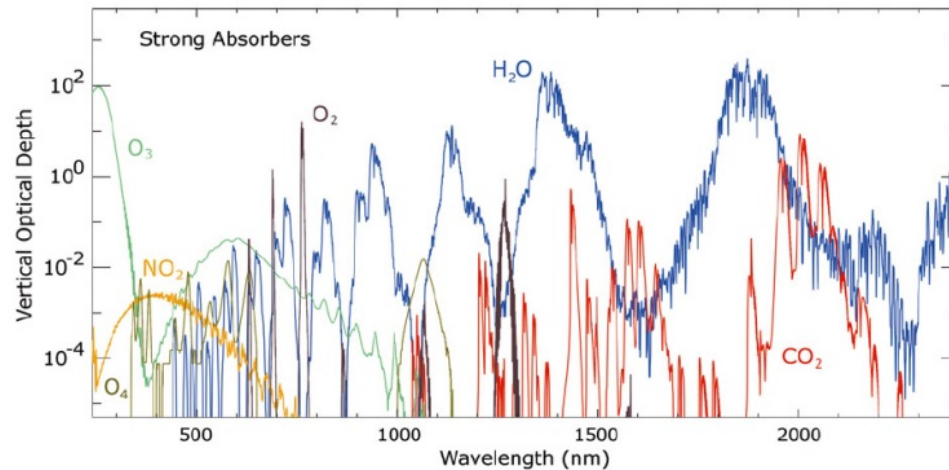


SLANT COLUMN RETRIEVAL: COMMON MODE

$$I = \left[(aI_o + \sum_i \alpha_i X_i) e^{-\sum_j \alpha_j X_j} + \sum_k \alpha_k X_k \right] ScalPoly + BasePoly$$

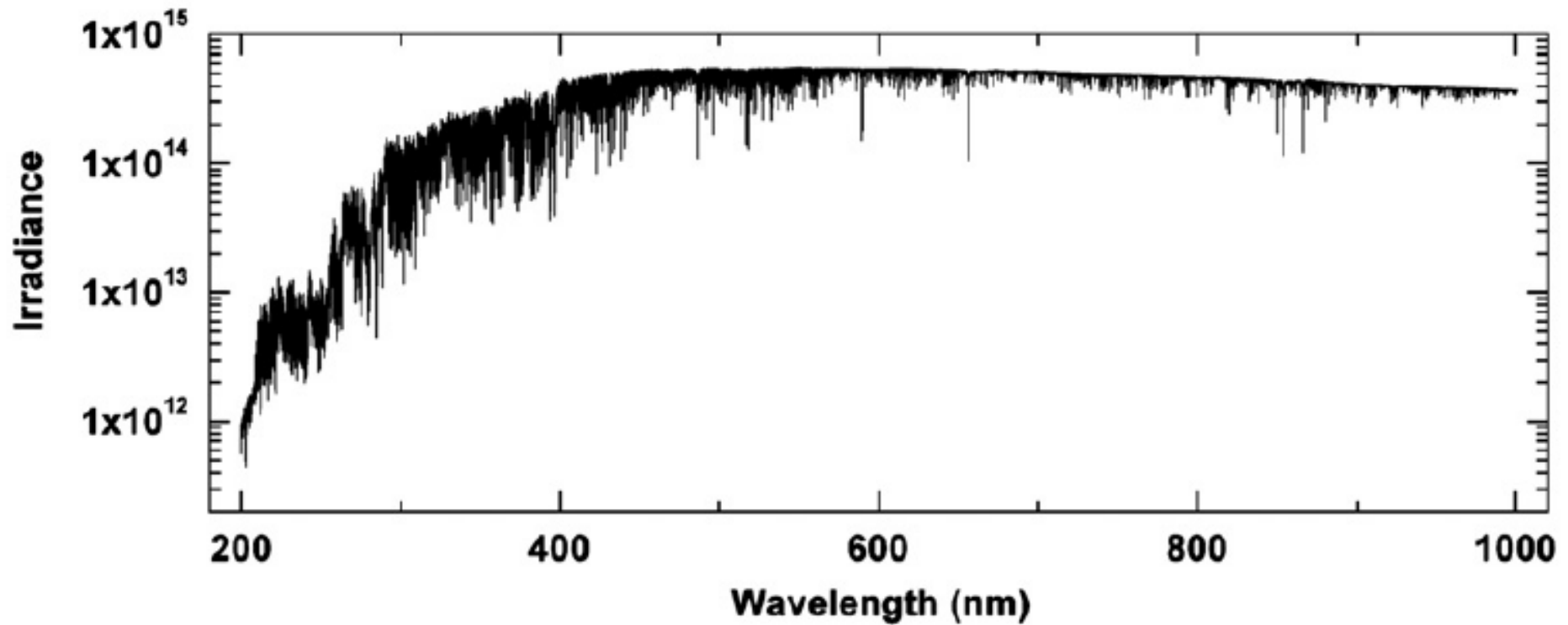


SLANT COLUMN RETRIEVAL: OPTICAL DEPTH

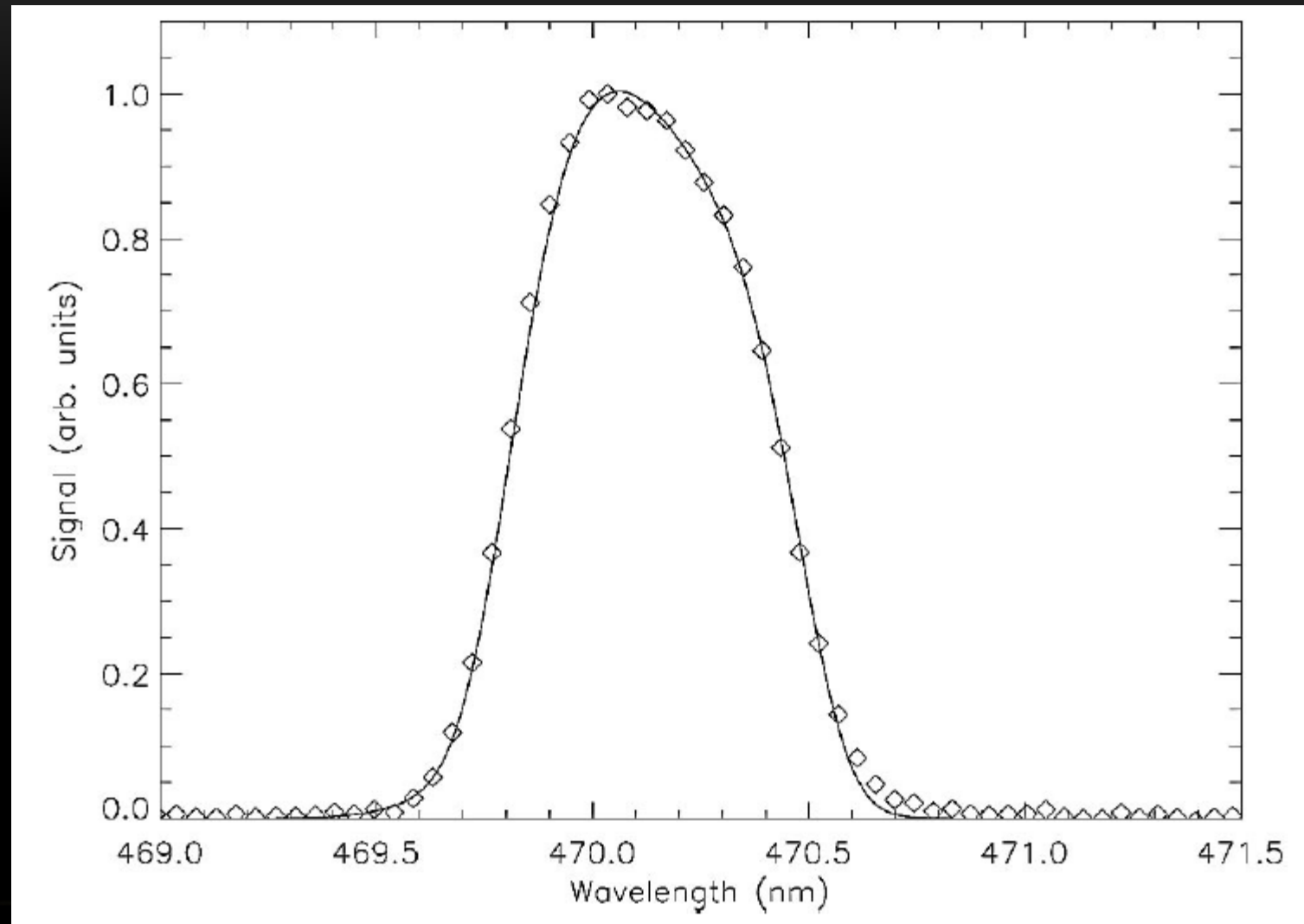


Courtesy: IUP-IFE,
University of Bremen

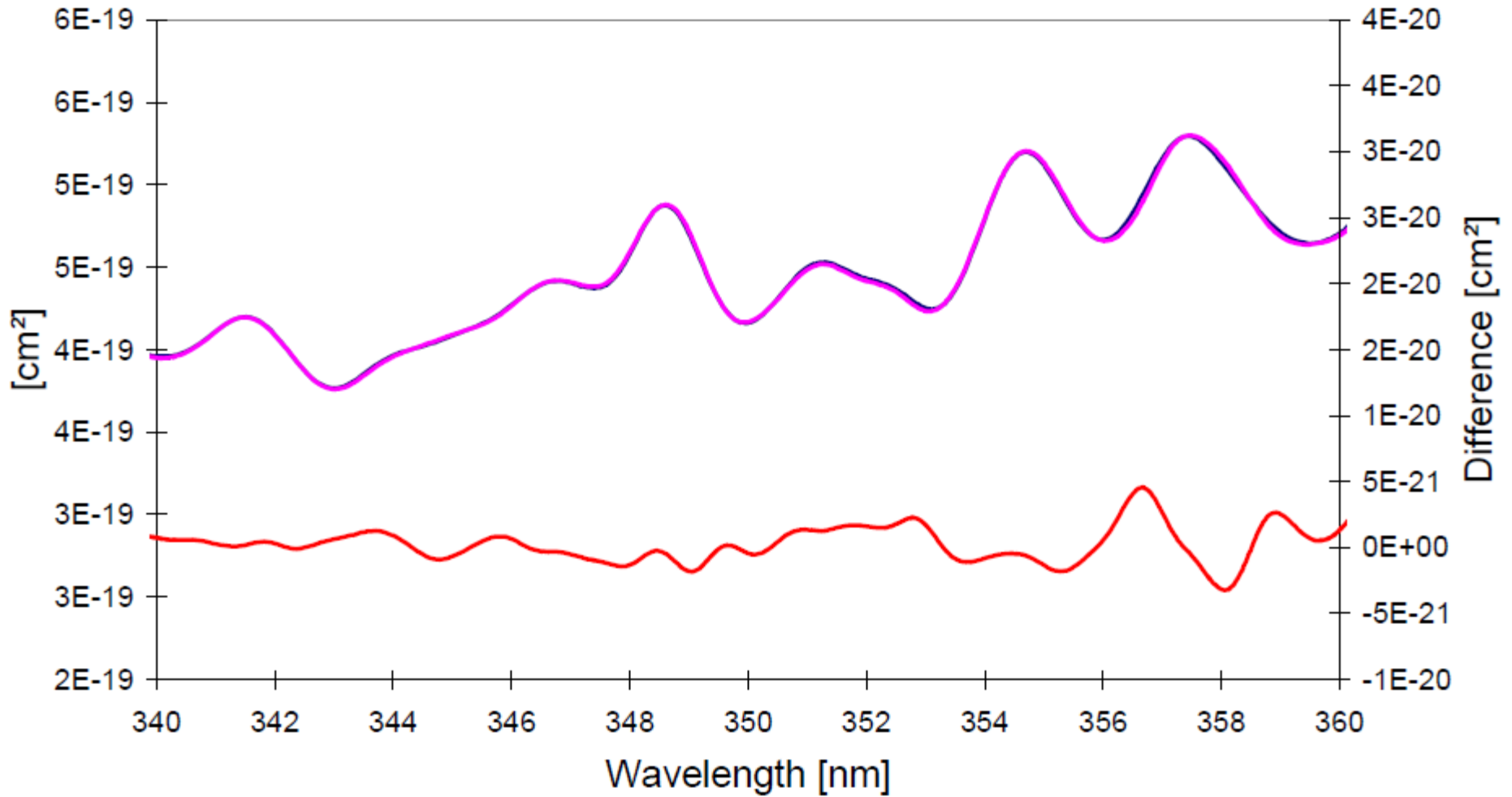
SLANT COLUMN RETRIEVAL: HIGH RESOLUTION SOLAR SPECTRUM



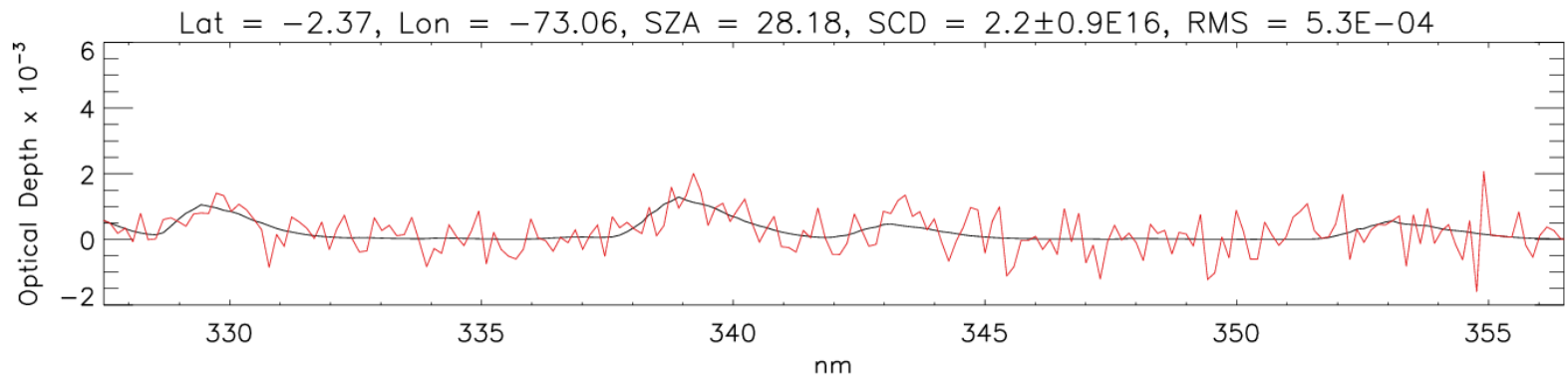
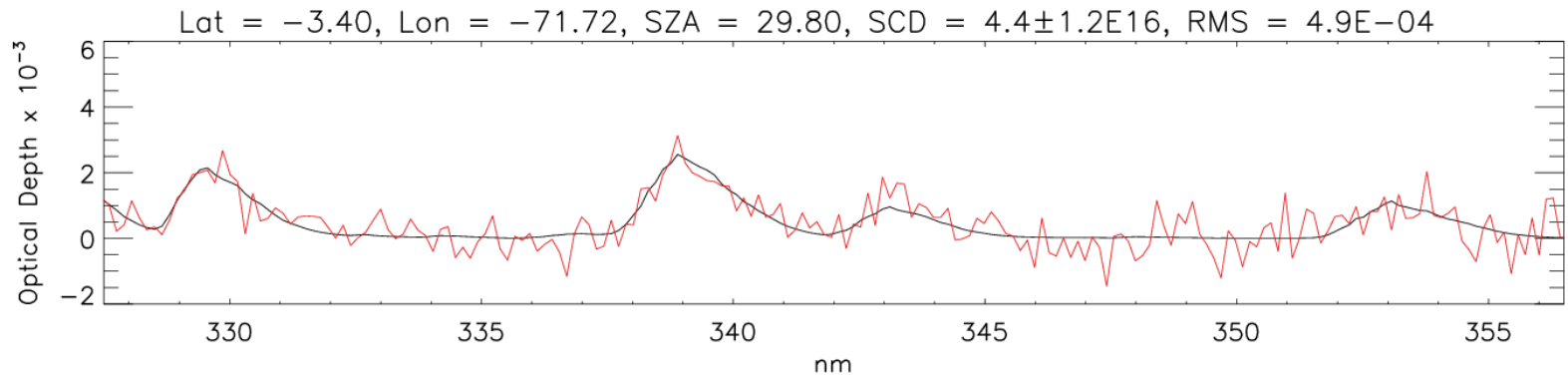
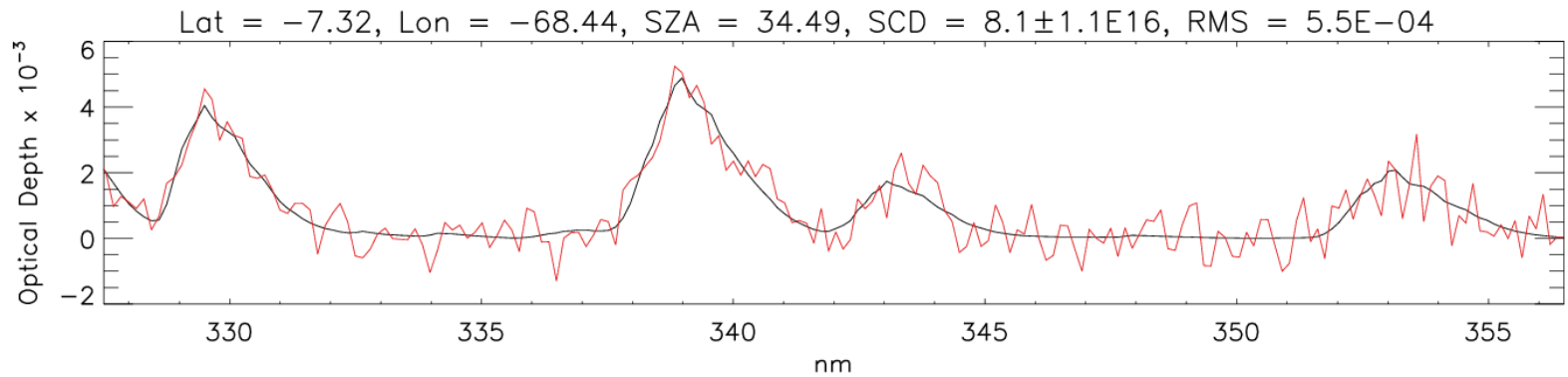
SLANT COLUMN RETRIEVAL: SLIT FUNCTION



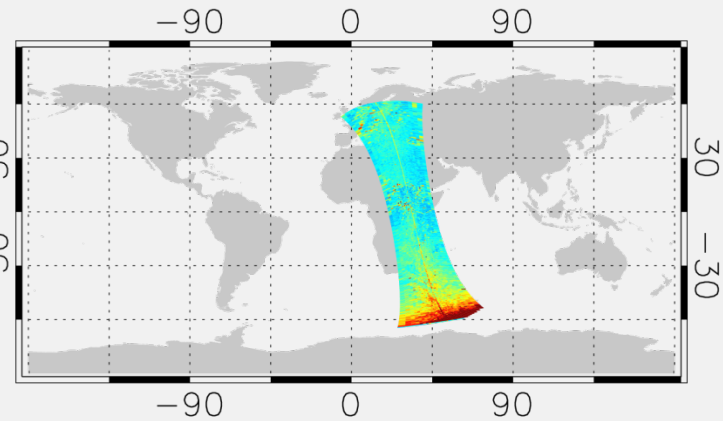
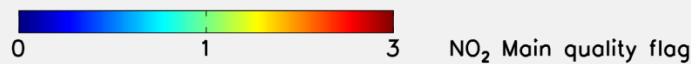
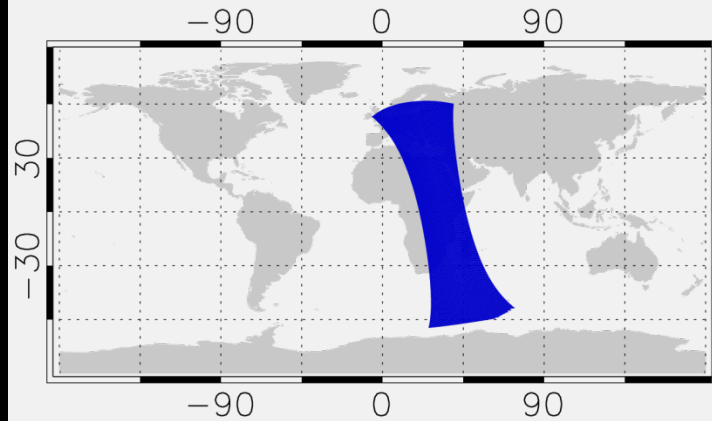
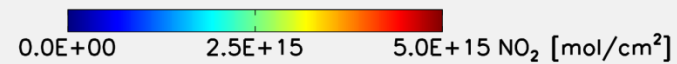
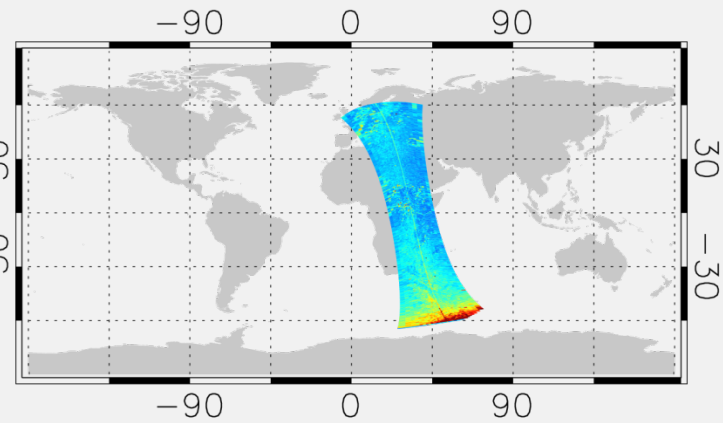
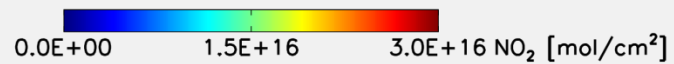
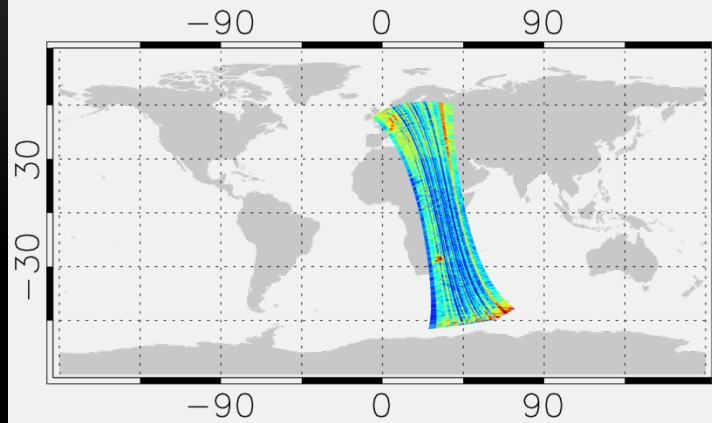
SLANT COLUMN RETRIEVAL: I_0 CORRECTION



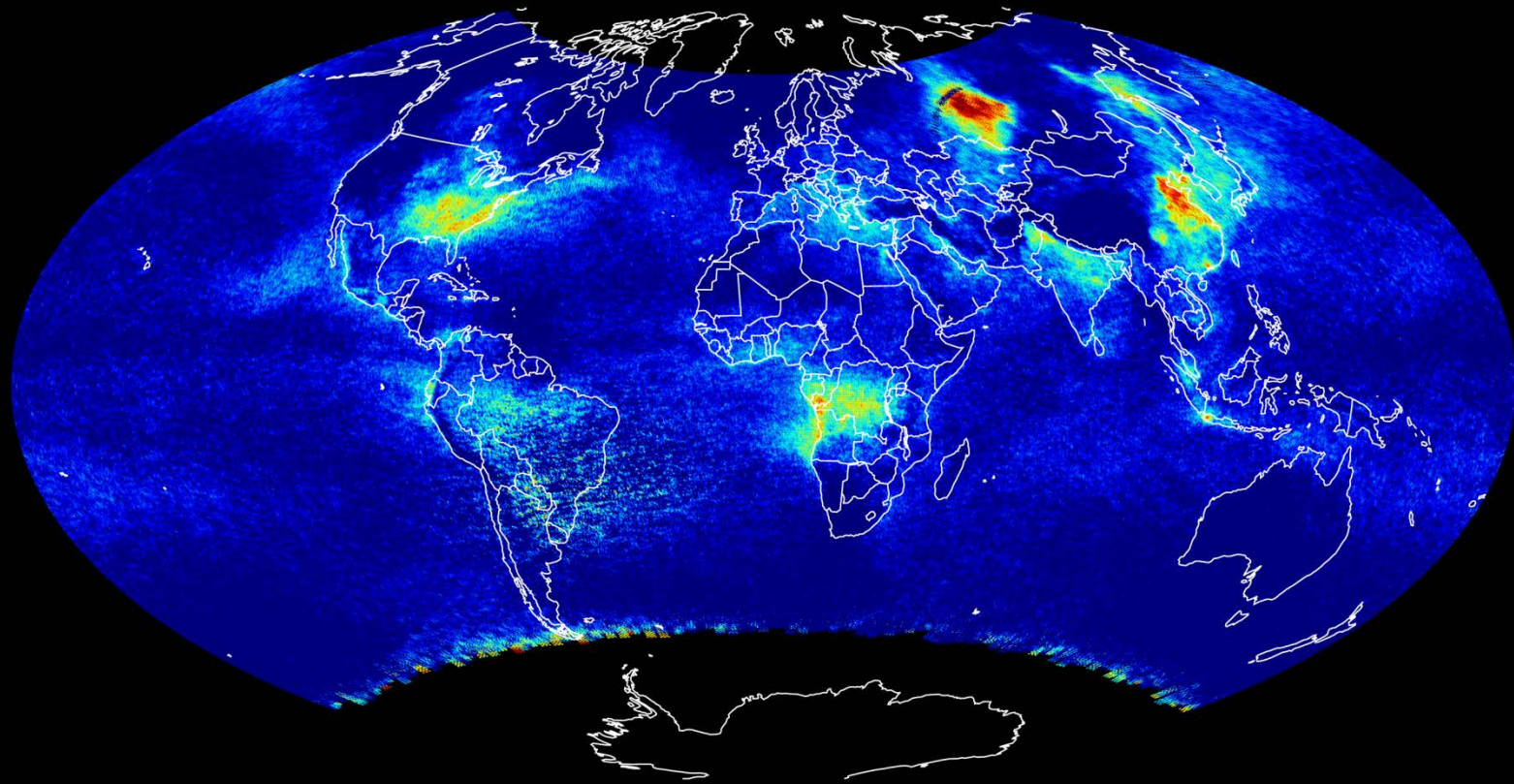
SLANT COLUMN RETRIEVAL: OMI C₂HO



SLANT COLUMN RETRIEVAL: OMPS NO₂

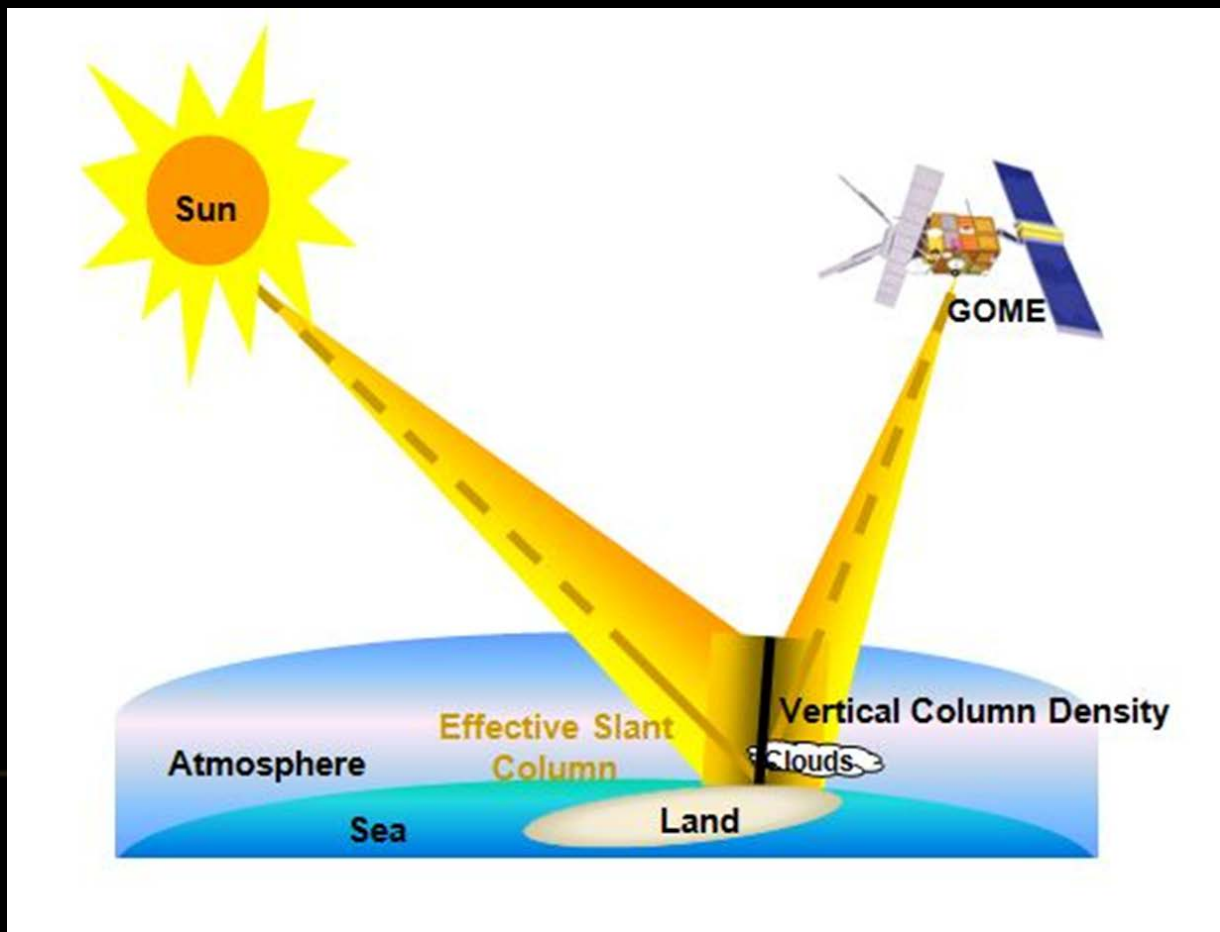


SLANT COLUMN RETRIEVAL: OMPS C₂H₄



VERTICAL COLUMN: EQUATIONS

$$VCD = \frac{SCD}{AMF} \quad \text{AMF (Air Mass Factor)}$$



VERTICAL COLUMN: EQUATIONS

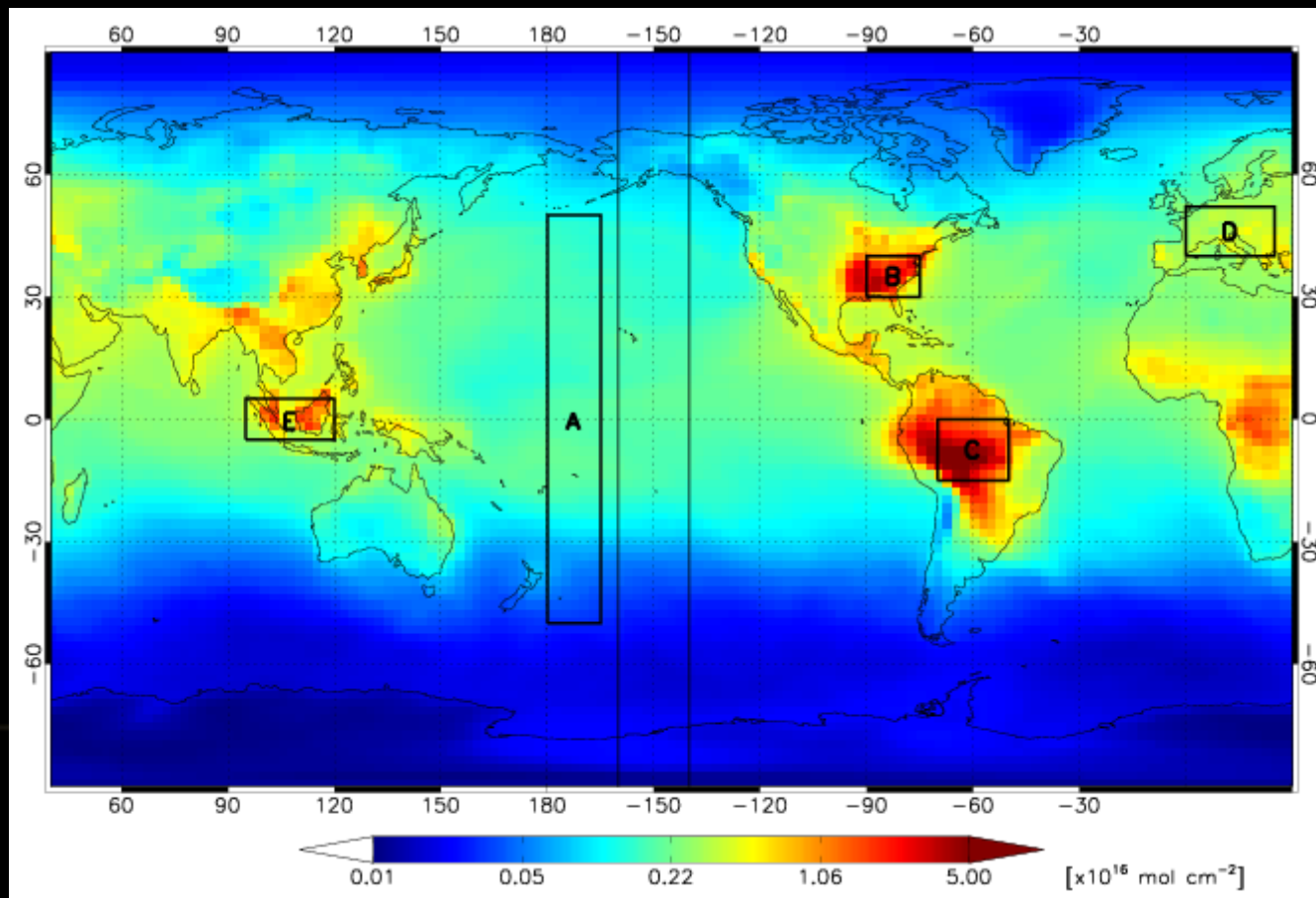
$$AMF = \int_{atm} w(z) S_z(z) dz$$

w(z) Scattering weights from Vlidor

*S_x(z) shape factors from a climatology for
example GEOS-Chem*

VERTICAL COLUMN: EQUATIONS, SHAPE FACTORS

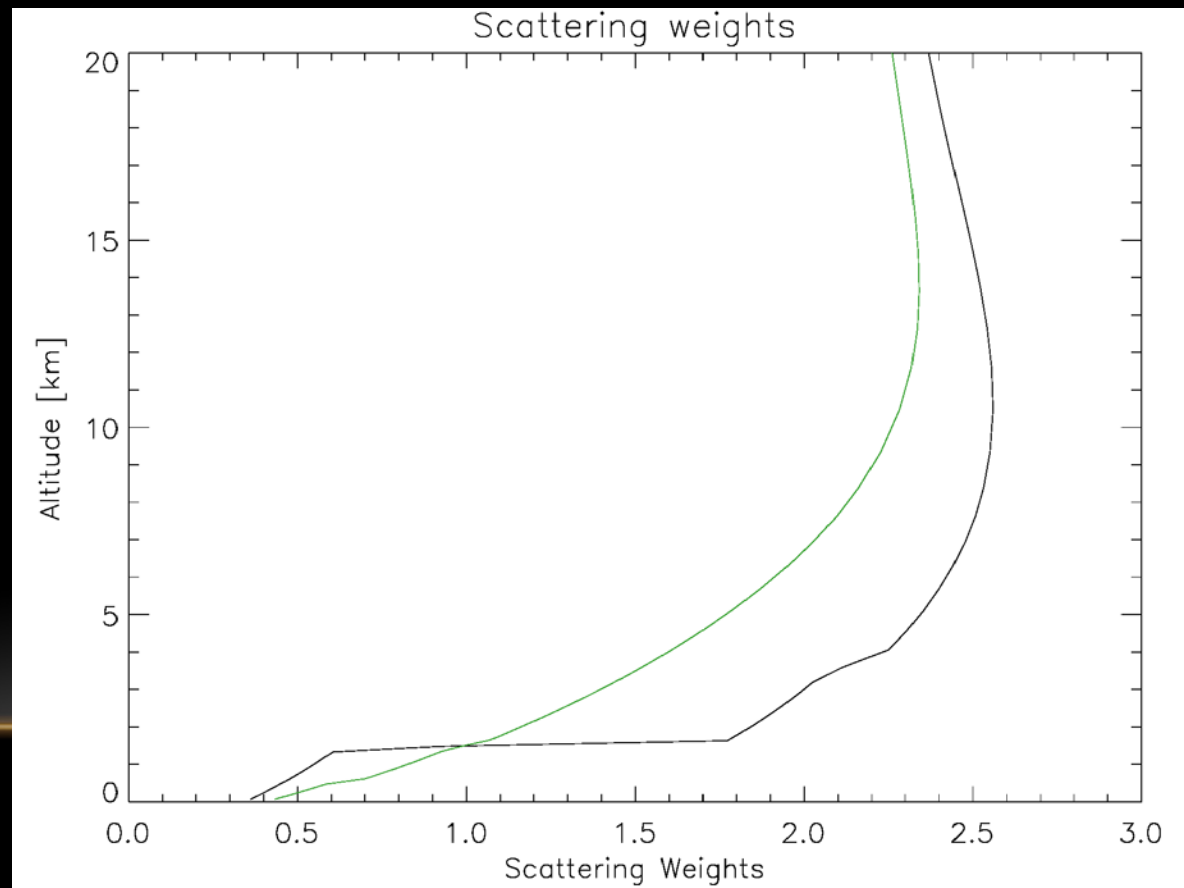
$$S(z) = \frac{x_a(z)}{\int_{atm} x_a(z) dz}$$



VERTICAL COLUMN: EQUATIONS SCATTERING WEIGHTS

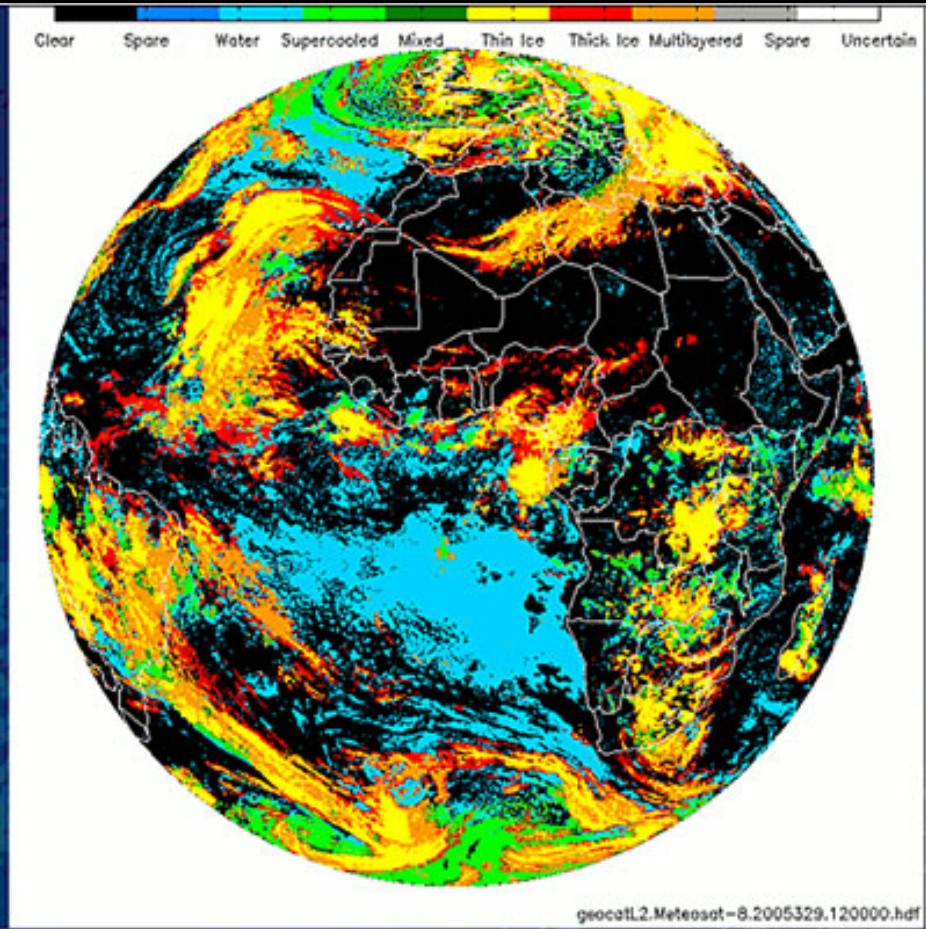
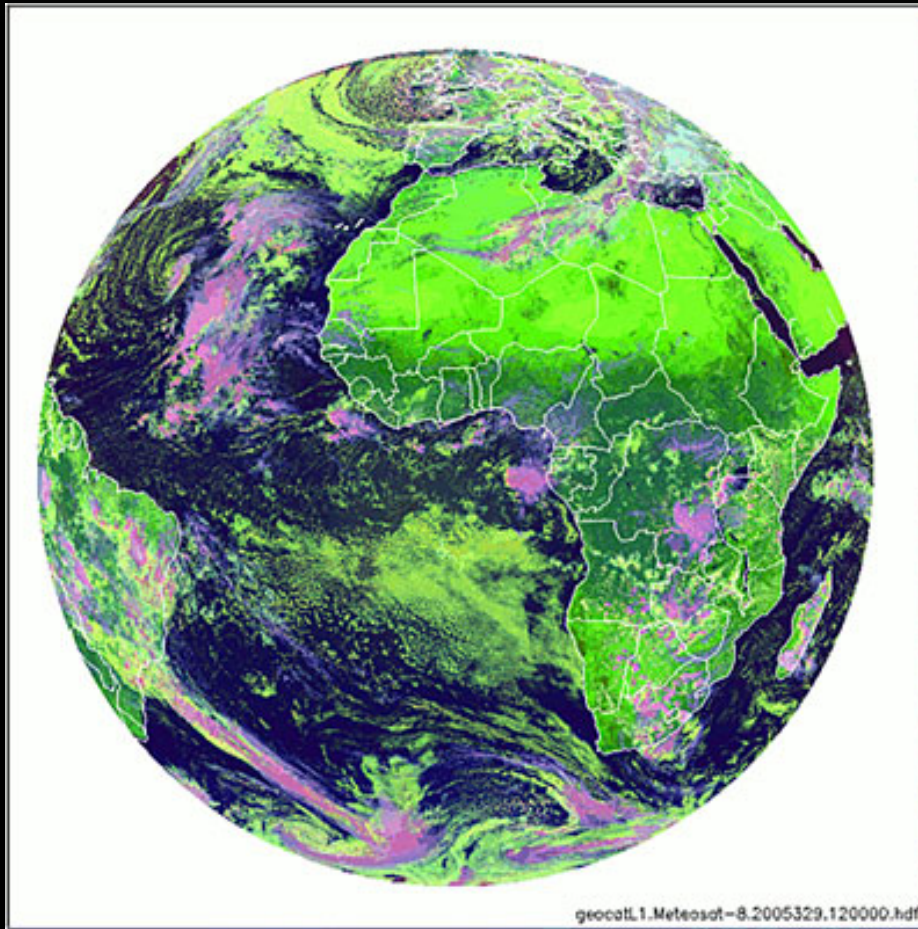
$$w(z) = \frac{\partial I / \partial z}{I \sum x_z}$$

Cloudy SZA: 28°
Clear SZA: 24°
Cloudy VZA: 16°
Clear VZA: 2.7°
Cloudy cloud fraction: 0.2
Clear cloud fractions: 0.01



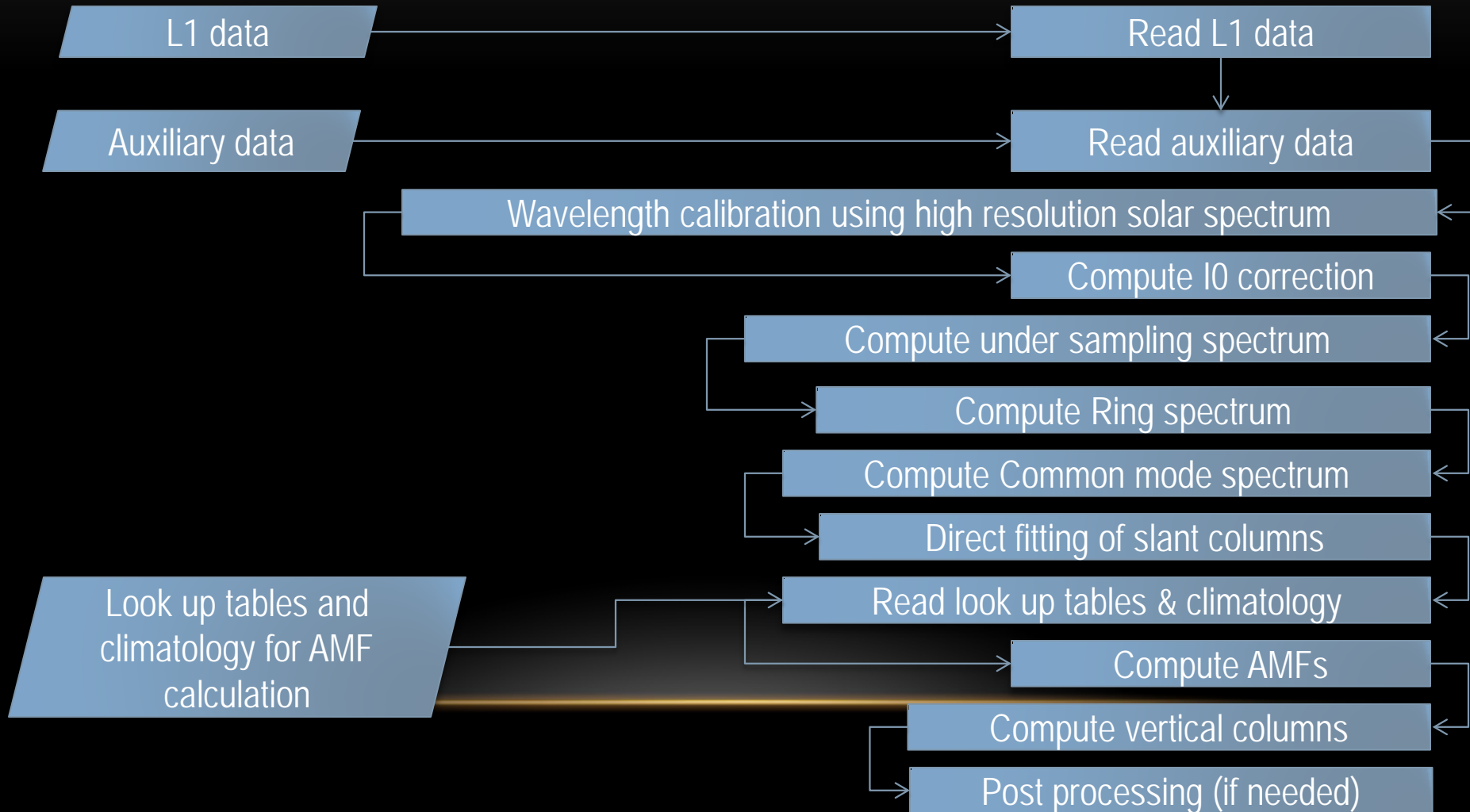
VERTICAL COLUMN: EQUATIONS SINGEL PIXEL APPROXIMATION

$$w(z) = (1 - \phi)w_{clear}(z, a_s, h_s) + \phi w_{cloud}(z, a_c, h_c)$$



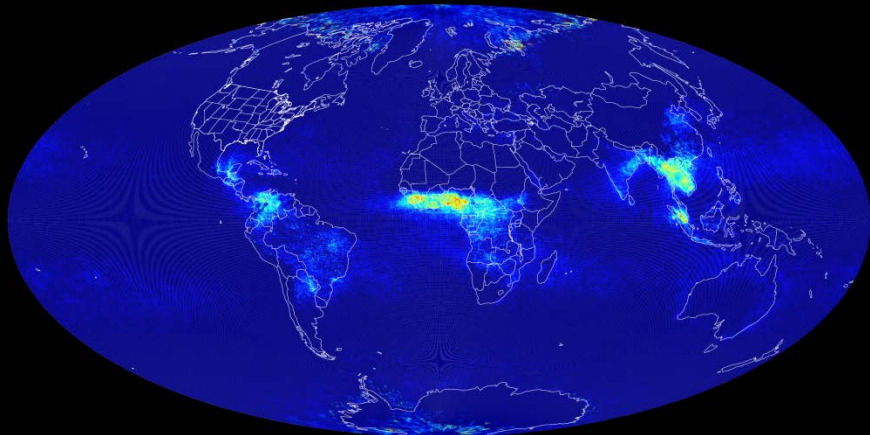
GENERIC RETRIEVAL FRAMEWORK

Typical retrieval flow chart for a retrieval

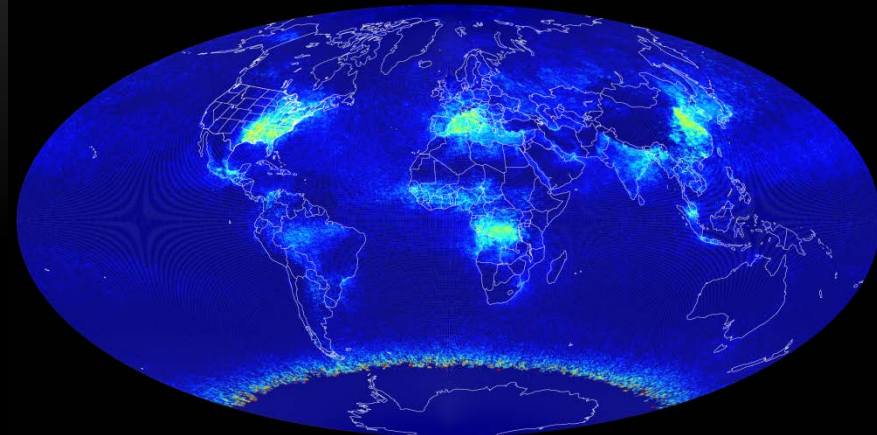


VERTICAL COLUMN RETRIEVAL: OMI C₂H₀

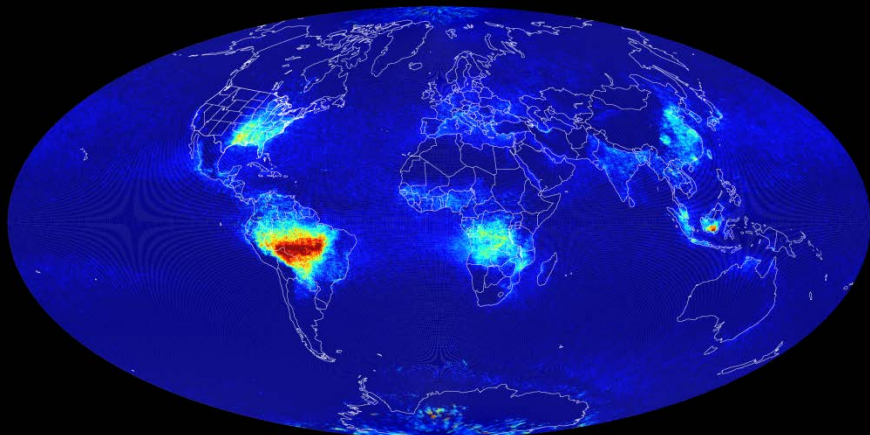
OMI HCHO 2005m03 ($\leq 40\%$ Cloud Cover)



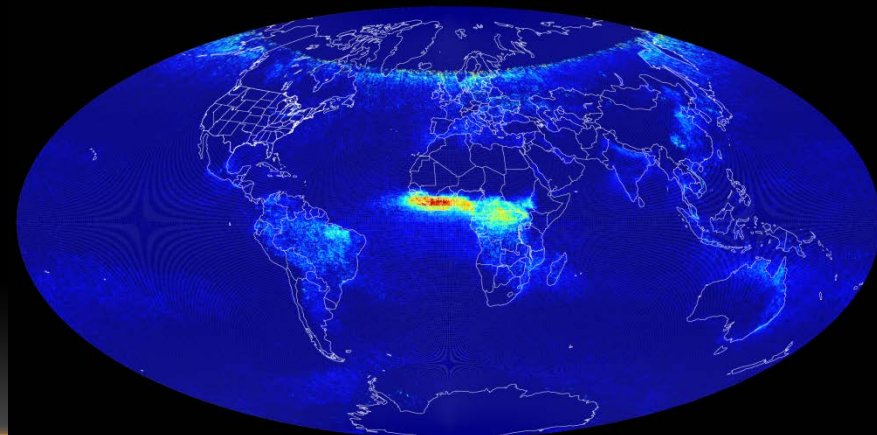
OMI HCHO 2005m06 ($\leq 40\%$ Cloud Cover)



OMI HCHO 2005m09 ($\leq 40\%$ Cloud Cover)

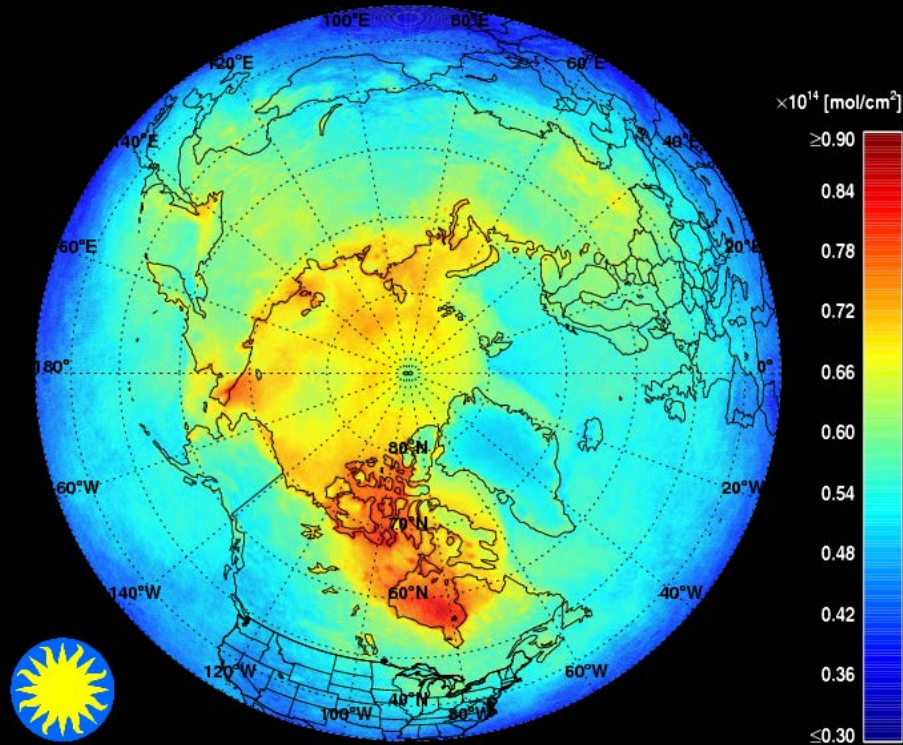


OMI HCHO 2005m12 ($\leq 40\%$ Cloud Cover)



VERTICAL COLUMN RETRIEVAL: OMI BRO

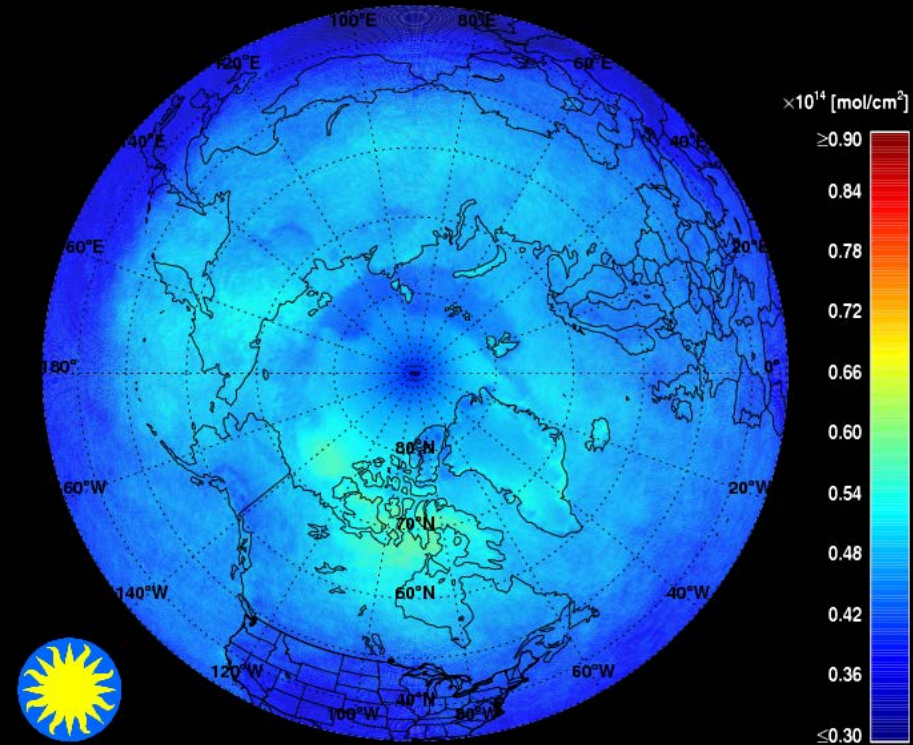
OMI BrO Vertical Columns 2005-03



Smithsonian Astrophysical Observatory

tkurosu@cfa.harvard.edu

OMI BrO Vertical Columns 2005-09

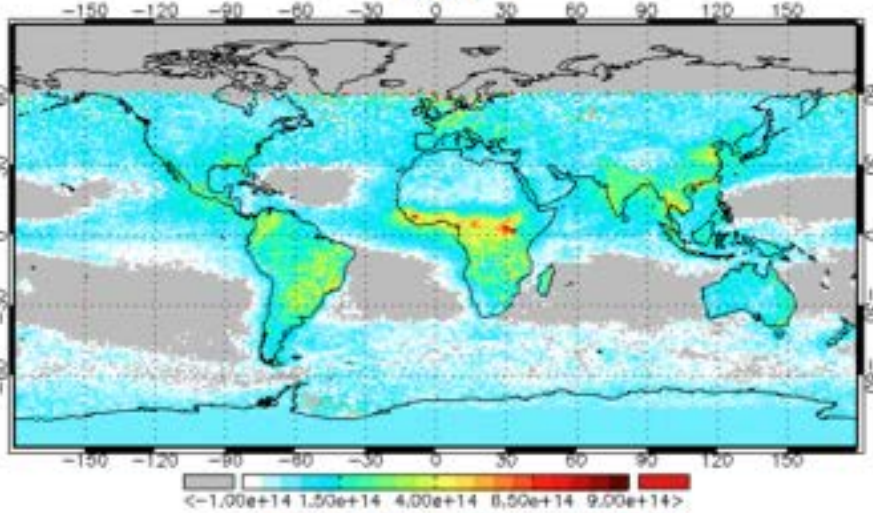


Smithsonian Astrophysical Observatory

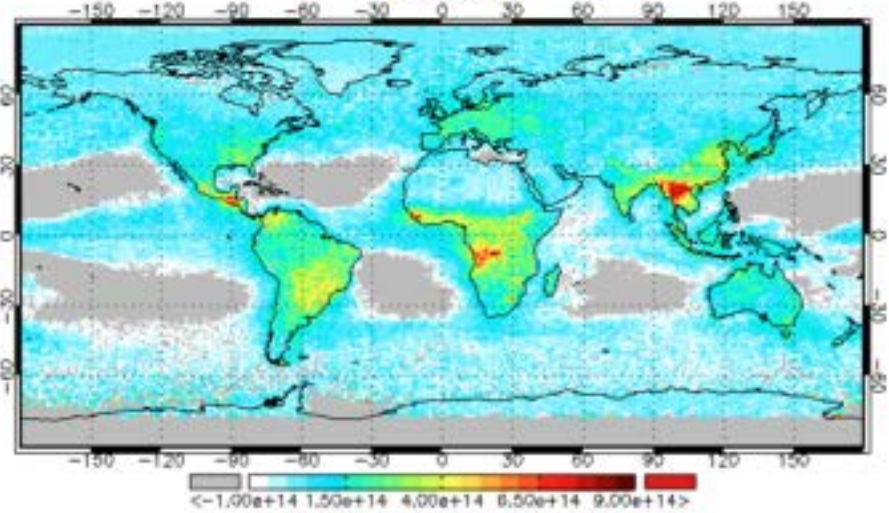
tkurosu@cfa.harvard.edu

VERTICAL COLUMN RETRIEVAL: OMI C₂H₂O₂

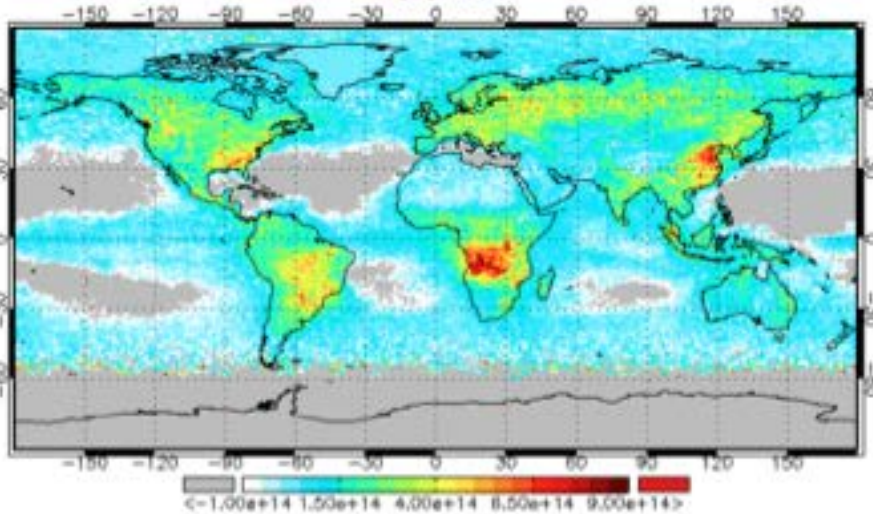
D-J-F



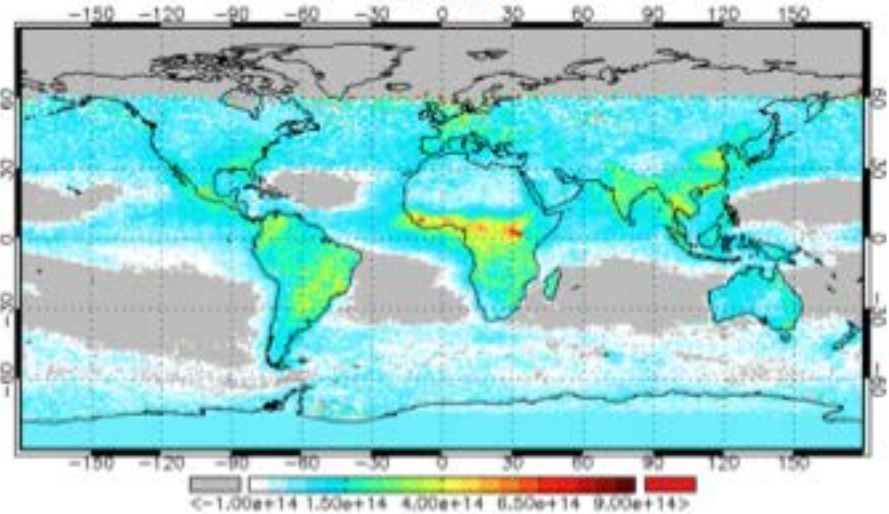
M-A-M



J-J-A

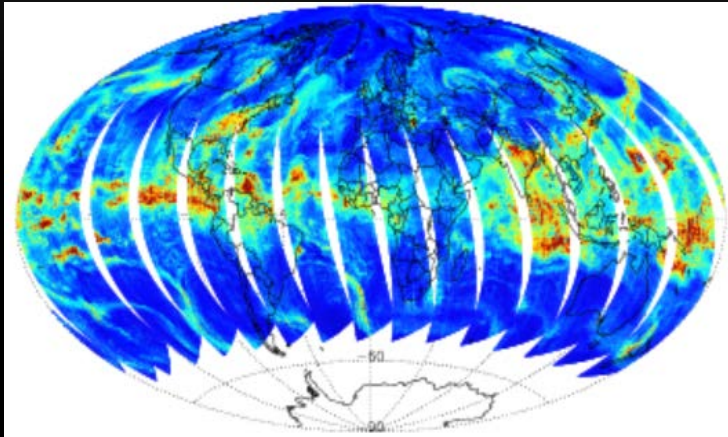


S-O-N

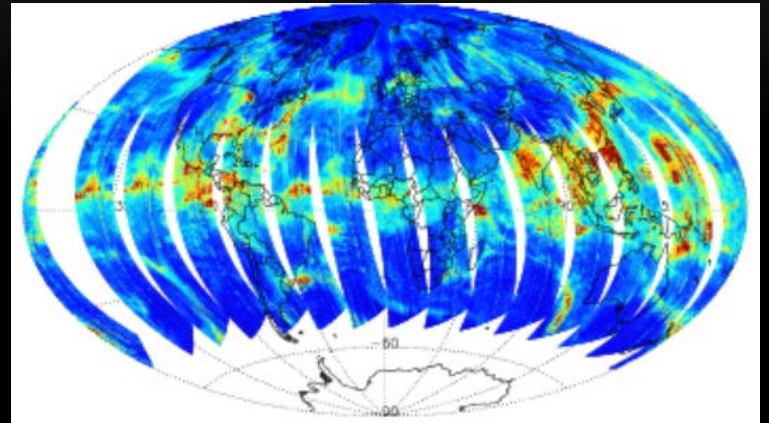


VERTICAL COLUMN RETRIEVAL: OMI H₂O

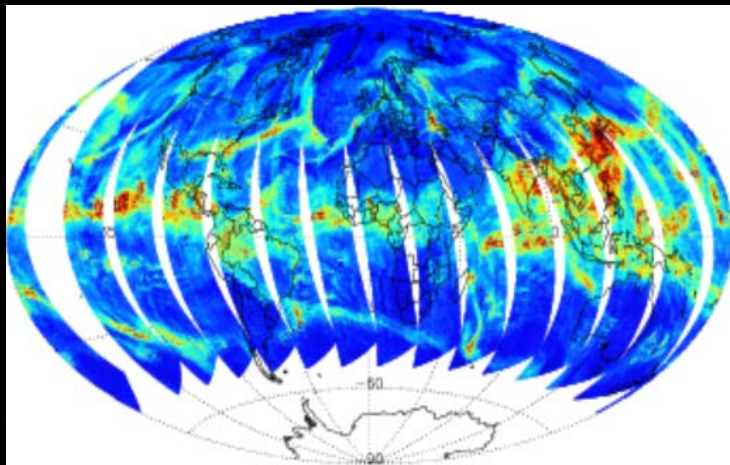
20050714



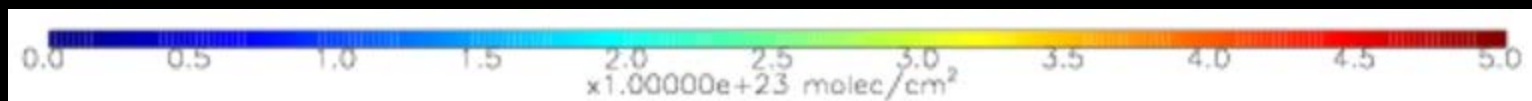
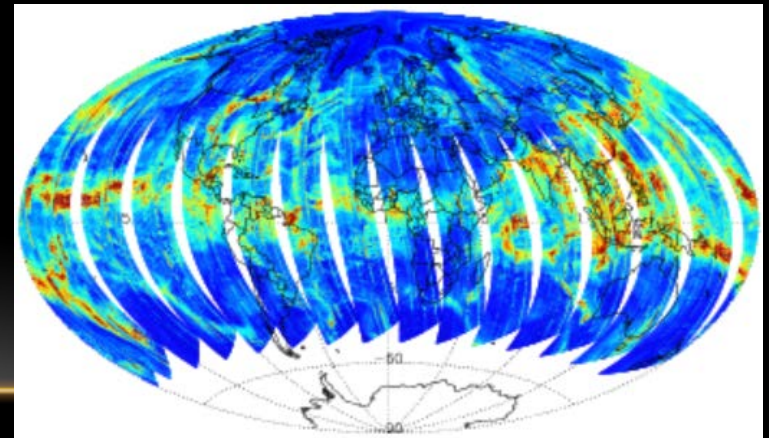
20110714



20070714



20130714



THE END



PLUS TEAM HANDBALL

OMI operational HCHO

Fitting window	327.5 – 356.5 nm
Baseline polynomial	3 rd order
Scaling polynomial	3 rd order
Instrument slit function	Hyper-parameterization of pre-flight measurements
Wavelength calibration	Spectral shift (no squeeze)
Solar Reference Spectrum	Chance and Spurr, 1997
HCHO cross sections	Cantrell et al., 1990; 300K
O ₃ cross sections	Macilet et al., 1995; 228K
NO ₂ cross sections	Vandaele et al., 1998 220K
BrO cross sections	Wilmouth et al., 1999; 228K
Molecular Ring cross sections	Chance and Spurr, 1997
Sampling correction	Computed on-line
Residual common mode spectrum	Computed on-line

OMI CHOCHO

Fitting window	431.5 – 469.5 nm
Baseline polynomial	4 th order
Scaling polynomial	3 rd order
Instrument slit function	Hyper-parameterization of pre-flight measurements
Wavelength calibration	Spectral shift (no squeeze)
Solar Reference Spectrum	Chance and Spurr, 1997
CHOCHO cross sections	Volkamer et al., 2005; 296K
O ₃ cross sections	Macilet et al., 1995; 228K
NO ₂ cross sections	Vandaele et al., 1998 220K
O ₂ -O ₂ cross sections	Hermans BISA, 2006; 294K
H ₂ O cross sections	HITRAN 2008, 280K
Liquid H ₂ O cross sections	Pope and Fry, 1997
IO cross sections	Spietz et al., 2005; 298K
Vibrational Raman Spectrum	Chance and Spurr, 1997
Molecular Ring cross sections	Chance and Spurr, 1997
Sampling correction	Computed on-line
Residual common mode spectrum	Computed on-line

OMI operational BrO

Fitting window	319.0 – 347.5 nm
Baseline polynomial	4 th order
Scaling polynomial	4 th order
Instrument slit function	Hyper-parameterization of pre-flight measurements
Wavelength calibration	Spectral shift (no squeeze)
Solar Reference Spectrum	Chance and Spurr, 1997
HCHO cross sections	Cantrell et al., 1990; 300K
O ₃ cross sections	Macilet et al., 1995; 218K; 295K
NO ₂ cross sections	Vandaele et al., 1998 220K
BrO cross sections	Wilmouth et al., 1999; 228K
O ₂ -O ₂ cross sections	Hermans BISA, 2006; 294K
SO ₂ cross sections	Hermans et al., 2009; 295K
OCIO cross sections	Kromminga et al., 2003; 213K
Molecular Ring cross sections	Chance and Spurr, 1997
Sampling correction	Computed on-line
Residual common mode spectrum	Computed on-line

OMI operational OCIO

Fitting window	358.5 – 392.0 nm
Baseline polynomial	3 rd order
Scaling polynomial	3 th order
Instrument slit function	Hyper-parameterization of pre-flight measurements
Wavelength calibration	Spectral shift (no squeeze)
Solar Reference Spectrum	Chance and Spurr, 1997
O ₃ cross sections	Macilet et al., 1995; 228K
NO ₂ cross sections	Vandaele et al., 1998 220K
BrO cross sections	Wilmouth et al., 1999; 228K
O ₂ -O ₂ cross sections	Hermans BISA, 2006; 294K
OCIO cross sections	Kromminga et al., 2003; 213K
Molecular Ring cross sections	Chance and Spurr, 1997
Sampling correction	Computed on-line
Residual common mode spectrum	Computed on-line

GOME2 SO₂

Fitting window	312.0 – 330.0 nm
Baseline polynomial	3 rd order
Scaling polynomial	3 rd order
Instrument slit function	Asymmetric Gaussian fitted online
Wavelength calibration	Spectral shift (no squeeze)
Solar Reference Spectrum	Chance and Spurr, 1997
SO ₂ cross sections	Bogumil et al., 2003; 293K
HCHO cross sections	Meller and Moortgat et al., 2000; 298K
O ₃ cross sections	Brion et al., 1993; 218K, 273K
NO ₂ cross sections	Vandaele et al., 1998 220K
BrO cross sections	Wilmouth et al., 1999; 228K
Molecular Ring cross sections	Chance and Spurr, 1997
Sampling correction	Computed on-line
Residual common mode spectrum	Computed on-line